Status of the frequency dependent squeezing experiment at TAMA

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Squeezing and gravitational wave detector

Quantum noise has already been the limiting noise for gravitational wave detector especially at high frequency. Squeezing is proven to be effective for reducing quantum noise and has already been used for O3.









Goal of experiment

Full scale filter cavity prototype to demonstrate frequency dependent squeezing with rotation at ~70Hz. . Cavity length: 300m 2. Finesse: 4400 @1064nm 3. Storage time 2.8ms @1064nm





Frequency independent squeezing:

Now frequency independent squeezing is used for O3. LLO and Virgo sensitivity with and without squeezing are shown below.



Low frequency quantum noise is due to radiation pressure noise and can be mitigated by amplitude squeezing.

High frequency quantum noise is due to shot noise and can be mitigated by phase squeezing.

Frequency dependent squeezing:

It is planned for O4 in LIGO and Virgo and is one of the candidate future upgrade plan for KAGRA. The optimal length should be hundred meter scale.





Squeezing degradation budget. Quantum noise relative to coherent vacuum in the signal quadrature for an ideal system (blue curve) is compared that with squeezing, taking into account degradation mechanisms (one by one).



Improvement in KAGRA sensitivity using 9 dB frequency-dependent squeezing, considering lossy cavity and other degradation mechanisms.

Status

•Filter cavity installation, control and characterization has been achieved •Squeezed vacuum source installation has been finished 4.4dB of squeezing and 15dB of anti-squeezing has been measured

•Commissioning of the squeezed vacuum source is on-going •KAGRA-like DGS system implemented •Automatic alignment system design is finished

> ____ green power = 10,20,30,40,49 mW — loss = 30 %, phase noise = 300 mrad

On-bench cavity parameters and control servo:

SHG and OPO have design Finesse of 70. Mode cleaners have design Finesse of 400.

We use seven different analog servos to control cavities length, pump power and phase fluctuation and LO phase fluctuation in the in air bench. They are all analog servos. The unity gain frequency of each loop is designed to be around several kHz.

Filter cavity losses measurement:

Assuming we know reflectivity and transmissivity of input mirror, we could get the information of losses by measuring the reflected power while cavity is on/off resonance.





The measurement is the ratio of filter cavity reflection on/off resonance.

We performed measurement many times. The difference is due to different alignment condition.

Filter cavity mirror characterization:



		diameter 0.05 m		diamet	
Filter cavity mirrors have been characterized by LMA. The data is shown on the right side.				0.02 n	
	Mirror	RMS	PV	RMS	
		(nm)	(nm)	(nm)	(
	#1	1.96	11.5	0.52	3
	#2	2.09	12.2	0.52	3







Homodyne noise spectrum while LO incidence homodyne

homodyne noise from LO

Degradation of squeezing Alignment situation of in current homodyne



By assuming phase noise of 300 mrad and total loss of 30%, we could explain the squeezing we measured for each pump beam power. There are long way to improve.

Squeezing and coherent control: Coherent control is used to stabilize the phase noise of squeezing.



We locked coherent control loop with a very low unity gain frequency(around 80Hz). With the locking of this loop, we could average noise spectrum of homodyne. We could see 4.4dB of squeezing in the

mechanical transfer function of





