

# PRFPMI or RSE for Joining O3

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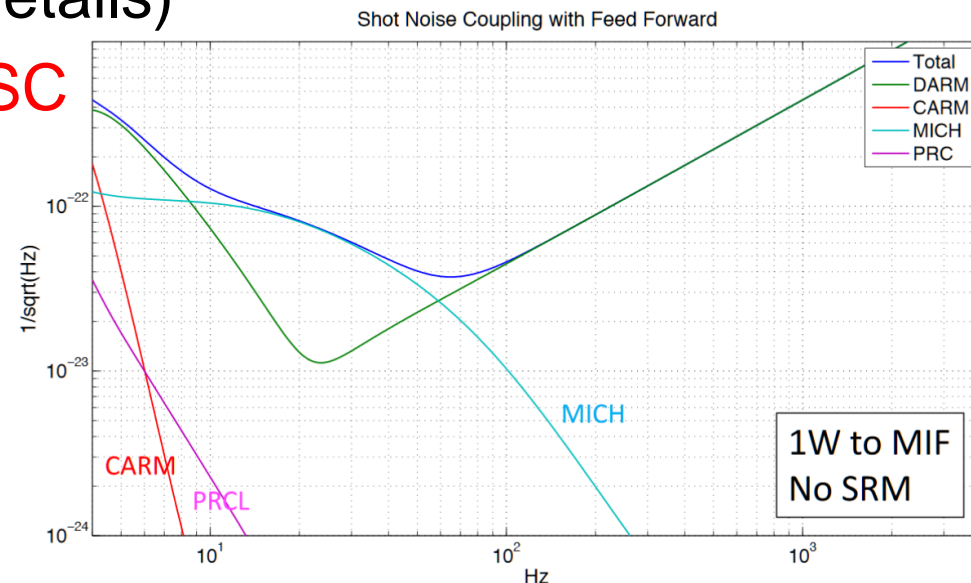
# Summary on PRFPMI or RSE

	<b>PRFPMI</b>	<b>RSE</b>
Locking scheme	green locking + f3 (or develop new scheme from scratch)	green locking + f3
# of degrees of freedom	LSC: 4, ASC: 10	LSC: 5, ASC: 11 → more DoF
Tolerable excess noise to achieve AdV O2 sensitivity	~ x4 O1 level (with 10 W input) → requires more noise hunting	~ x8 O1 level (with 10 W input)
SRM Installation	Blank SRM	70 % SRM

**Our suggestion: RSE for O3**

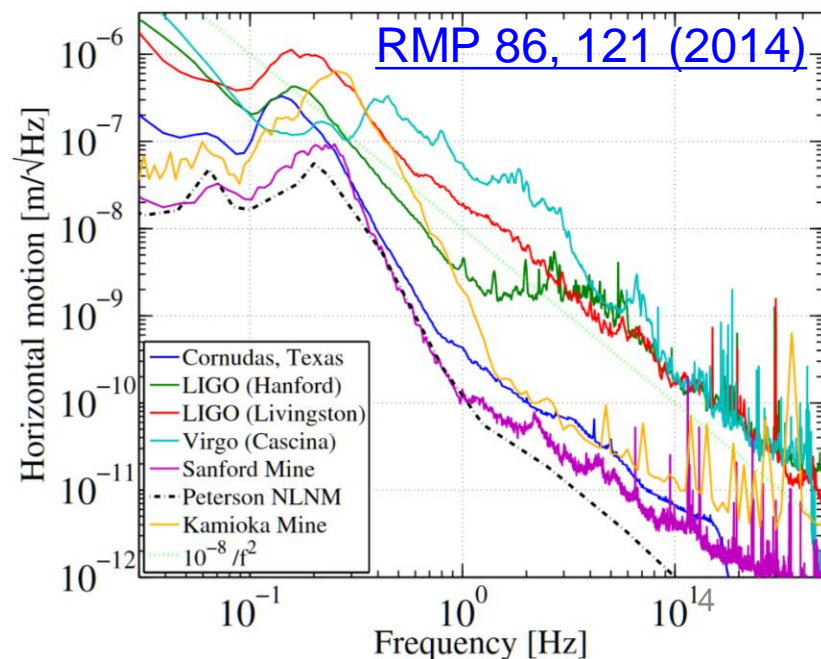
# Locking Scheme

- Green locking and f3 AM generation is **necessary for both PRFPMI and RSE**
- For **PRFPMI**, other locking scheme is also possible if we develop new scheme **from scratch**
- In **PRFPMI** case, resonant condition for f1 will change, which result in **~ x3 worse shot noise**, and couples to DARM (see [JGW-G1707479](#) for details)
- We also need to design **ASC** for **PRFPMI** from scratch
- New locking scheme for **PRFPMI** without green will likely to be inapplicable to **RSE**



# New Locking Scheme for PRFPMI?

- Advanced Virgo could lock **PRFPMI** because they had long experience on their variable finesse scheme
  - detailed time-domain simulation (e2e)
  - guided lock
  - experienced people from Initial Virgo
- KAGRA has **higher arm cavity finesse** (1530) than Advanced LIGO (450) and Advanced Virgo (460), or any other 1<sup>st</sup> gen. detectors
- Micro-seismic noise for KAGRA is not low.
  - Stochastic locking will be tougher
  - Locking KAGRA **PRFPMI** without green **will be tough**



# Sensitivity

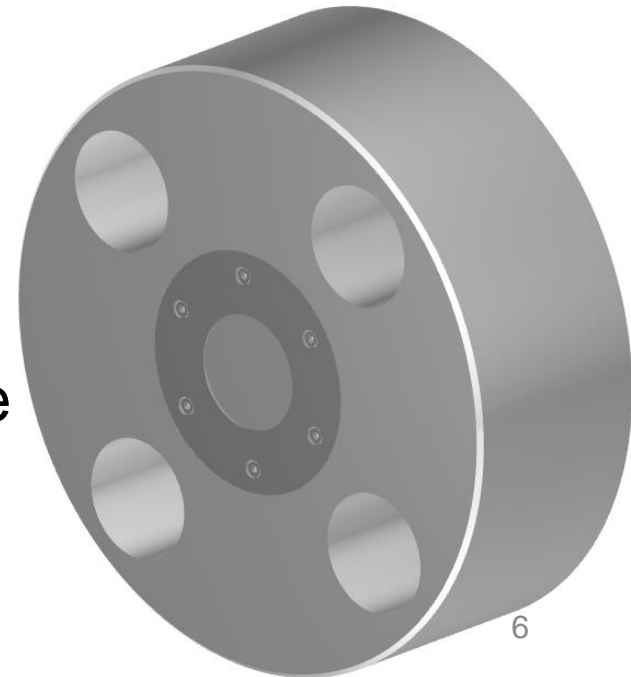
- Inspiral range of **PRFPMI** relies on low frequencies
- **PRFPMI** requires **more noise hunting** than **RSE** to achieve same inspiral range
- See [JGW-T1707334](#) for details

<b>Excess noise</b>	<b>PRFPMI</b>	<b>RSE</b>
No excess	BNS: 58 Mpc BBH: 0.82 Gpc	BNS: 93 Mpc BBH: 1.4 Gpc
x1 O1 level (~ KAGRA suspension thermal noise)	BNS: 48 Mpc BBH: 0.66 Gpc	BNS: 71 Mpc BBH: 1.1 Gpc
x4 O1 level	BNS: 27 Mpc BBH: 0.37 Gpc	BNS: 42 Mpc BBH: 0.62 Gpc
x8 O1 level	BNS: 19 Mpc BBH: 0.26 Gpc	BNS: 30 Mpc BBH: 0.45 Gpc

\* Assumed 10 W input for inspiral range calculation. Includes shot noise coupling <sup>5</sup>

# Schedule

- Both **PRFPMI** and **RSE** requires SRM installation
  - Blank one for **PRFPMI** for mode-matching to OMC
- Both **PRFPMI** and **RSE** requires green locking and f3 AM
  - If we give up green locking, **PRFPMI** requires development of new locking scheme.
- Giving up green locking at this point **do not accelerate** the schedule
  - Independent manpower from CRY and VIS
- We already ordered blank SRM and 70% SRM. Switching to blank one can be done **at later stages** if we had some trouble in locking DRMI (~Dec 2018) or RSE (~ March 2019)



# Our Suggestion: RSE for O3

- PRFPMI can be done without green only if we successfully develop new locking scheme from scratch. Concentrating our resources to green (not new scheme) seems to be a better idea.
- RSE requires one more degrees of freedom to lock. PRFPMI requires more noise hunting. Latter is more unpredictable in terms of scheduling.
- Switching to PRFPMI from RSE can be done at later stages, only by replacing SRM. Switching to RSE from PRFPMI without green is almost impossible.
- Making a solid schedule is important for joining O3. PRFPMI relies on fragile assumptions (new locking scheme not guaranteed and more noise hunting).