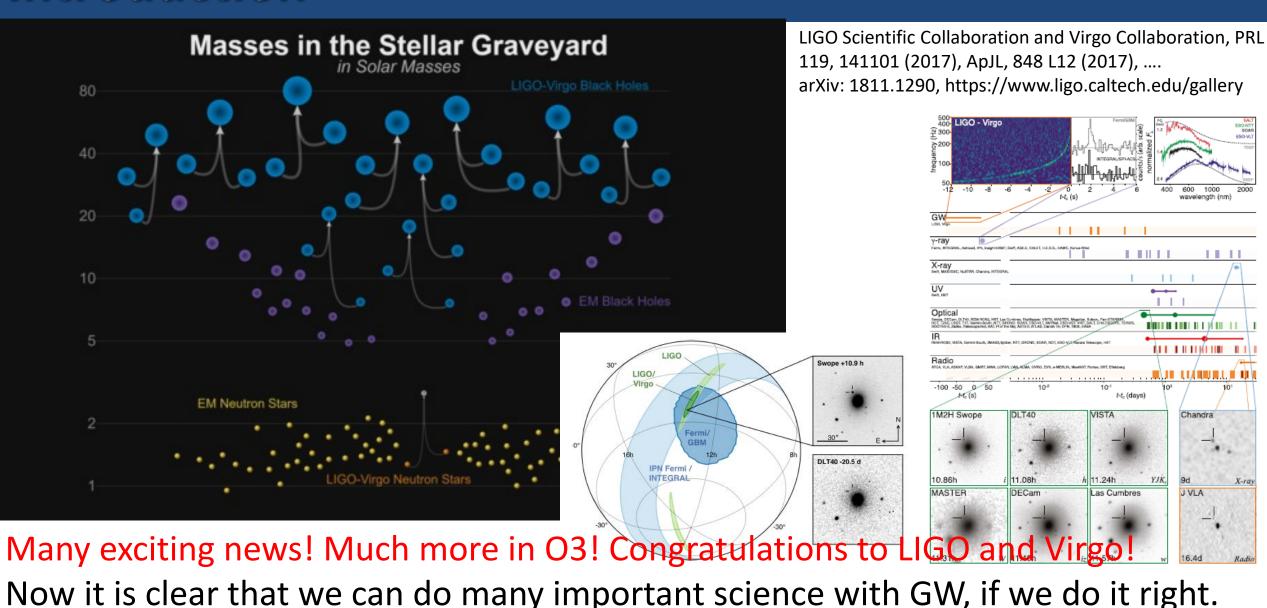


Outline

- Introduction
- Status of KAGRA
- KAGRA's contribution to the GW science
- Future plan of KAGRA
- Summary

Introduction

Introduction



4



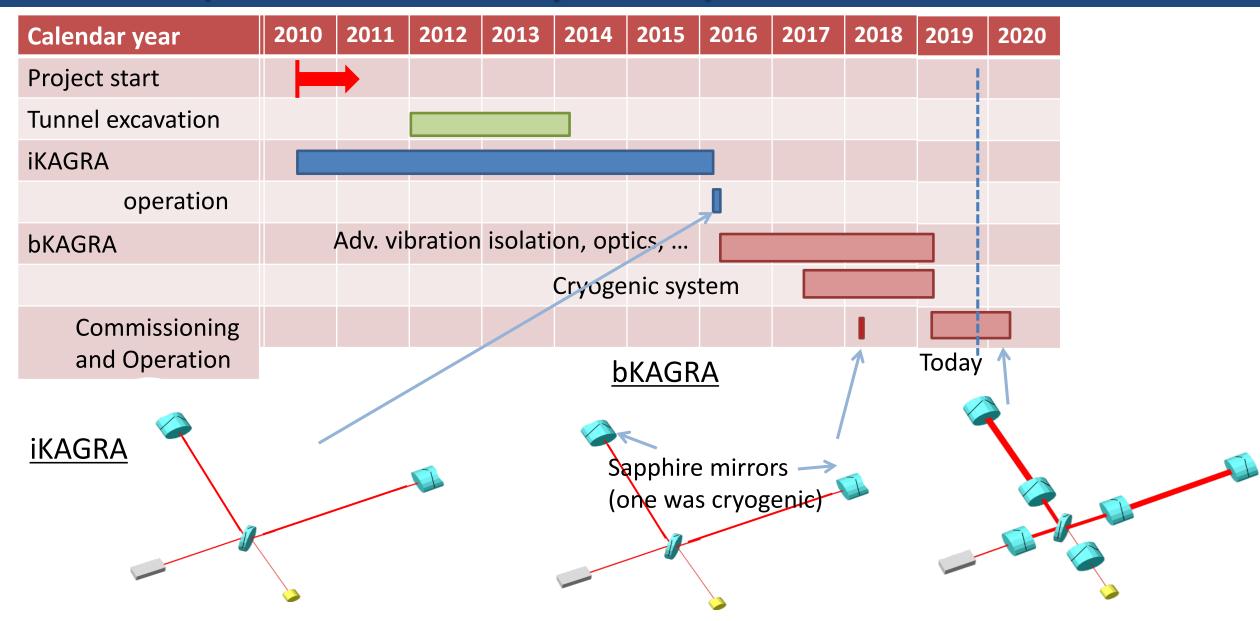
Status of KAGRA

KAGRA collaboration

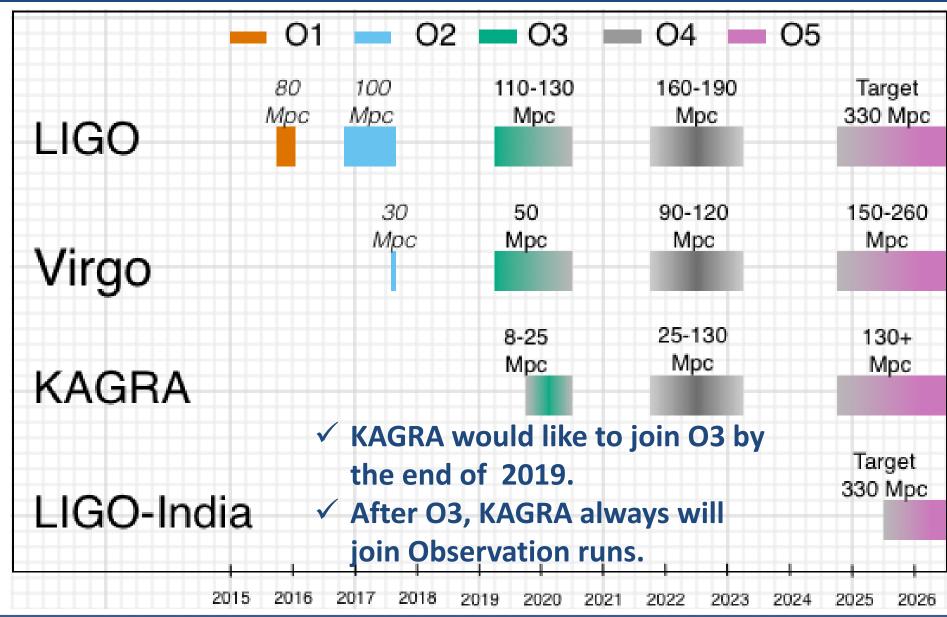


10 countries, >300 members

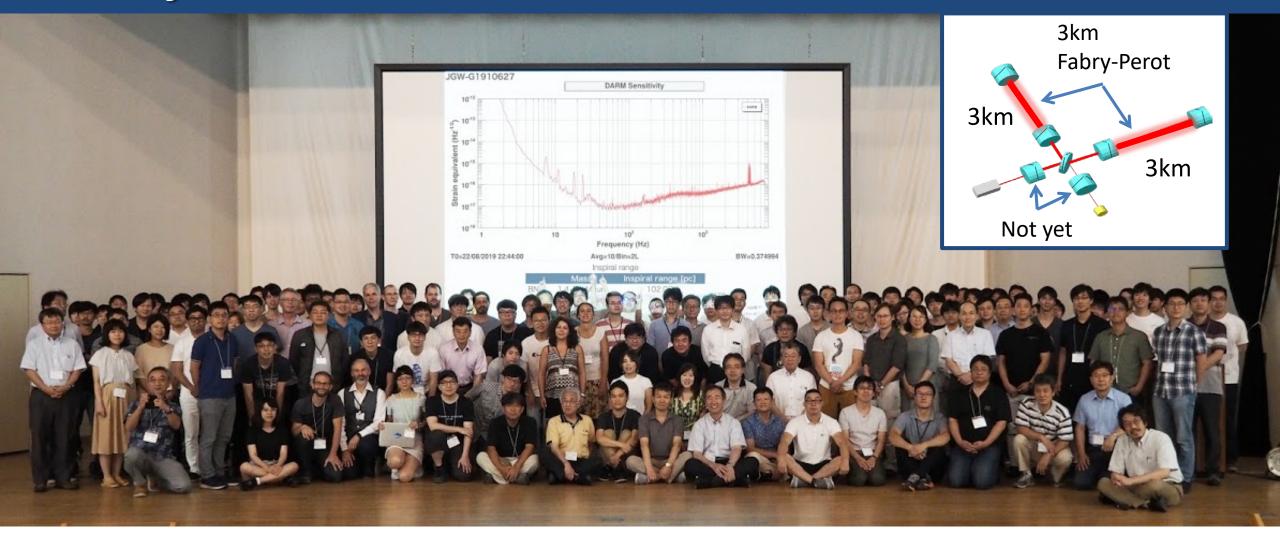
Timeline (Construction and Operation)



"Observation scenario paper"



Status of KAGRA



On Aug. 23, 2019, KAGRA has successfully operated the interferometer with the 3km Fabry-Perot cavities. (Photo during the KAGRA Face-to-Face meeting on the same day.)

Signing Memorandum of Agreement with LIGO and Virgo

M1900145-v2, VIR-0091A, and JGW-M1910663

Memorandum of Agreement

between

VIRGO,

KAGRA,

and the

Laser Interferometer Gravitational Wave Observatory (LIGO)

October 2019

Purpose of agreement:

The purpose of this Memorandum of Agreement (MOA) is to establish and define a collaborative relationship between VIRGO, KAGRA and the Laser Interferometer Gravitational Wave Observatory (LIGO) to develop and exploit laser interferometry to measure and study gravitational waves.

We enter into this agreement in order to lay the groundwork for decades of world-wide collaboration. We intend to carry out the search for and analysis of gravitational waves in a spirit of teamwork, not competition. Furthermore, we remain open to participation of new partners, whenever additional data can add scientific value to the detection and study of gravitational waves. All partners in the world-wide collaboration should have a fair share in the scientific governance of the collaborative work.

Among the scientific benefits we hope to achieve from this collaboration are: better confidence in detection of signals, better duty cycle and sky coverage for searches, better estimation of the location and physical parameters of the sources, and gravitational wave studies based on the detected signals. Furthermore, we believe that the sharing of ideas will also offer additional benefits.

This MOA supersedes the MOU LIGO-M060038-v5 between VIRGO and LIGO, established in March 2019. This MOA also supersedes the MOU JGW-M1201315-v3 between KAGRA, LSC

Approved: David Reitze Date LIGO Executive Director and LIGO Principal Investigator Albert Lazzarini Date LIGO Laboratory Deputy Director Patrick Brady Date LSC Spokesperson James Hough Date GEO Representative 30-Sep-2019 Stavros Katsanevas Date Director of EGO Jo van den Brand Date Virgo Collaboration Spokesperson Albert Lazzarini Date 30-Sep-2019 Stavros Katsanevas Date Director of EGO Jo van den Brand Date Virgo Collaboration Spokesperson Albert Lazzarini Date Sign Sign Sign		
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Patrick Brady LSC Spokesperson James Hough GEO Representative 30-Sep-2019 Stavros Katsanevas Date Director of EGO Jo van den Brand Virgo Collaboration Spokesperson May Date KAGRA Principal Investigator H. Shinlan Ox 4, 2019 Sign		
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H. Shinlan Oa 4, 2019 Sia		
HISAAKI Shinkai Date	H. Shinlan Oct 4, 2019 Hisaaki Shinkai Date	Sig

KSC Board Chair



Sign-up ceremony on Oct. 4, 2019

KAGRA's contribution to the GW science

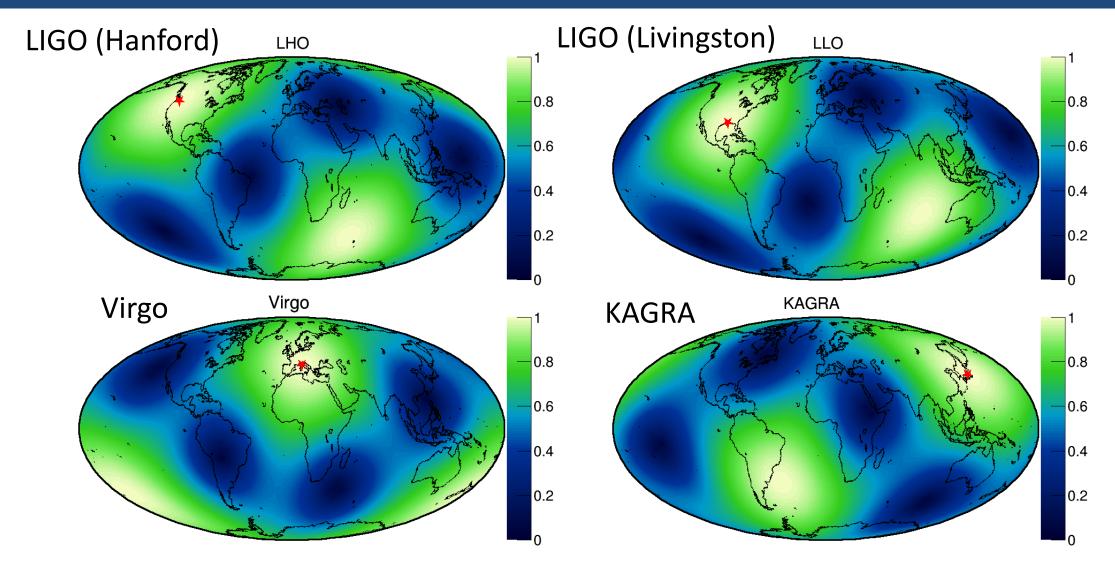
Importance of multiple antennas

◆ Let us require at least 3 detector operation for the determination of the source direction.

Duty cycle of a single detector	70%	80%
3 detectors (LHO, LLO, Virgo)	34%	51%
4 detectors (LHO, LLO, Virgo, KAGRA)	65%	81%
5 detectors (LHO, LLO, Virgo, KAGRA, LIGO-India)	83%	94%

Adding KAGRA (and LIGO-India) has a significant impact on the 3 detector coincidence!

Importance of Global GW Network: Detector antenna patterns



KAGRA is complementary in the sensitive direction to other detectors.

Importance of Global GW Network: Sky localization

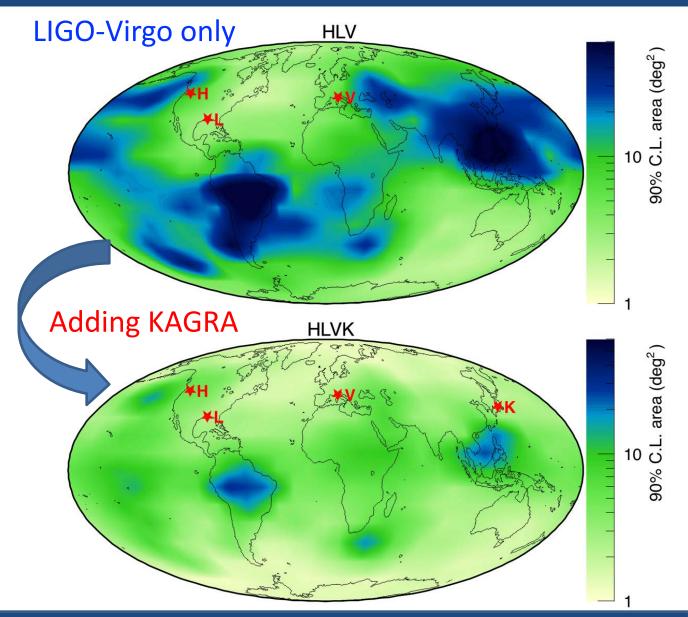
Assuming the sensitivity of;

LIGO	Virgo	KAGRA
205 Mpc	126 Mpc	152 Mpc

LV: LIGO-P1200087, K: JGW-T1707038

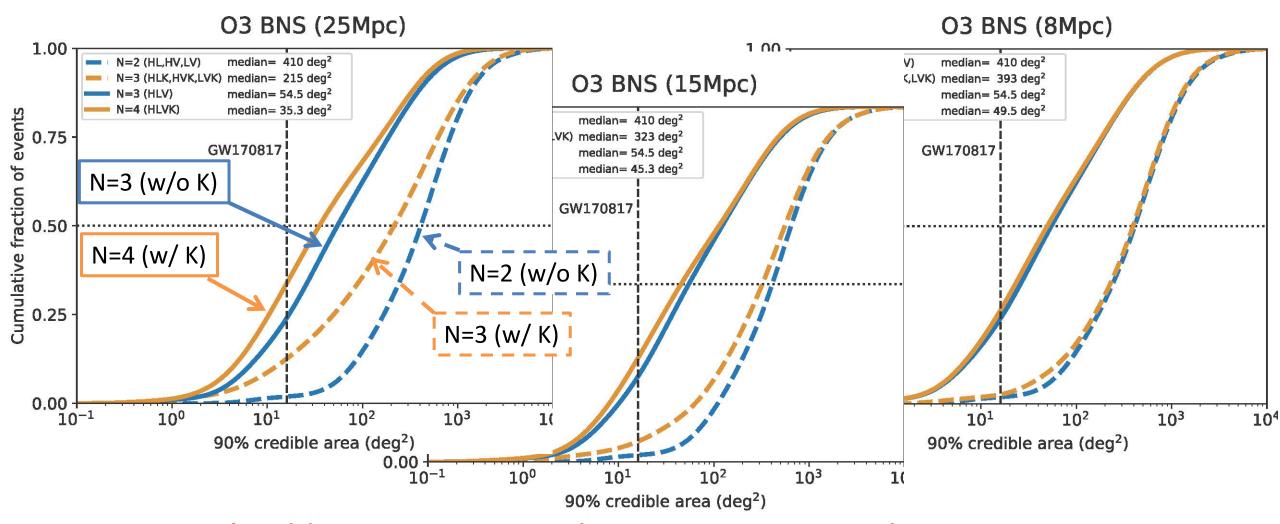
Also, assuming NS-NS merger
 (1.4 M_{Sun} -1.4 M_{Sun}) at 150 Mpc

However, the expected sensitivity of KAGRA (at least) during O3 is much lower than the design sensitivity...



Source localization during O3

Assumed sensitivity (Binary NS range): LIGO 120Mpc, Virgo 60Mpc, KAGRA 8, 15 and 25Mpc



KAGRA should try to maximize the sensitivity as much as we can....

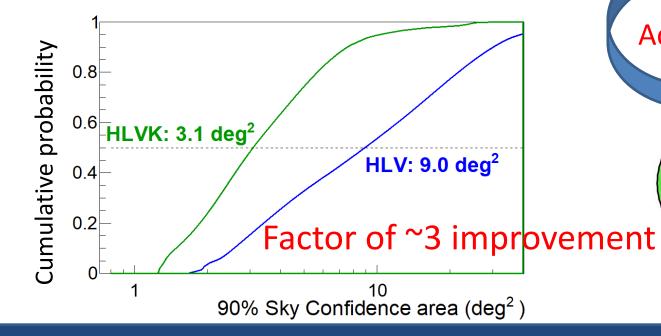
Importance of Global GW Network: Sky localization

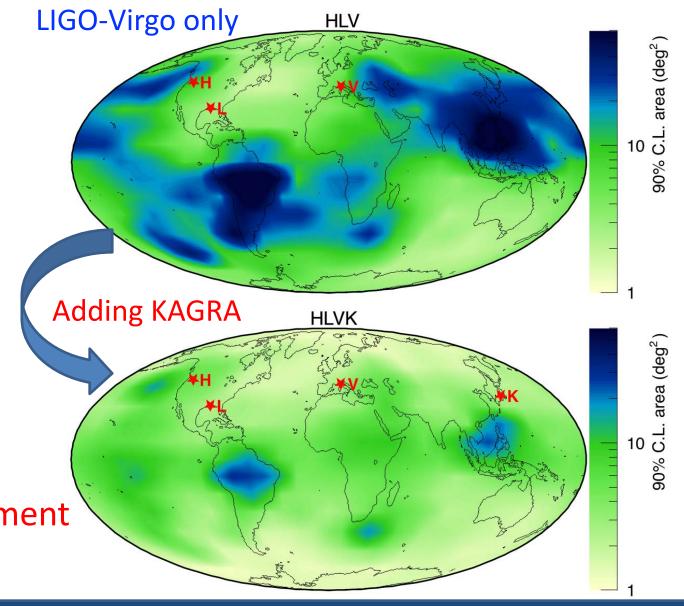
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Future plan of KAGRA

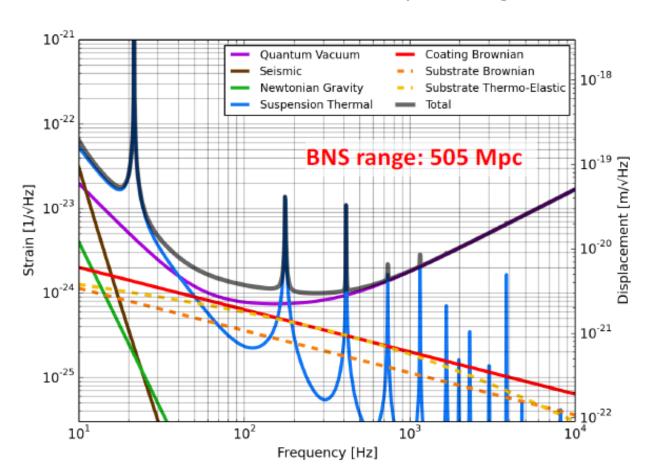
After 03

- After finishing the O3 run at the end of April 2020, KAGRA will soon begin the works to achieve a higher sensitivity by improving the present instruments (and software).
- The target sensitivity of KAGRA for O4 (fall 2021 to spring 2023) is between 25 and 130 Mpc for BNS mergers.
- In O5 (from fall 2024?), KAGRA would like to achieve > 130 Mpc for BNS mergers. Between O4 and O5, KAGRA may install something new, such as much better Sapphire Input Test Masses (ITMs) and squeezed light in order to get a high sensitivity.
- After O5, KAGRA would like to improve the sensitivity further (and substantially). Various options are discussed in the KAGRA Future Planning Committee. (Target sensitivity around 500 Mpc for BNS mergers?)

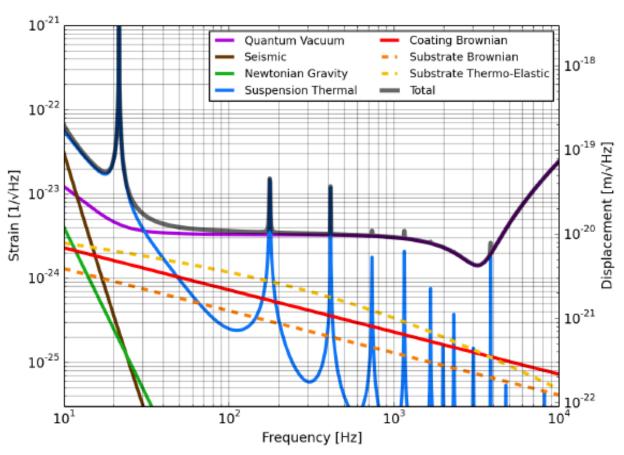
Possible future sensitivities of KAGRA

S. Haino (FPC) Aug. 2019 KAGRA F2F meeting

Assuming: 200kg, 10ppm/cm sapphire,
 500W laser, and 10dB FD squeezing



High frequency option with 500m SRC



Summary

- KAGRA plans to join O3 in late 2019. We have signed a MoA with LIGO and Virgo.
- KAGRA has began the discussions to improve the sensitivity in the future.
- KAGRA would like to contribute to the global network of gravitational wave detectors and to the science of gravitational wave astronomy.