KABRA Optomechanical Instability in the KAGRA Gravitational Wave Telescope

Yuta Michimura for the KAGRA Collaboration Department of Physics, University of Tokyo Email: michimura@granite.phys.s.u-tokyo.ac.jp

[3] L. Barsotti et al., Class. Quatum Grav. 27, 084026 (2010)

KAGRA drawings and photos are taken from JGW-G1503311 by S. Kawamura and JGW-G1402288 by M. Ando

[5] M. Evans et al., https://github.com/Optickle/Optickle

[4] E. Morrison et al., Appl. Opt. 33, 5041 (1994)

Mathematica 3D rigid body suspension model is made by T. Sekiguchi

http://www.astroarts.co.jp/news/2013/12/04binary_bh/

Drawings of GW sources are from

http://youtu.be/g8s81MzzJ5c

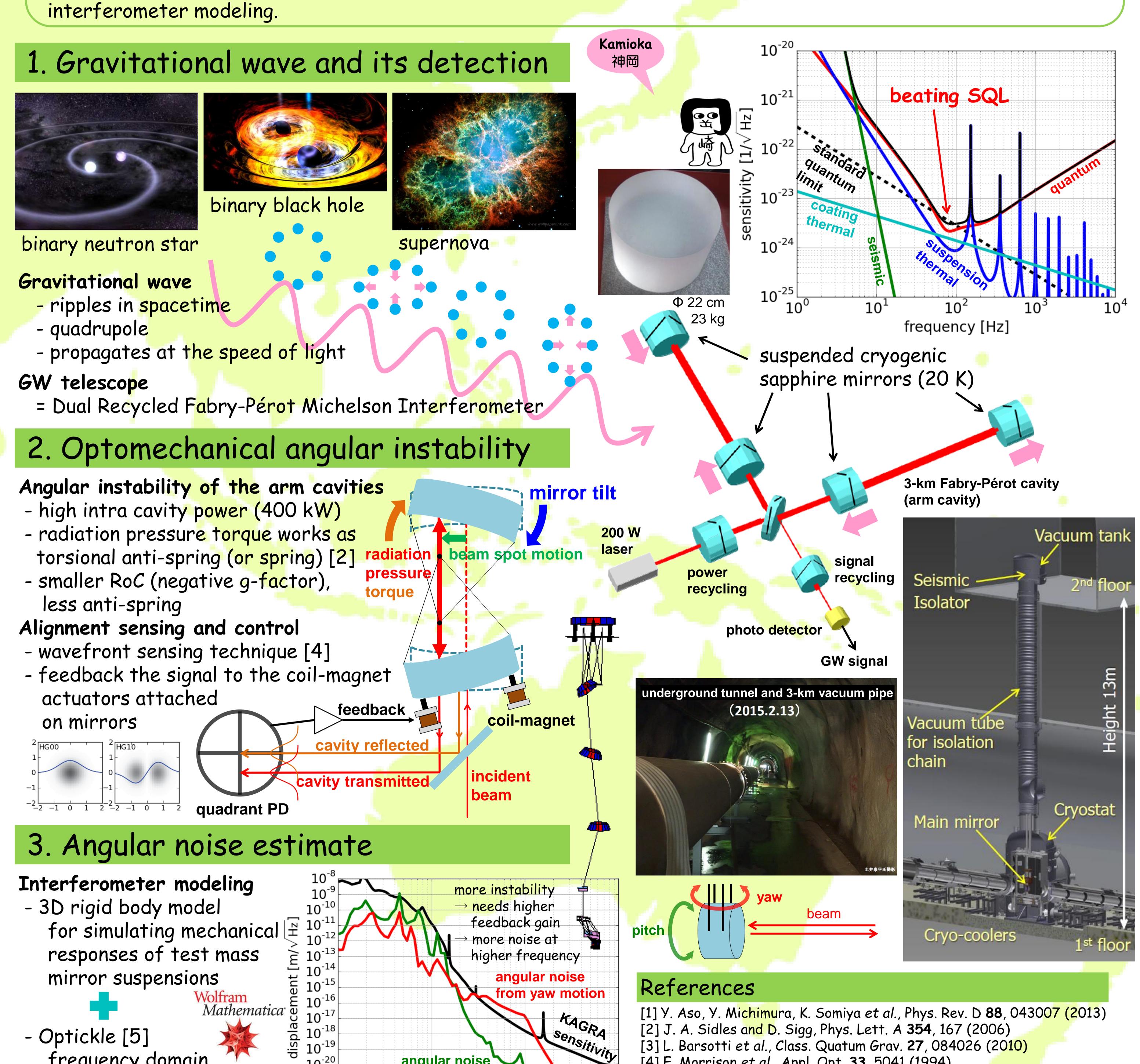
http://en.wikipedia.org/wiki/Supernova

Abstract

KAGRA is a 3-km interferometric gravitational wave telescope which started construction in 2010 at Kamioka, Japan [1]. We reduce seismic noise by constructing the interferometer in the quiet underground site, and we reduce thermal noise by cooling down the test mass mirrors to 20 K. These advanced technologies help KAGRA detect gravitational waves from binary neutron stars 150 Mpc away.

In order to achieve such a high sensitivity, longitudinal and angular motions of the mirrors must be finely controlled. However, the alignment control will be one of the most challenging issue because of the optomechanical angular instability of the arm cavities [2,3].

Here, we present our interferometer design to reduce this instability, and show angular noise estimate from the



10^{-19]}

10⁻²⁰

10⁻²¹

10⁻¹

angular noise

10°

from pitch motion

frequency [Hz]

10²

frequency domain

simulation tool

MATLAB

quantum-optomechanics