Gravitational Wave Physics and Astronomy Workshop Fast localization (Annecy, May 30th – June 2nd, 2017) with a hierarchical network of gravitational wave detectors



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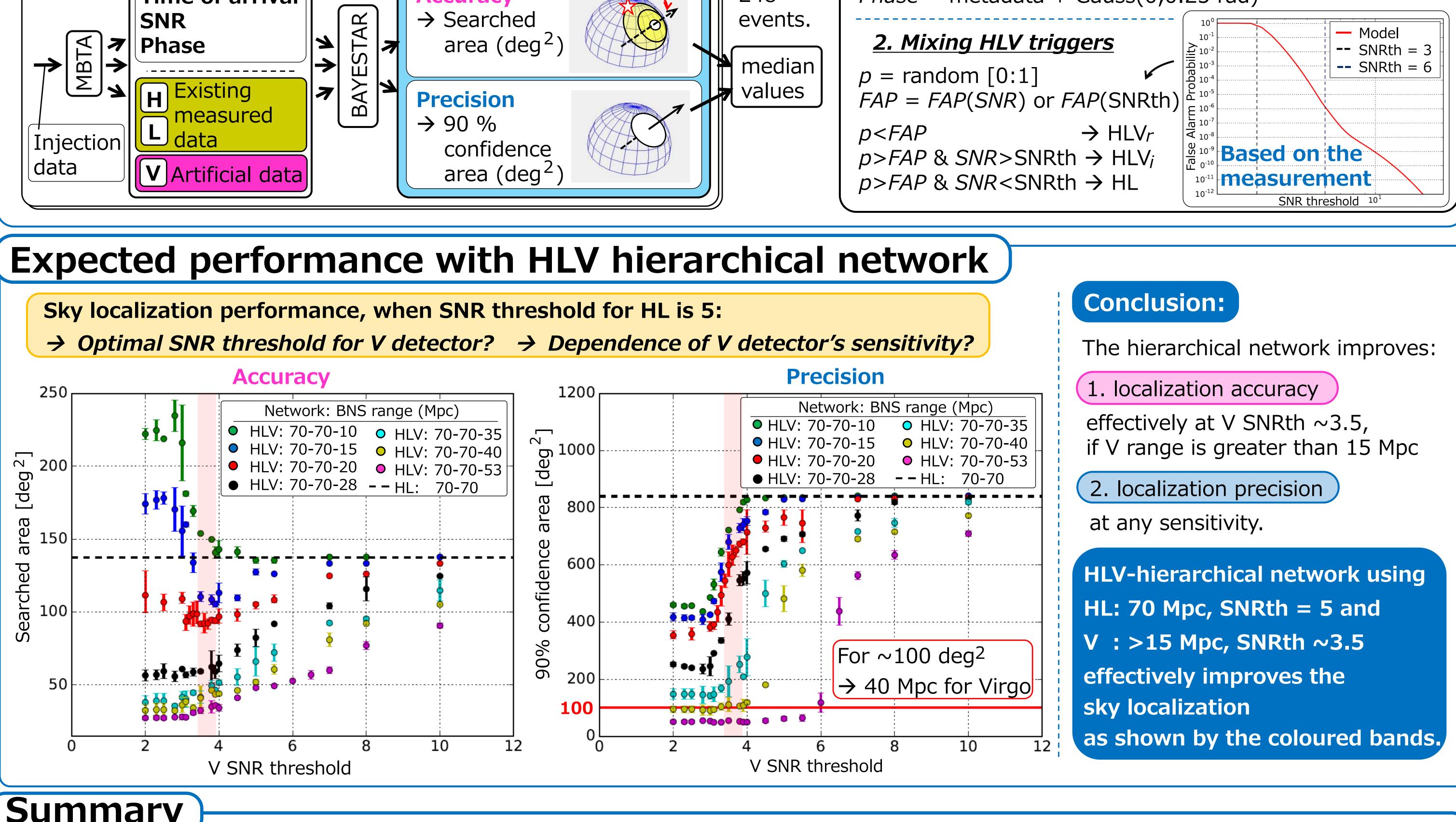
Introduction

We present expected fast sky localisation of coalescing binaries with a hierarchical search using three gravitational wave (GW) detectors, HLV (Hanford/Livingston/Virgo).

A hierarchical search can be used with a network of GW detectors with varying sensitivities, and is aimed at making effective use of the least sensitive detector's information. Here we demonstrate the sky localisation using a hierarchical search with the two higher sensitivity LIGO detectors and the less sensitive Virgo detector, using simulated signals.

Hierarchical network How to analyze: threshold Higher sensitivity detectors sub network detects For precise source localization: Lower sensitivity Higher sensitivity One template \(\frac{\pi}{2} \) Triple (or more) coincidences candidate event. Virgo (V) **LIGO** Hanford (H) At the beginning: **Detectors with different sensitivity** threshold Less sensitive detectors For getting more coincidences: are added into network Set a lower threshold, as long as 1. with lower SNR threshold not too many background triggers 2. using same parameters & Higher sensitivity Lower sensitivity 3. a small window around time Analyze hierarchically! **LIGO** Livingston (L) KAGRA(K) of double coincidences. (At the beginning) → How does this approach improve the localization? Generating & mixing artificial Vtriggers **Assumptions:** SNR distribution 1. Generating V triggers

Calculation setup Higher sensitivity: HL \rightarrow 70 Mpc, Lower sensitivity: V (for 1.4–1.4 M_{\odot} BNS range) Model Measurement V_r : V trigger based on random parameters EM GW detectors) $|MBTA[1]| \rightarrow |BAYESTAR[2]| \rightarrow$ = random following telescopes $\mathsf{H}\,\mathsf{L}\,\mathsf{V}$ **Measured** measurement Signal Sky map Compact **Event info:** E 10-4 In O1 $Time = t_{H1} \text{ or } t_{L1}$ probability **B**inary • Time of arrival, + random [-35ms:35ms] etc. Coalescence Phase = random $[0:2\pi]$ V_i : V trigger based on injection parameters **Calculation main flow:** = metadata + Gauss(0,1) Localization performance Inputs Time = metadata + Gauss(0,0.66 ms* $\frac{6}{SNR}$) Using 248 Phase = metadata + Gauss(0,0.25 rad)Time of arrival Accuracy **SNR** → Searched events. Model 2. Mixing HLV triggers area (deg²) Phase **--** SNRth = 3 median -- SNRth = 6 p = random [0:1]**H** Existing values FAP = FAP(SNR) or FAP(SNRth)**Precision** measured **→** 90 % p<FAP \rightarrow HLV_r **L** data Injection confidence Based on the p>FAP & SNR>SNRth → HLV_i data area (deg²) V Artificial data p>FAP & SNR<SNRth → HL 10-11 measurement SNR threshold 101



Summary

- 1. We investigated the expected fast localization performance with a hierarchical network using HLV.
- 2. We demonstrated that the hierarchical network effectively improved the accuracy & precision when V threshold is set to \sim 3.5, if BNS range of V detector is greater than 15 Mpc.
- 3. The hierarchical search will be most useful when adding new detectors, which are less sensitive as they are undergoing commissioning, to the network.

Future work:

- 1. Investigate the localization with HLVK hierarchical network
- 2. Implement in online analysis

References: [1] T. Adams et. al., Class. Quant. Grav. 33 (2016) [2] L. P. Singer, L. R. Price, Phys. Rev. D **93**, 024013 (2016)