

# Double Balanced Homodyne Detection and Fabry-Perot gravitational-wave detector

Kouji Nakamura (NAOJ)  
with Masa-Katsu Fujimoto (NAOJ)

This research is on a readout scheme of gravitational-wave detectors.

In interferometric gravitational-wave detectors, gravitational-wave signals are included in the output photon quadratures as

$$\hat{b}_\theta(\Omega) = \cos \theta \hat{b}_1(\Omega) + \sin \theta \hat{b}_2(\Omega) = R(\Omega) (\hat{h}_n(\Omega) + h(\Omega))$$

$$\left\{ \begin{array}{l} \hat{b}_1(\Omega) : \text{amplitude quadrature (in two-photon formulation); } \hat{b}_2(\Omega) : \text{phase quadrature (in two-photon formulation);} \\ \theta : \text{homodyne angle; } R(\Omega) : \text{response function of the detector (complex function);} \\ h(\Omega) : \text{gravitational-wave signal (classical function); } \hat{h}_n(\Omega) : \text{noise operator (signal referred).} \end{array} \right.$$

In the previous KAGRA f2f meeting, we showed that the expectation value of the operator  $\hat{b}_\theta(\Omega)$  can be measured through the “double balanced homodyne detection.”

In this poster, we discuss an application of this readout scheme to a Fabry-Perot GW detectors as an example.

Although the advantage of the frequency-dependent homodyne detection is lost in our double balanced homodyne detection, this example indicates that *we should discuss the signal-noise trade-off relation through the total quantum measurement process including its readout scheme*, since the quantum measurement process of GW detectors is completed by the inclusion of its readout scheme.

See you at Poster C02 !!