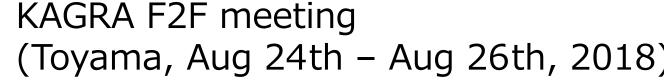
# Fast localization with







# a heterogeneous network of gravitational wave detectors

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# Introduction

We present expected fast sky localisation of coalescing binaries with a hierarchical search using three gravitational wave (GW) detectors, HLV (Hanford/Livingston/Virgo).

A hierarchical search can be used with a network of GW detectors with varying sensitivities, and is aimed at making effective use of the least sensitive detector's information. Here we demonstrate the sky localisation using a hierarchical search with the two higher sensitivity LIGO detectors and the less sensitive Virgo detector, using simulated signals.

## Hierarchical network

For precise source localization: Triple (or more) coincidences

#### At the beginning:

**Detectors with different sensitivity** 

#### For getting more coincidences:

Set a lower threshold, as long as not too many background triggers

Analyze hierarchically!

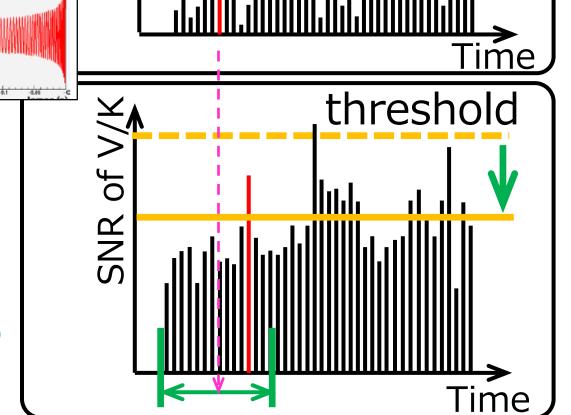
## How to analyze: Lower sensitivity Higher sensitivity Virgo (V) **LIGO** Hanford (H) Higher sensitivity Lower sensitivity **LIGO** Livingston (L) KAGRA(K)

(At the beginning)

Higher sensitivity detectors sub network detects One template  $\stackrel{\sim}{z}$ candidate event.

Less sensitive detectors

- are added into network 1. with lower SNR threshold
- 2. using same parameters &
- 3. a small window around time of double coincidences.



SNR distribution

**Measured** 

E 10-4 In O1

Model

Measurement

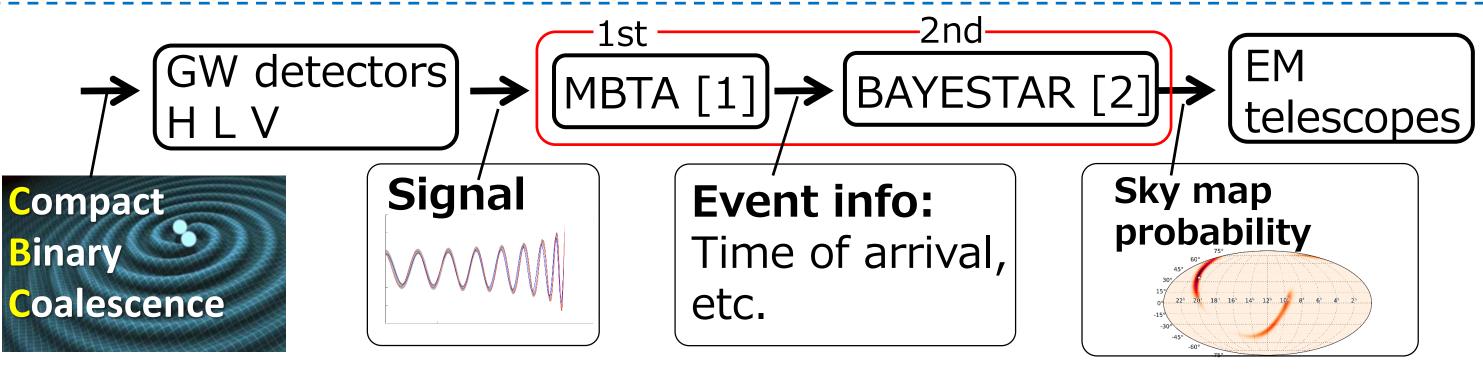
threshold

→ How does this approach improve the localization?

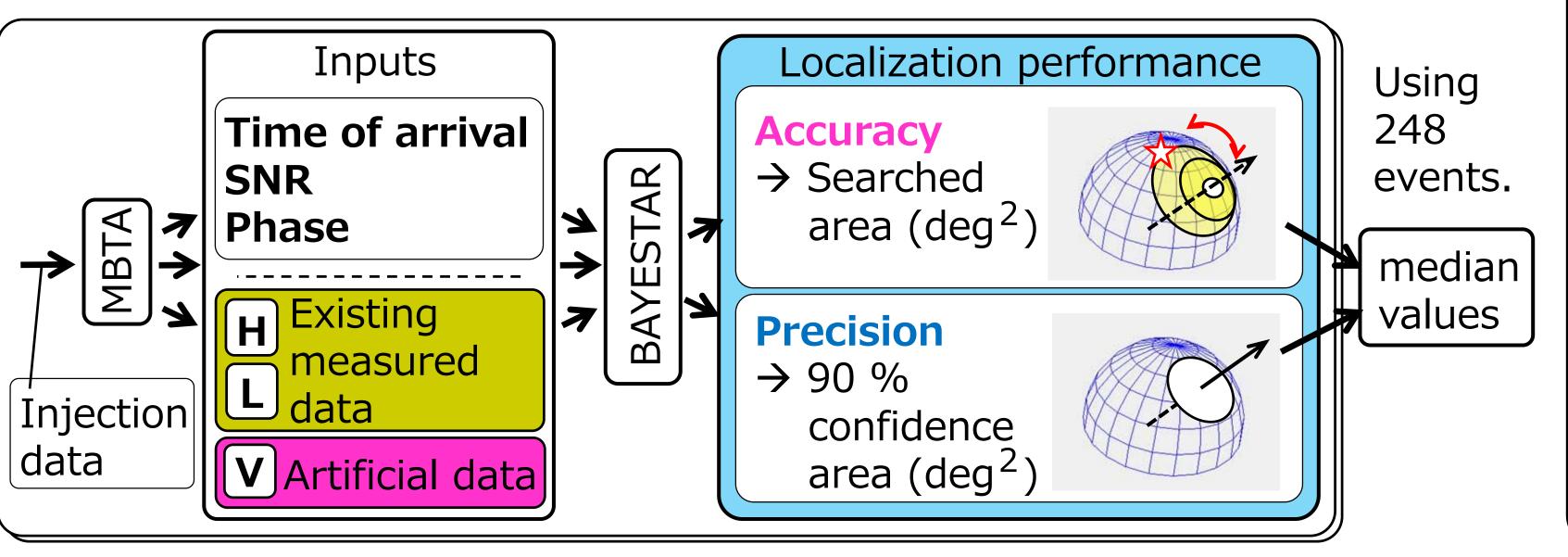
# Calculation setup

#### **Assumptions:**

Higher sensitivity: HL  $\rightarrow$  70 Mpc, Lower sensitivity: V (for 1.4–1.4 M<sub> $\odot$ </sub> BNS range)



#### **Calculation main flow:**



## Generating & mixing artificial Vtriggers

# 1. Generating V triggers

 $V_r$ : V trigger based on random parameters

= random following measurement

 $Time = t_{H1} \text{ or } t_{L1}$ 

+ random [-35ms:35ms] Phase = random  $[0:2\pi]$ 

## $V_i$ : V trigger based on injection parameters

= metadata + Gauss(0,1)

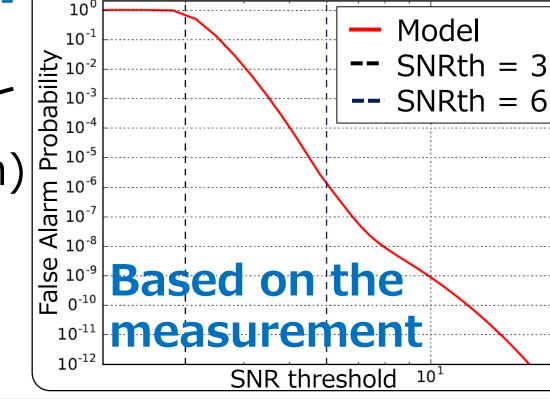
Time = metadata + Gauss(0,0.66 ms\* $\frac{6}{SNR}$ )

Phase = metadata + Gauss(0,0.25 rad)

#### 2. Mixing HLV triggers

p = random [0:1]FAP = FAP(SNR) or FAP(SNRth)

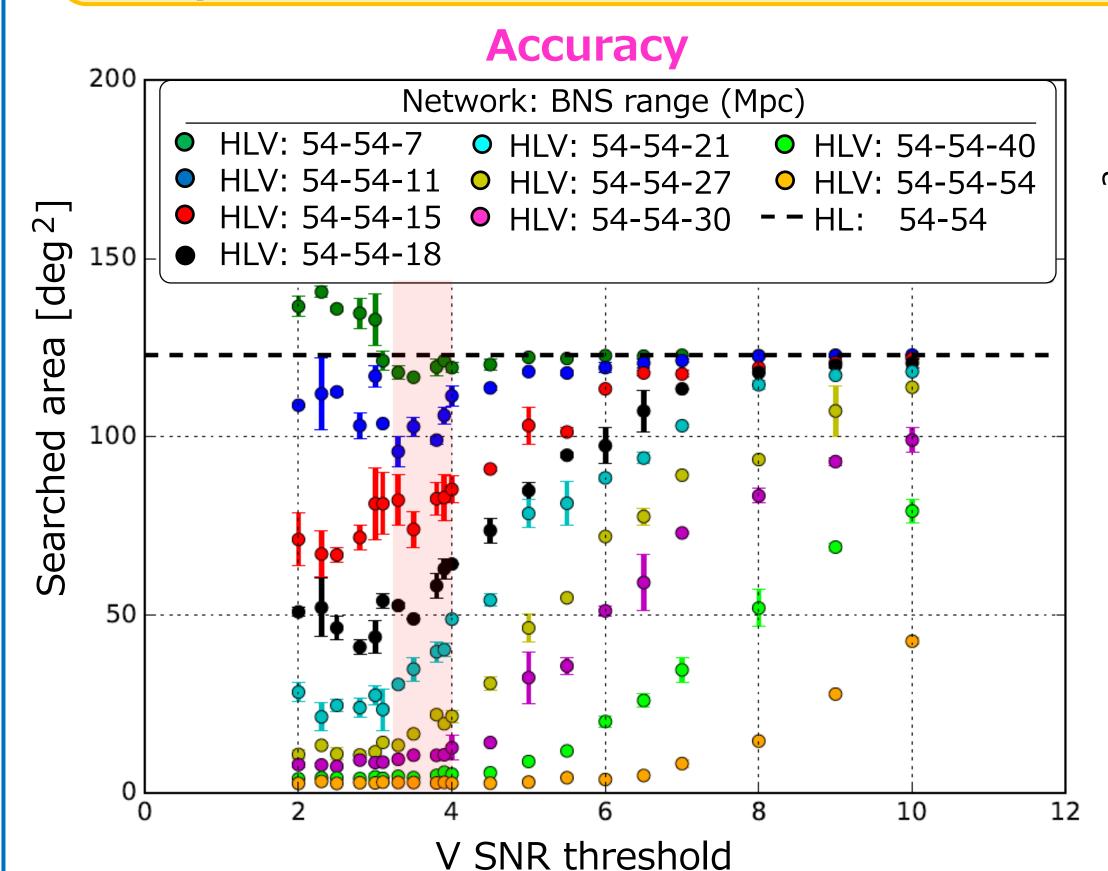
p<FAP  $\rightarrow$  HLV<sub>r</sub> p>FAP & SNR>SNRth → HLV<sub>i</sub> p>FAP & SNR<SNRth → HL

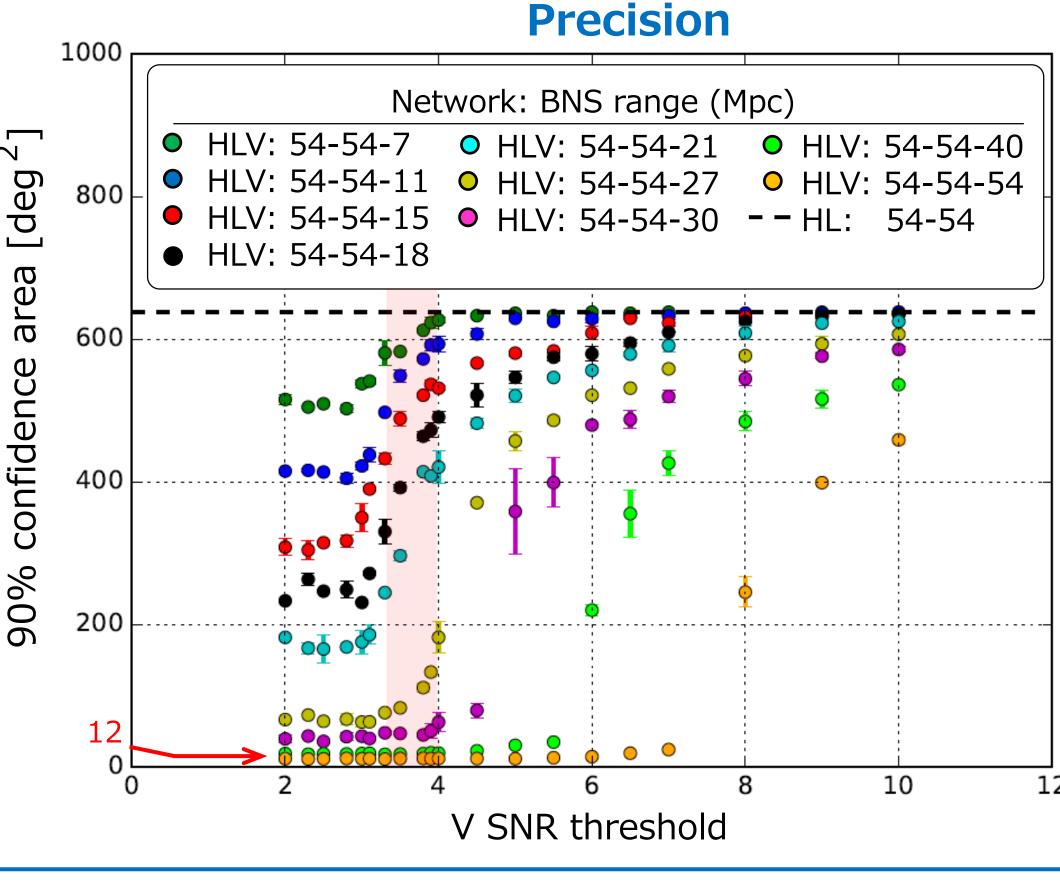


# Expected performance with HLV hierarchical network

Sky localization performance, when SNR threshold for HL is 5:

→ Optimal SNR threshold for V detector? → Dependence of V detector's sensitivity?





#### **Conclusion:**

The hierarchical network improves:

1. localization accuracy

effectively at V SNRth  $\sim$ 3.5, if V range is greater than 11 Mpc

2. localization precision

at any sensitivity.

**HLV-hierarchical network using** HL: 54 Mpc, SNRth = 5 and V:>11 Mpc, SNRth ~3.5 effectively improves the sky localization as shown by the coloured bands.

## Summary

- 1. We investigated the expected fast localization performance with a heterogeneous network using HLV.
- 2. We demonstrated that the hierarchical network effectively improved the accuracy & precision when V threshold is set to  $\sim$ 3.5, if BNS range of V detector is greater than 11 Mpc.
- 3. The hierarchical search will be most useful when adding new detectors, which are less sensitive as they are undergoing commissioning, to the network.

#### Ongoing work:

1. Investigate the localization with HLVK hierarchical network

References: [1] T. Adams et. al., Class. Quant. Grav. 33 (2016) [2] L. P. Singer, L. R. Price, Phys. Rev. D **93**, 024013 (2016)