

# Commissioning works for KAGRA

KAGRA International Workshop  
@Ewha Womans University, Seoul, Korea  
2018/6/29(Fri.)

Osamu Miyakawa  
and KAGRA Collaboration

- iKAGRA commissioning: 2016/3~2016/4
  - Simple 3km Michelson interferometer with room temperature, and no mirrors in vacuum
- bKAGRA commissioning: 2017/9~
  - **phase I** (~2018/5): Michelson with low temperature operation
  - phase II (~2019 middle): full configuration
  - phase III (2019 middle~): observation

- Stage 1 (2017/9/25~)
  - Beam alignment to EX and EY area.
- Stage 2 (2017/12/4~)
  - Beam alignment from ETMY to center area
- Stage 3 (2018 March, April)
  - Beam alignment from ETMX to center area
  - Locking of MI with cryogenic ETMY with calibrated sensitivity curve
- ~~Stage 4 (April)~~
  - ~~– Locking of MI in vacuum with both cryogenic end~~
  - ~~– Operation of IFO~~

- 9/19: All PR suspension installation completed
- 9/21: BS suspension installed
- 10/19: Main beam reached X end
- 10/26: Main beam reached Y end



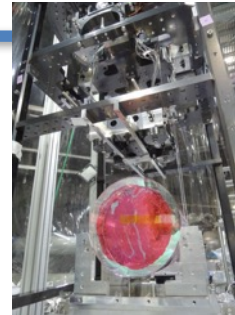
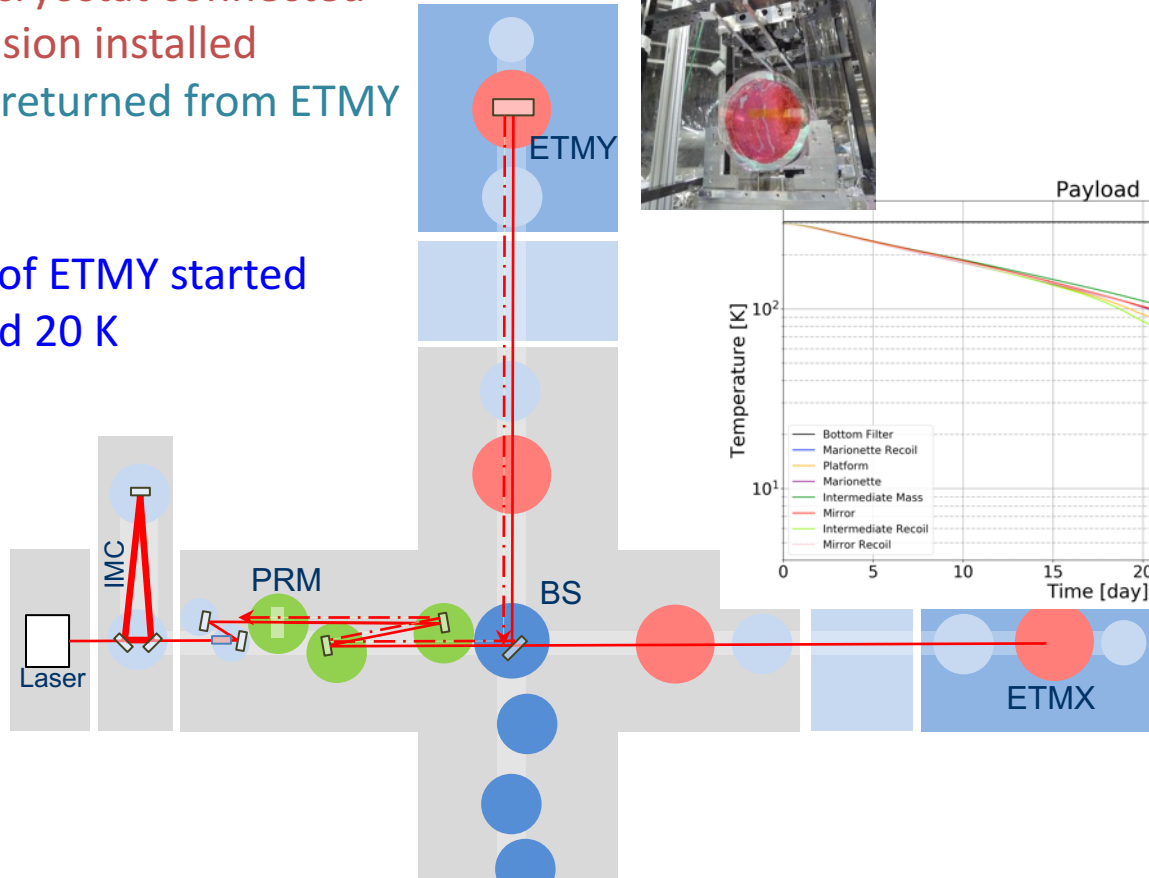
11/27-12/1: Commissioning Work Shop

- 11 visitors from LIGO, VIRGO
- 13 KAGRA members, 3 LOCs

11/29: Type A and cryostat connected  
 12/1: ETMY suspension installed  
 12/19: Main beam returned from ETMY

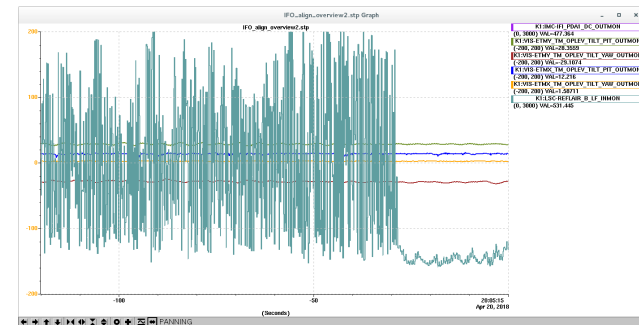
2018

2/7: Cooling down of ETMY started  
 3/11: ETMY reached 20 K



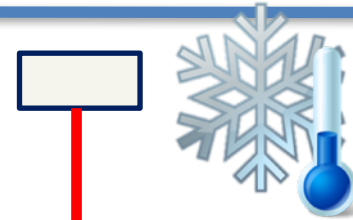
- It took ~2months from connection Type A and Cryostat to start cooling down. Is it possible shrink time for the next mirrors?
  - ✓ Some troubles happened for the first time work, but it can be avoided in the next time.

- Installation of ETMX was **much faster** with the experiences of ETMY.
  - ✓ Behavior of ETMX is also much smoother than ETMY.

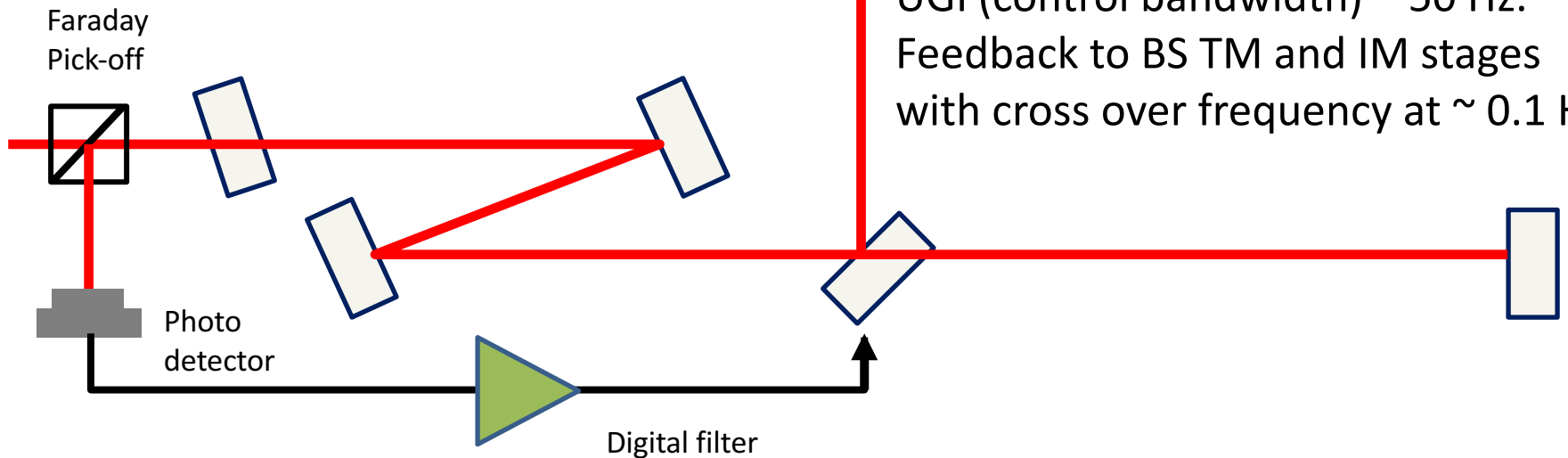


- Infrastructure
  - PRs, BS, ETMs are implemented and worked.
  - All mirrors in vacuum: 3km arms, X-end, Y-end and central area.
  - Cooled ETMY with a spare mirror.
  - Room temperature ETMX with a real mirror.
- IFO
  - 3km Michelson lock: ~10hours (longest).
  - Sensitivity:  $2\sim 3e-17/\text{rHz}$
  - 9 days operation.

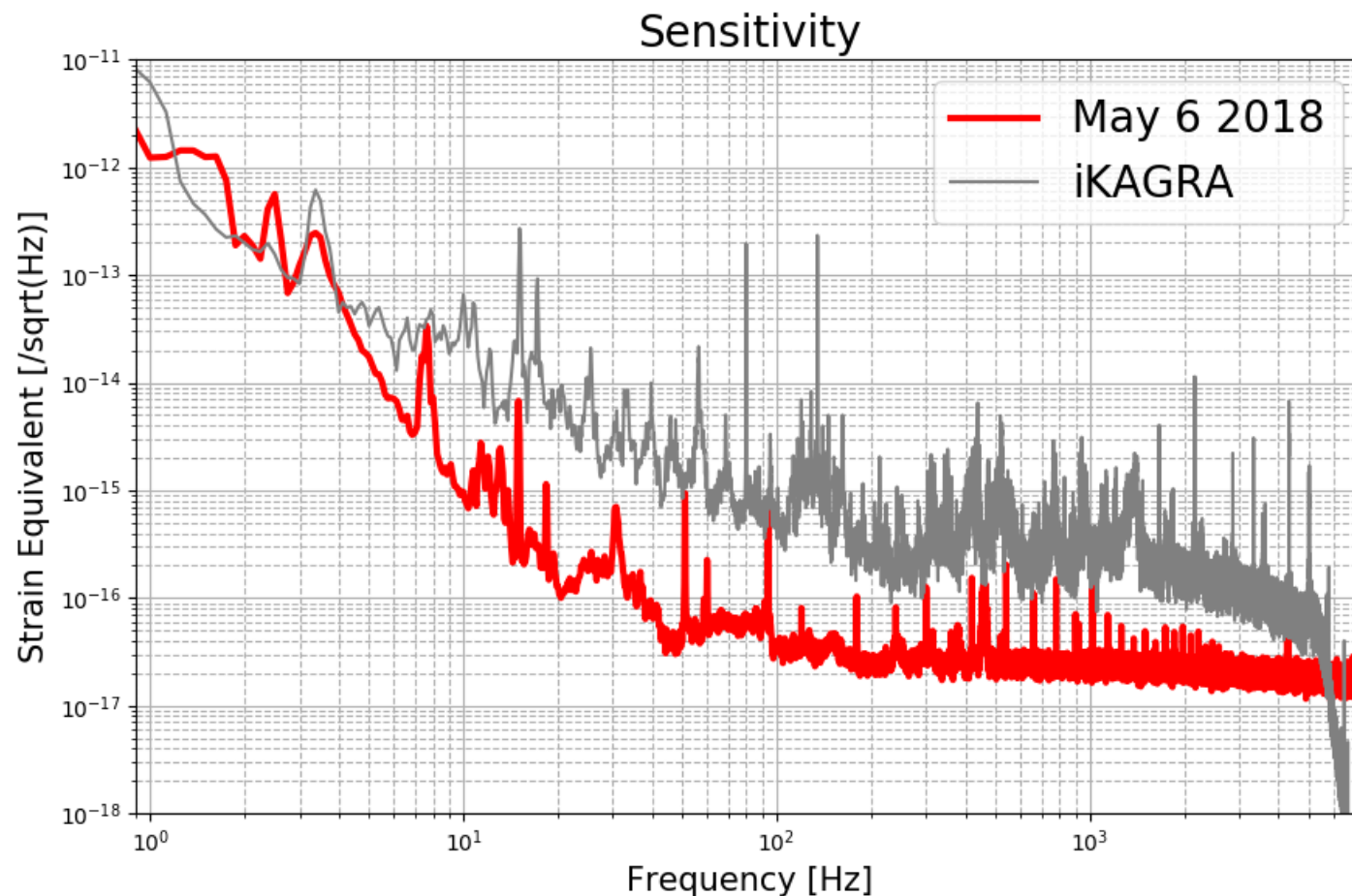
# LENGTH CONTROL



UGF(control bandwidth)  $\sim 50$  Hz.  
Feedback to BS TM and IM stages  
with cross over frequency at  $\sim 0.1$  Hz



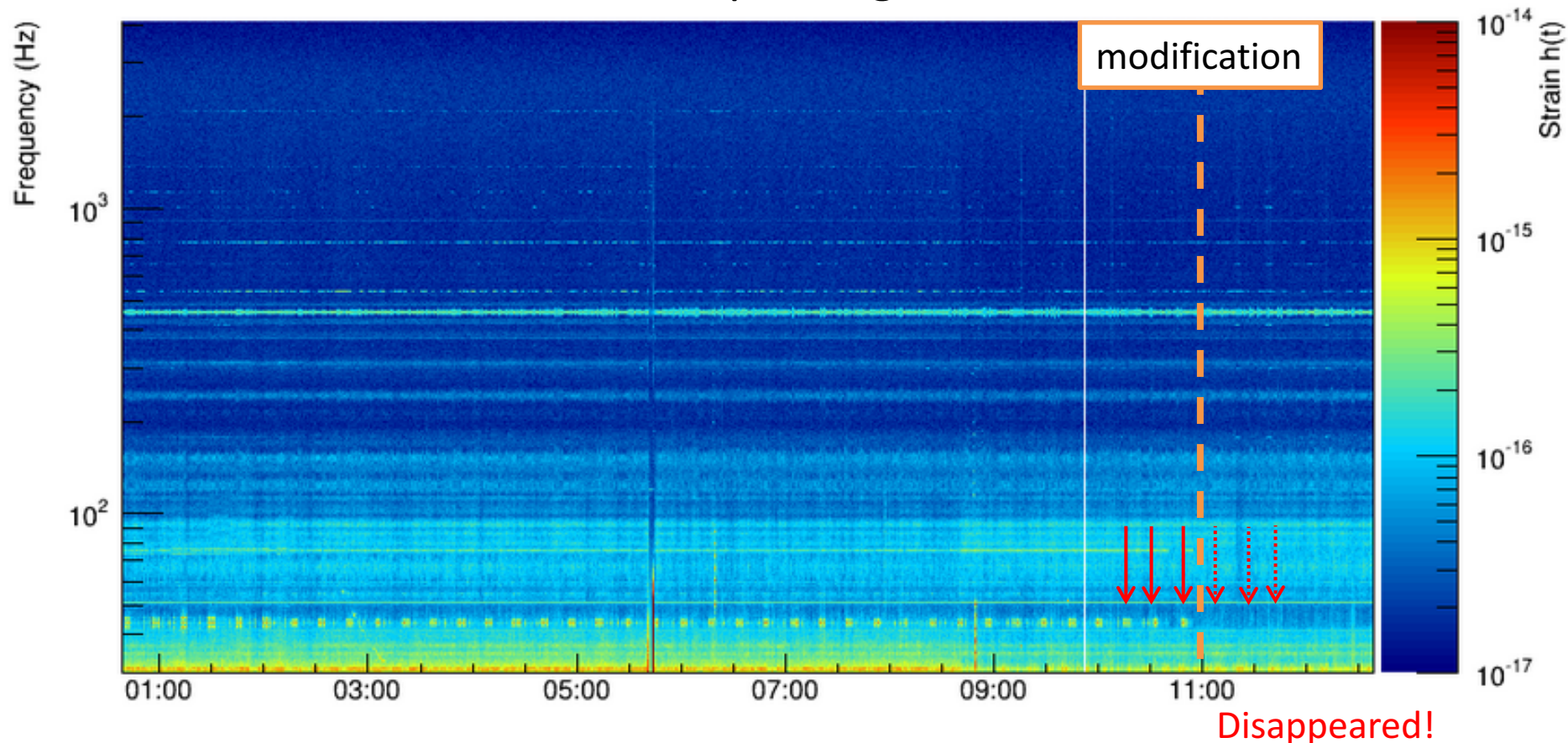




- Sensitivity was improved 1~1.5 order from iKAGRA in almost frequency.
- This is mainly because all the mirrors were in vacuum, and acoustic noise was reduced.

- It turned out this noise has coherence with PR2 optical lever signal
- => we modified optical lever control loop of PR2

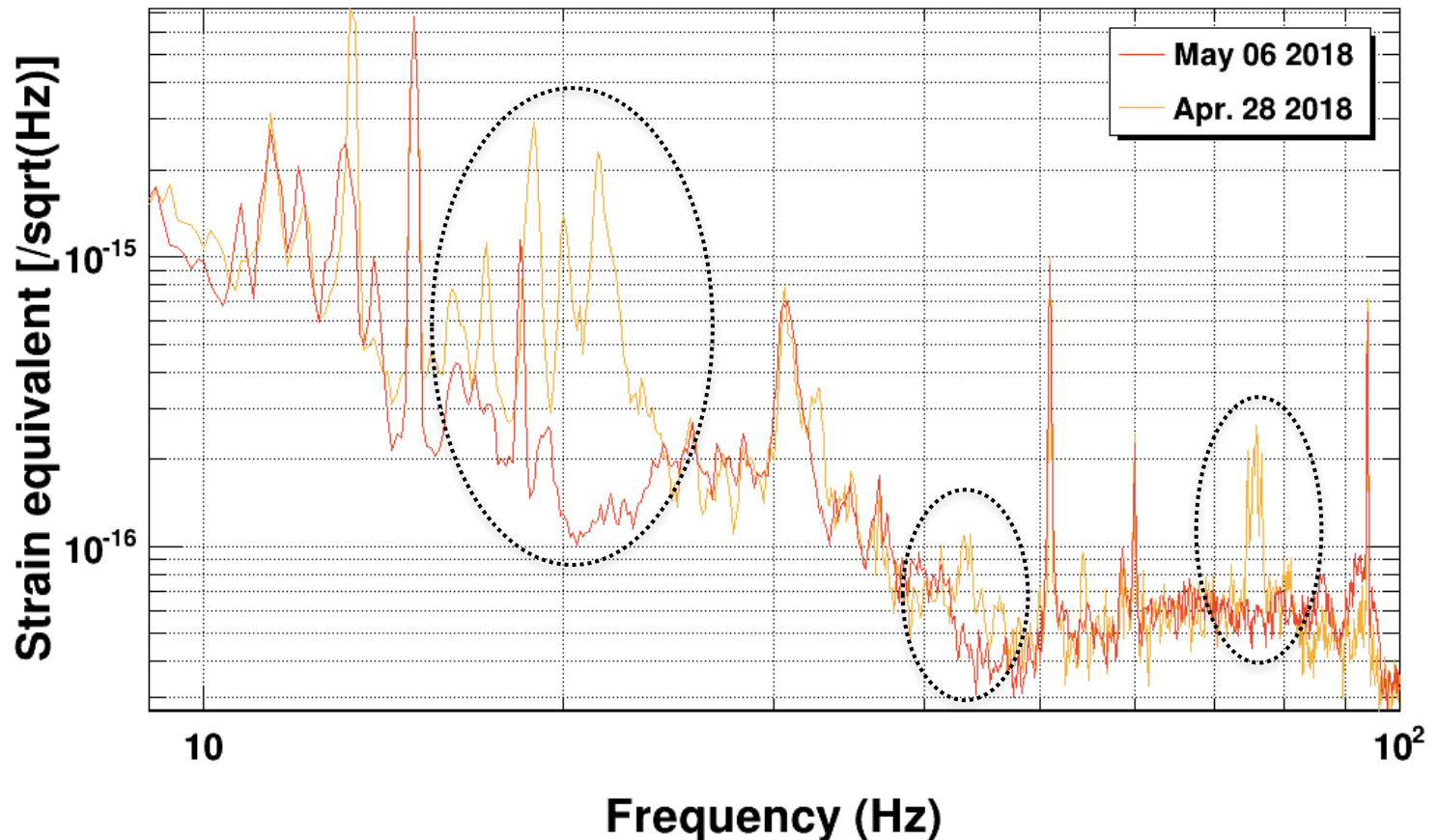
### MICH spectrogram



- It turned out this noise has coherence with PR2 optical lever signal  
=> we modified optical lever control loop of PR2

Strain  $h(f)$

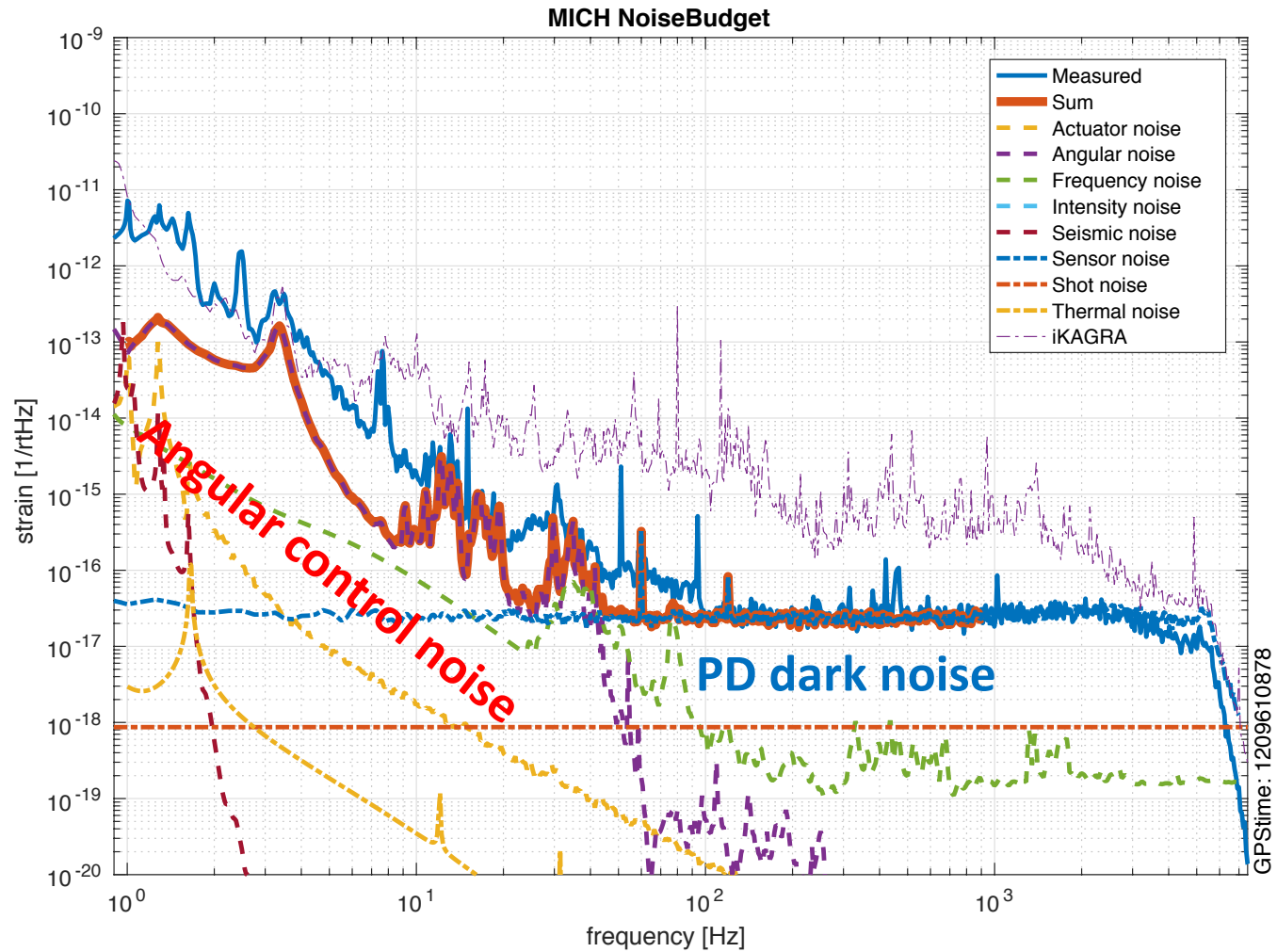
Some other structures also vanished thanks to this.



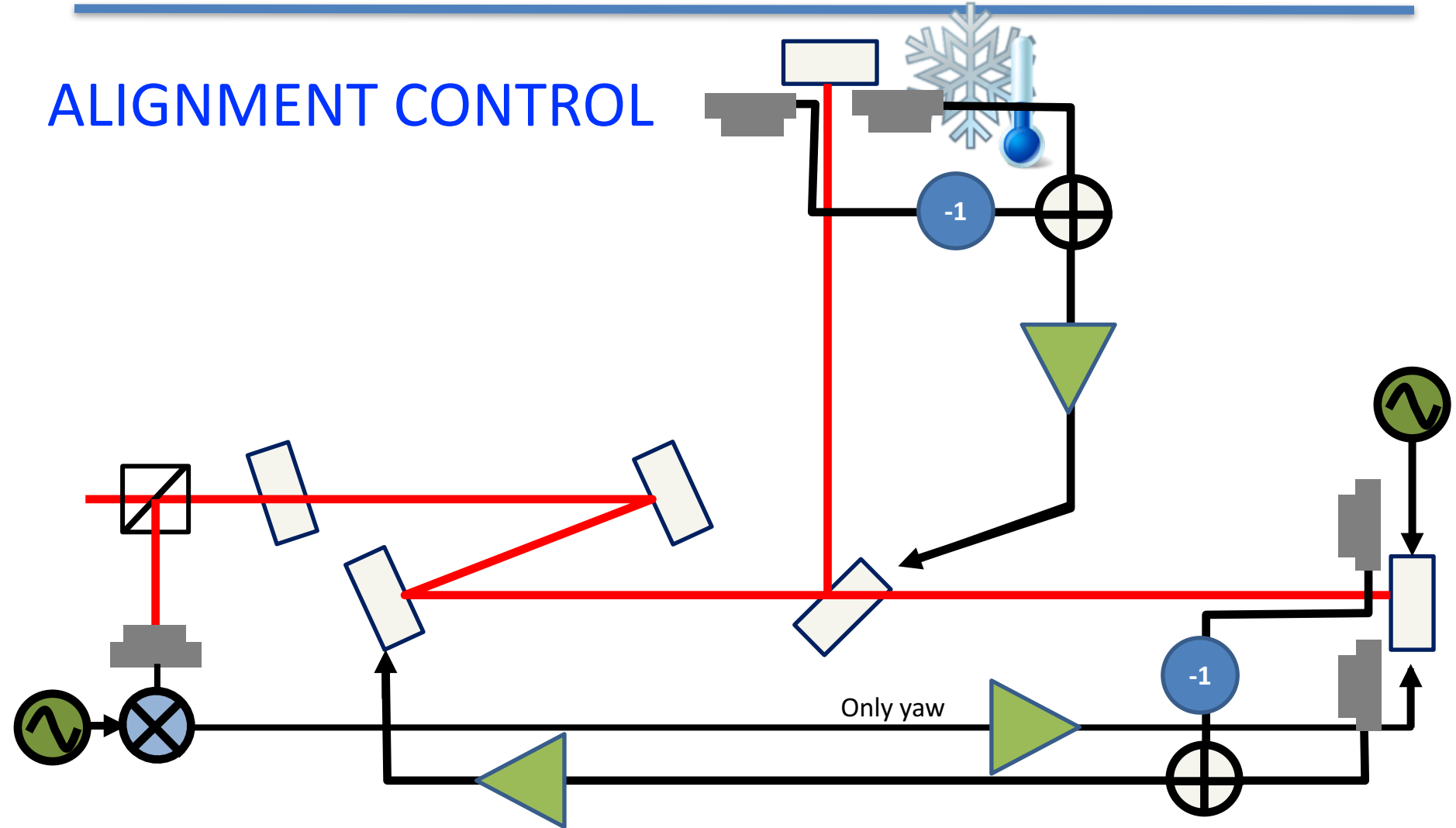
\*T0=06/05/2018 03:01:00

\*Avg=27

BW=0.187493



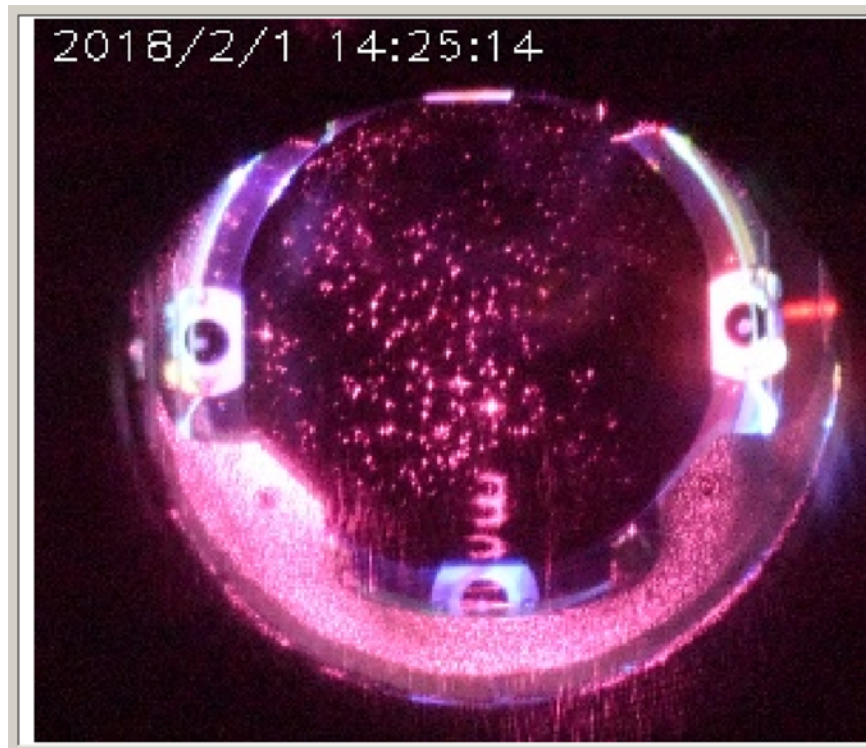
# ALIGNMENT CONTROL



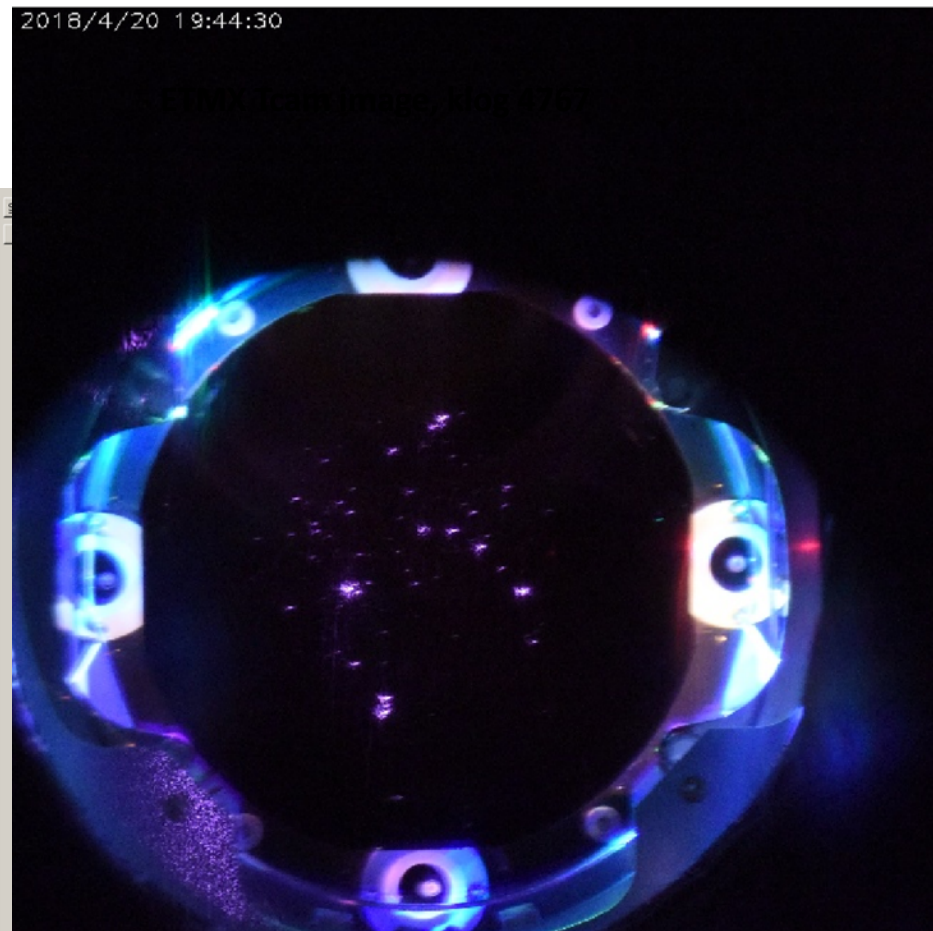
- Manual fine adjustment ~ every half a day
- No global control on ETMY angle
  - Optical lever local control was good enough
- ETMX needed a global control
  - Servoing to optical lever had a long term drift issue (~ a day)
  - Global control lets ETMX follow ETMY
  - Seemingly worked fine



ETMY Tcam image: spare mirror



ETMX Tcam image: real mirror



- Observed in ETMY, PRs 2 and 3, and BS controllers.



Klog 4254

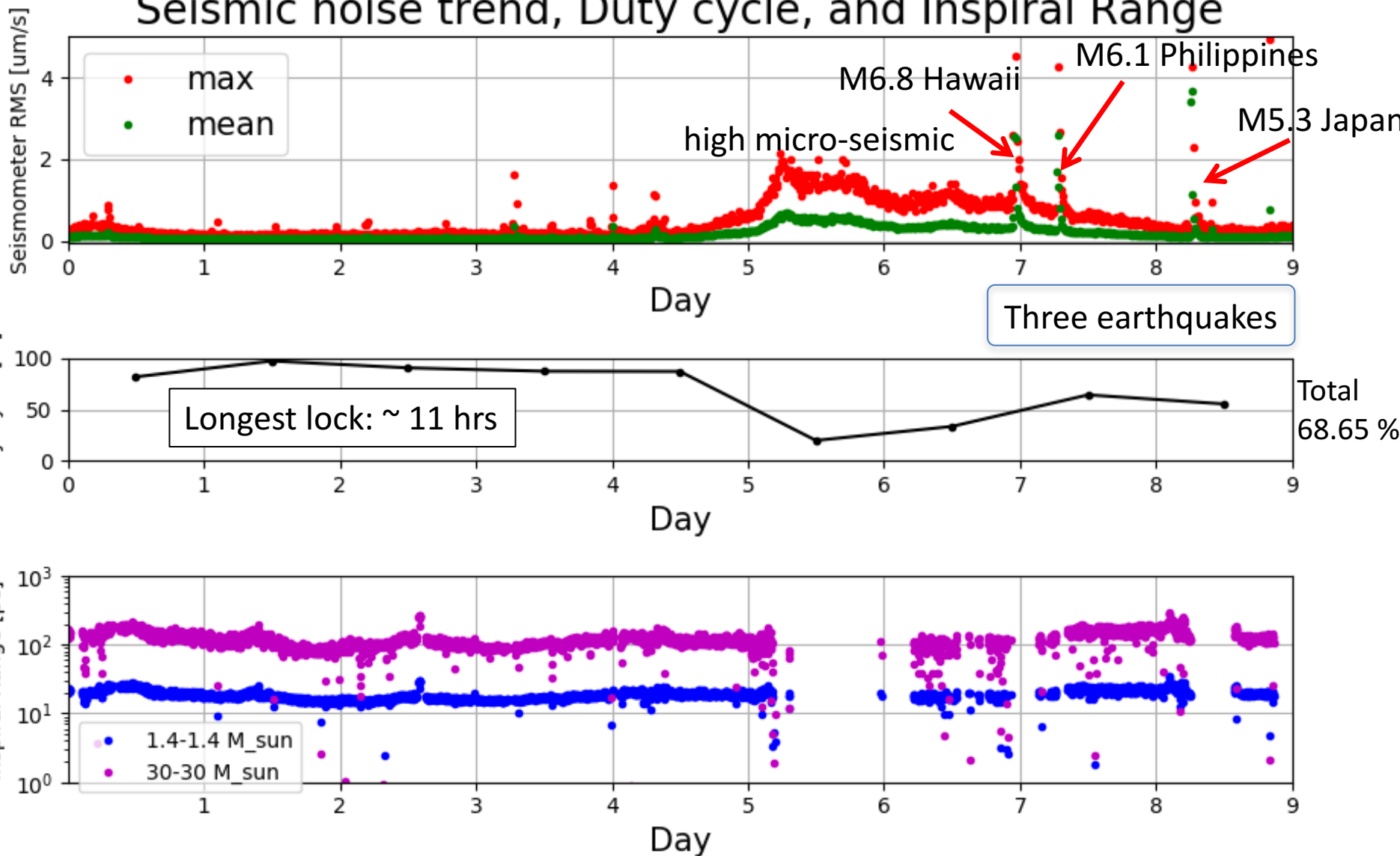
- This is because of too much load on CPU for real-time control PC.
- Reducing the sampling rate (16 -> 2 kHz) reduced the glitch rate to an OK-level.



	4/28	4/29	4/30	5/1	5/2	5/3	5/4	5/5	5/6	5/7
Day (9:00-17:00)	OLG measure ments	Type-A Yend TRF	BS TRF	Type-A Xend TRF	Noise injection Center	Noise injection YEND	Schnupp Assymmetry & IFO noise budget	Noise injection XEND	CRY Extra EXP. 1 & 2	Phase 2
Night (17:00-9:00)	OLG measure ments	Type-A Yend TRF	CW injection	Type-A Xend TRF	CBC injection	CBC injection	Schnupp Assymmetry & IFO noise budget	OLG measure ments	CRY Extra EXP. 1 & 2	
Parallel	Data transfer, Pipeline tests, GIF									

- Originally this was planned that 4/23(Mon) started.
- Actual starting date was 4/28(Sat), that was 5days delay due to vacuum leakages.
- IFO was also helped by this delay. The first lock was 4/20. We needed some days for tuning interferometer.
- Some of results of this operation will be talked in this workshop.

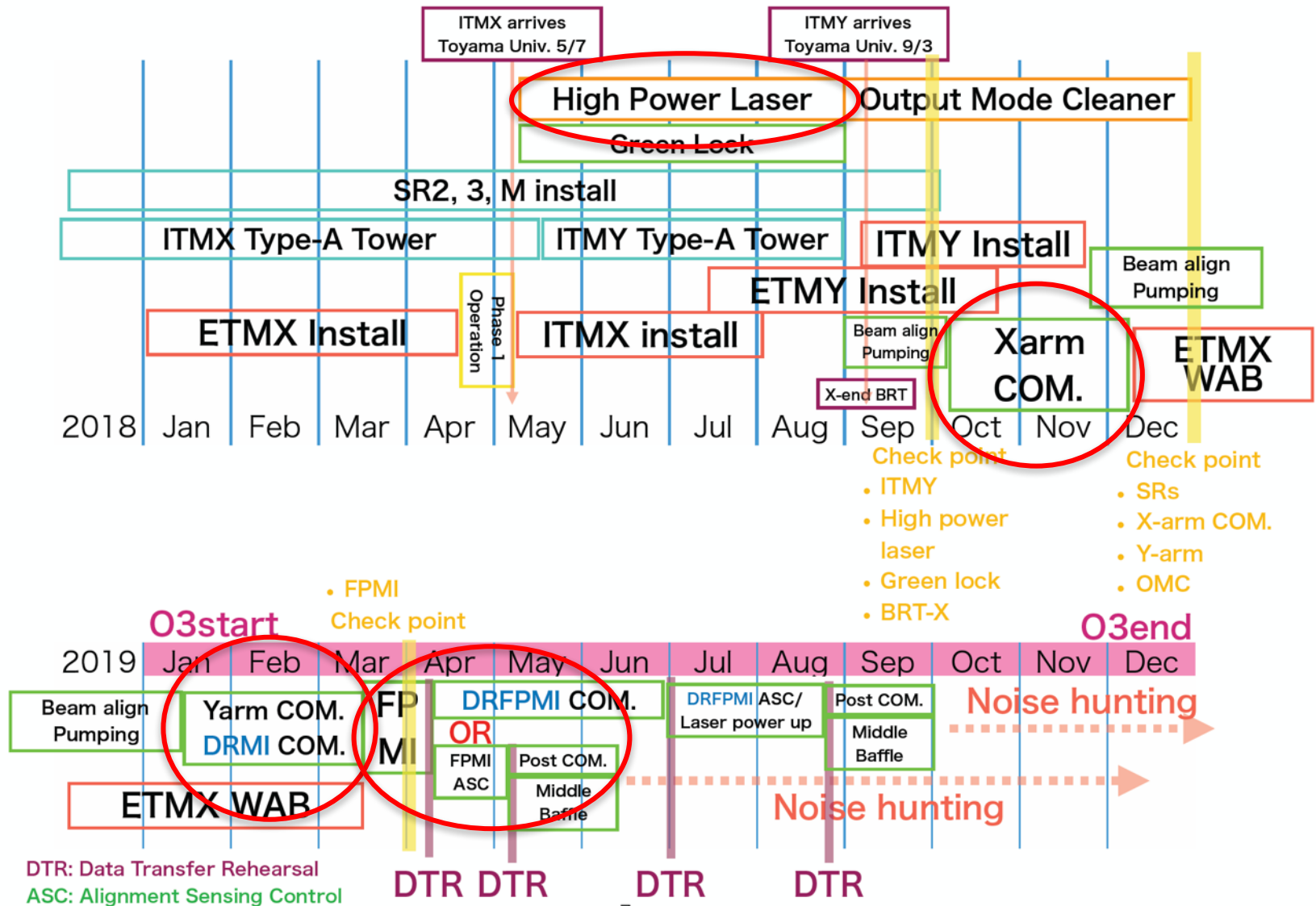
## Seismic noise trend, Duty cycle, and Inspiral Range



- Proposing several days engineering run after each step.
  - HPL stage: for High Power Laser
  - XA stage: 1 day engineering run for X Arm lock.
  - YA stage: 3 day engineering run for Y Arm lock.
  - FPM stage: 1week engineering run for Fabry Perot Michelson lock.
  - RSE stage: 1week engineering run for RSE lock.
- Additional run.
  - 2 weeks run IFO with some sensitivity. This will be a good training for O3.

2018/6/18

## Not Using ITMY dummy case



- Commissioning of bKAGRA phase 1 was successfully done.
  - Cryogenic Michelson was operated.
  - ETMY was cooled down to 20 K.
- We had 1 week operation. We experienced a lot of things and obtained many results.
  - Issues around the payload was identified to some extent.
- Commissioning of bKAGRA Phase 2 started.
  - Installation and preparation for joining late O3 is NOW on-going.
  - High Power Laser will be provided by the end of August.