Parameters for the latest estimated sensitivity of KAGRA

Kentaro Komori, Yuta Michimura, Kentaro Somiya

September 18, 2017



Figure 1: Comparison of KAGRA sensitivity curves with different detune and homodyne phase. These sensitivities are optimized on maximum inspiral range of a binary neutron star merger.

1 Parameters

Parameters
VRSE
$3 \mathrm{km}$
$1064~\mathrm{nm}$
$674~{\rm W}$ $^{\rm i}$
0.004 [1]
0.92 [1]
$100 \text{ ppm}^{\text{ii}}$
0.002 ⁱⁱ
0.1 $^{ m ii}$
$86.5^{\circ iii}$
$119.1^{\circ iii}$
$135.1^{\circ \text{ iii}}$

Table 1: Interferometer parameters reviewed by MIF chief Y. Michimura.

Item	Parameters
Material	Sapphire
Radius	$11 \mathrm{~cm}$
Thickness	$15~\mathrm{cm}$
Mass	$22.8 {\rm ~kg} {\rm ~iv}$
Temperature	$22~{ m K}^{-{ m i}}$
Loss angle	$1.0 \times 10^{-8} [2]$
Absorption	50 ppm/cm [3]

Table 2: Parameters of the mirror reviewed by MIR chief E. Hirose.

ⁱsee section2.

ⁱⁱrequirement.

ⁱⁱⁱBNS inspiral range optimized.

 $^{^{\}rm iv}{\rm reported}$ by E. Hirose, based on CAD drawing, including the mass of the ears.

^vreported by E. Hirose.

^{vi}measured to be 0.35 ppm by LMA.

^{vii}reported by A. Hagiwara, based on CAD drawing. Actual mass of IM could vary depending on the masses of the parts.

viii requirement. Measured to be 2e-7 with bulk [6], 1e-5 with 20um wire [7].

^{ix}reported by K. Yamamoto.

^xreported by T. Ushiba

^{xi}requirement calculated by [9]

^{xii}reported by T. Ushiba based on CAD drawing.

Item	Parameters
Material	silica/tantala
Numbers of coating layers	22/40 (ITM/ETM) ^v
Loss angle of silica	$3.0 \times 10^{-4} [4, 5]$
Loss angle of tantala	$5.0 \times 10^{-4} [4, 5]$
Beam Radius	3.5 cm [1]
Absorption	$0.5~{ m ppm}^{ m vi}$

Table 3: Parameters of the coating reviewed by MIR chief E. Hirose.

Item	Parameters
IM suspension	
Material	CuBe
Number	4
Length	$26.1 \mathrm{~cm}^{\mathrm{~vii}}$
Diameter	$0.6 \mathrm{~mm}^{\mathrm{~vii}}$
Loss angle	5×10^{-6} viii
IM mass	$20.5 { m ~kg} { m ^{vii}}$
Temperature	$16 \mathrm{~K}^{\mathrm{~ix}}$
Blade spring	
Material	Sapphire
Number	4
Horizontal res. freq.	$2 \rm ~kHz~^x$
Vertical res. freq.	14.5 Hz (with TM) [8]
Loss angle	7×10^{-7} xi
Mass	$55~{ m g}^{-{ m xii}}$
Temperature	$16~{ m K}~{ m ^{ix}}$
TM suspension	
Material	Sapphire
Number	4
Length	$35~{ m cm}^{ m vii}$
Diameter	$1.6 \mathrm{~mm}^{\mathrm{~vii}}$
Loss angle	2×10^{-7} [9]
Temperature	$19~\mathrm{K}$ (average of 16 K and 22 K)
Vertical-Horizontal Coupling	1/200

Table 4: Parameters of the suspension reviewed by CRY group K. yamamoto and T. Tomaru.

2 Process to decide power and mirror temperature

In this section we describe how to set laser power and mirror temperature. Laser power and mirror temperature are in one-to-one correspondence via the mirror absorption. Firstly, the sensitivity of DRSE configuration is optimized with changing homodyne and detune angle, and mirror temperature. The best temperature is 22 K, and then corresponding laser power at BS is calculated to be 674 W to keep the mirror temperature 22 K. In broadband configuration without homodyne detection, the best sensitivity in terms of inspiral range of a binary neutron star merger, 131 Mpc, can be reached at the power of 1250 W and the temperature of 25 K. However, power over 1 kW at BS is too high and 128 Mpc can be realized even at 674 W and 22 K. Therefore, we decided to fix the power and temperature. The relationship between laser power and mirror temperature is derived as follows.

Heat flow of fibers in steady state is the same as energy absorbed by a suspended mass,

$$K_{\rm abs} = N_f \int_{T_u}^{T_l} \frac{S_f \kappa(T)}{l_f} dT, \qquad (1)$$

where K_{abs} is the absorbed energy, N_f is the number of fibers, T_u (T_l) is the temperature of an upper (lower) end, S_f is the area, l_f is the length, and $\kappa(T)$ is the thermal conductivity of the fiber. A marionette of the suspension system is connected to a platform by a thick heat link and the CuBe fibers have high thermal conductivity. Therefore, temperature of the IM, CuBe fibers and the marionette can be cooled down to $T_u = 16$ K. Here we decide that the test mass temperature is $T_l = 22$ K. $\kappa(T)$ is measured by Sascha to be [10]

$$\kappa(T) = 7.98 \times T^{2.2}.\tag{2}$$

Substituting $N_f = 4$, $l_f = 0.35$ m, and $S_f = \pi (1.6 \times 10^{-3}/2)^2$ m², total heat flow is calculated as $K_{abs} = 0.724$ W.

The absorbed energy can be divided to that from sapphire substrate, coating, and radiation from outside,

$$K_{\rm abs} = 2\beta_{\rm sub}t_m P_{\rm mich} + \gamma_{\rm coa}P_{\rm circ} + K_{\rm rad},\tag{3}$$

where β_{sub} is the absorption rate of the substrate, t_m is the mirror thickness, P_{mich} is the power between BS and ITM, γ_{coa} is the absorption of the coating, P_{circ} is the intra-cavity power of the main arm, and K_{rad} is the radiation energy from outside. β_{sub} is measured to be 30 ppm/cm [3], but that value has a large error and the requirement was 50 ppm/cm. We assumed the absorption rate of 80 ppm/cm before 2016, but here $\beta_{\text{sub}} = 50$ ppm/cm is assumed because it is likely to meet the requirement. $t_m = 15$ cm is designed and $\gamma_{\rm coa} = 0.5$ ppm is assumed. Transmittance of ITM is designed to be $T_{\rm ITM} = 0.004$, so the intra-cavity power is $P_{\rm circ} = 4P_{\rm mich}/T_{\rm ITM} = 10^3 P_{\rm mich}$. Radiation from an opened window for beam is dominant in that from outside and estimated to be $K_{\rm rad} = 50$ mW [11]. With these parameters laser power between BS and ITM can be increased to be $P_{\rm mich} = 337$ W in order to keep the mirror temperature of 22 K. Therefore the maximum laser power at BS is twice than that, $P_{\rm BS} = 674$ W.

References

- [1] Y. Aso *et al.* Phys. Rev. D 88, 043007 (2013).
- [2] T. Uchiyama *et al.* Phys. Lett. A **261**, 5 (1999).
- [3] E. Hirose, JGW-G1706429-v1 (2017).
- [4] K. Somiya *et al.* Class. Quantum Grav. **29**, 124007 (2012).
- [5] E. Hirose *et al.* Phys. Rev. D **90**, 102004 (2014).
- [6] W. Duffy *et al.* Cryogenics **32**, 1121 (1992).
- [7] R. Newman et al. Phil. Trans. R. Soc. A 372, 20140025 (2014).
- [8] R. Kumar JGW-G1503717-v1 (2015).
- [9] D. Chen Ph. D. thesis JGW-P1605622-v1 (2016).
- [10] A. Khalaidovski et al. Class. Quantum Grav. **31**, 105004 (2014).
- [11] Y. Sakakibara *et al.* Class. Quantum Grav. **31**, 224003 (2014).