KAGRA Detector Characterization

Kazuhiro Hayama (Osaka City U) On behalf of The KAGRA Detector Characterization

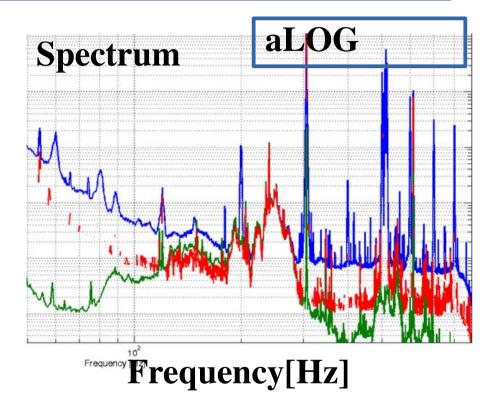
Human Resources

Hayama(OCU), U. Tokyo: Araya, Itoh,Miyakawa,Ono,Uchiyama OCU : Asano, Kanda,Miyamoto,Yamamoto, Yokozawa, Yuzurihara ISM : Mano Osaka U: Narikawa, Ueno

Development of KAGRA

Design→**Make**→**Installation**→**Design Sensitivity**

- Difficult to understand the kmscale instrument at a glance
- Important to have diagnostics system to know its behavior via ~1000 PEM and instrumental channels.
 - → DetChar group is developing the system and diagnostics tools



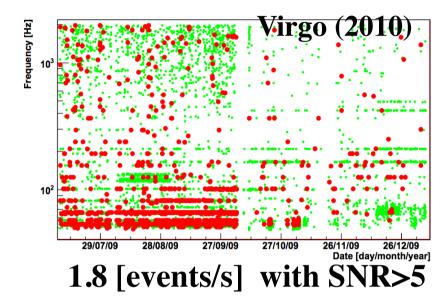
Experts finds noise sources from large degree of freedom

→ System localizes the sources and reduce the DoG

Observation of Gravitational Waves

Evaluation of data quality

- Is KAGRA working properly? What about Environmental situation?
- Can we do science using the data taken today?
- The triggered signal, which is above the detection threshold statistically, is really gravitational wave? Or just artificial noise?



	Line categories	Number of identified lines	
	Violin modes	127	
Intrinsic lines	Mechanical resonances	26	
	Calibration and control	32	
Noise lines	Power line and harmonics	40	
	Vibration	24	
	Magnetic	-	
	Digital	73	
	Sidebands	640	

960/1390 identified

KAGRA detector characterization

- To select tools for detector diagnostics, evaluation of data quality which has been useful for TAMA、LIGO、Virgo. We are developing the tools from scratch in order to do tuning for KAGRA deeply.
- So far we still have unidentified signals. We try to develop tools to understand such unidentified noise and make contribution to the GW committee.
- At the weekly meeting, we review aLOG and study what kind of tools are useful, what they need at the aLIGO commissioning.
- User friendly GUI interface
- Web-based summary page

Primary Projects

DetChar Projects

- To maintain Diagnostics Test Tool
-] Detchar GUI
- **Glitch Monitor**
- Detchar web page
- Line Monitor
 - correlation finder
- Noise Modeling
- **Rayleigh Monitor**
- Noise Floor Monitoring
- Range Monitor (Inspiral, Ringdown, Insp-Merger-Ringdown, Stochastic)
- Noise Budget
- Health Monitor
- 📑 Data base
- **Quality flag**

Special Projects

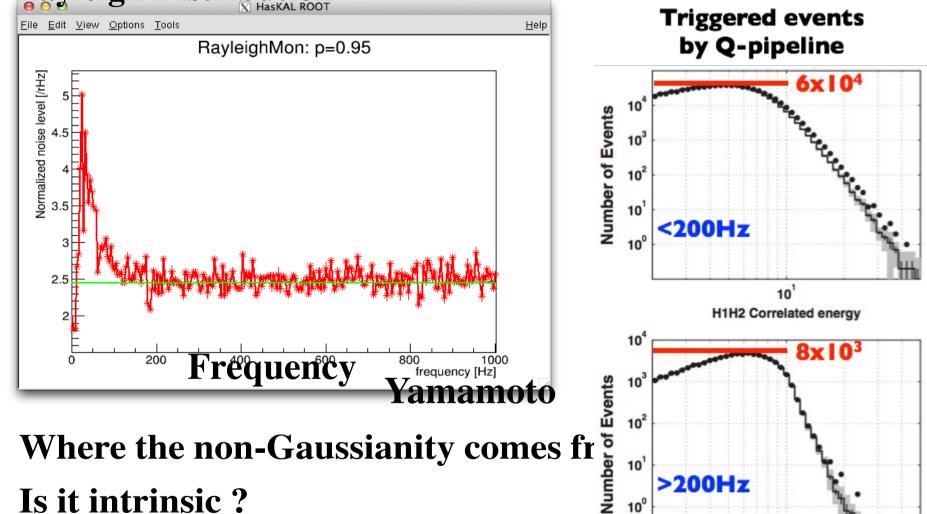
- Globally correlated mag noise
- 📄 Violin mode
- Multi-Channel Analysis (with Korea detchar, Mano)
-] Detchar shift plan
- Newtonian Noise
 - in progress
 - in slowly progress

Noise Characterization at the KAGRA site



(Non-/) Stationary non-Gaussian Noise feature

Rayleigh Distribution



>200Hz

10¹ H1H2 Correlated energy

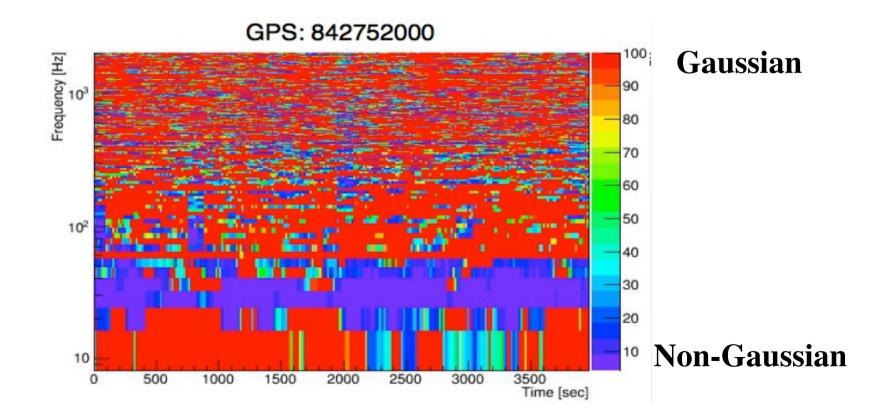
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Is it intrinsic?

Non-Gaussian Noise Modeling Yamamoto

- Noise Modeling using Student-t distribution.
- Characterizing non-Gaussianity using 1parameter.

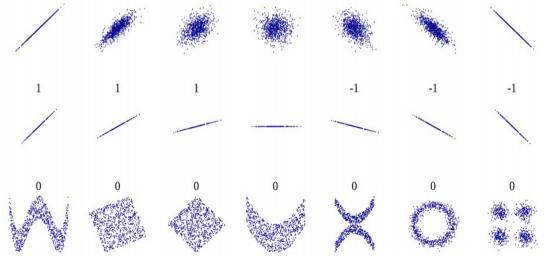


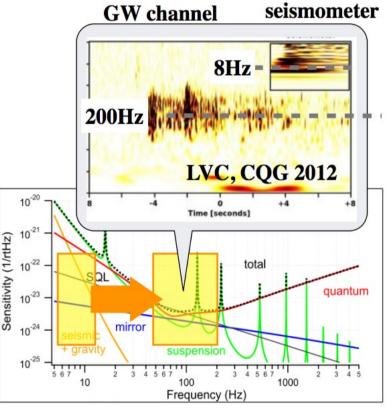
Correlation Finder

Yuzurihara, Hayama, Mano

Some noise sources make correlated noise in multiple channels

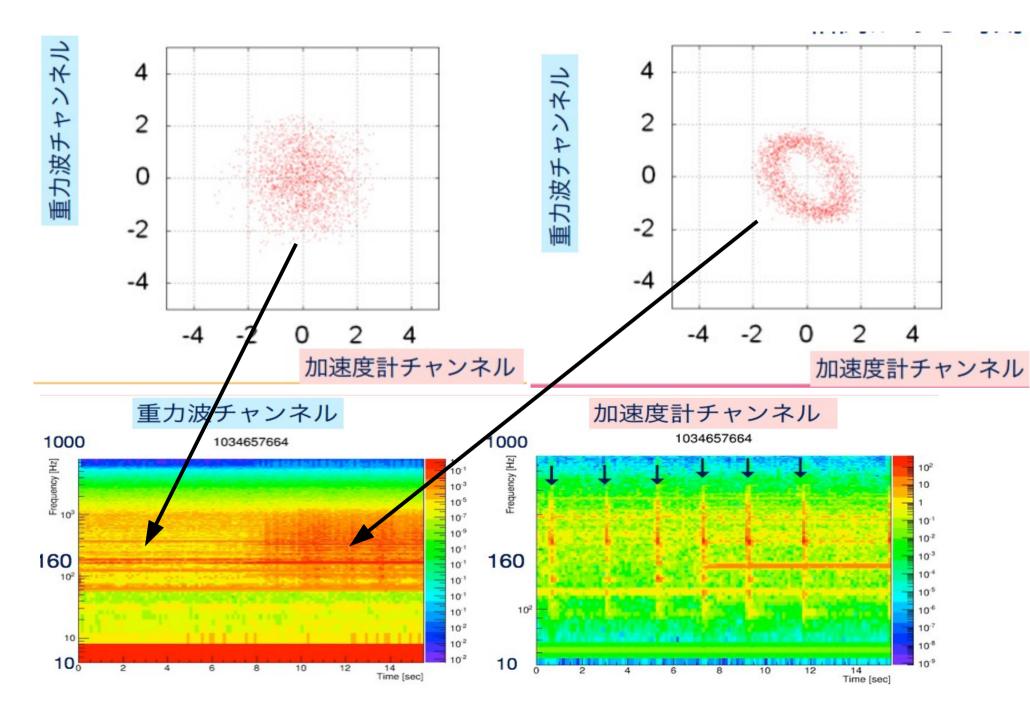
- Finding Linear and non-Linear correlation from enormous channe
- Pearson : Find Linear correlation
- MIC : Find non-Linear correlation





Sensitivity curve of KAGRA

CLIO Case (2012 Sep)



Correlation Heat Map

	ch1	^{///} ers ch2 de	ch3	ch4	ch5	ch6
51	1.00	-0.41	-0.51	0.75	0.80	-0.75
ch2	-0.41	1.00	0.05	-0.46	-0.57	0.42
ch3	-0.51	0.05	1.00	-0.33	-0.35	0.34
ch4	0.75	-0.46	-0.33	1.00	0.82	-0.98
ch5	0.80	-0.57	-0.36	0.82	1.00	-0.82
ch6	-0.75	0.43	0.34	-0.98	-0.82	1.00

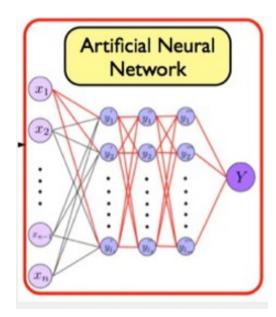
Yuzurihara

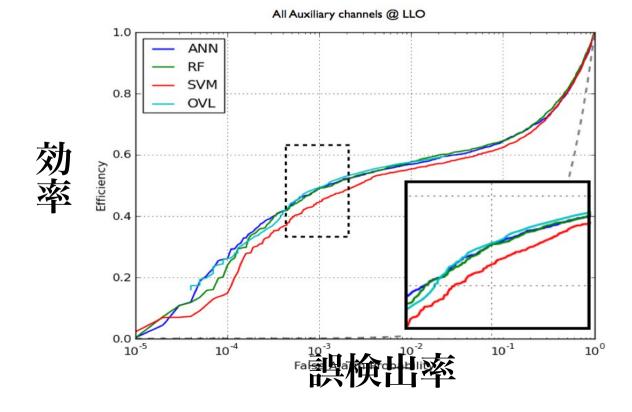
Line Characterization

Asano, Ueno Line Tracking(Ueno) Line Removal (Asano) Time variation of amplitudes Q=10^3 Q=10^4 ~ 345 Hz ~370 Hz Q=10^5 Q=10^6 ~403 Hz Noise spectral density [1//Hz] 10⁻²⁰ 10⁻²¹ 10⁻²² LIGO-Hanford LIGO-Livingston Virgo VSR2 Virgo VSR3 0-22 10³ 10^{2} 10 LIGO,Virgo Frequency [Hz]

Localizing Noise sources using Multichannel Analysis KGWG(Korea)

 Neural Network based method (KGWG)





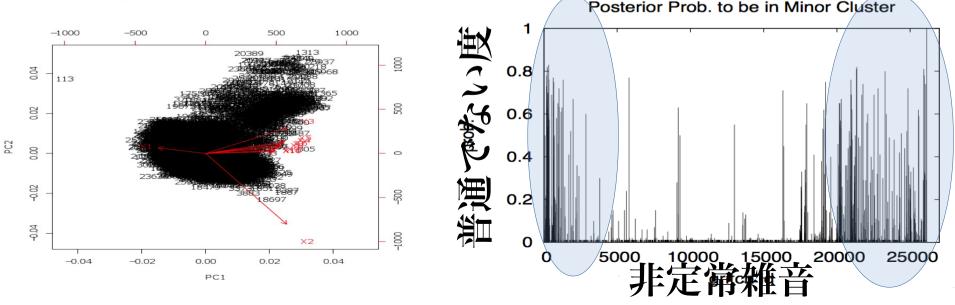
Localizing Noise sources using Multichannel Analysis

Bayesian Non-parametric Clustering Mano(ISM)

- Non-Supervised Machine Learning
- To show "uncommon" noise events

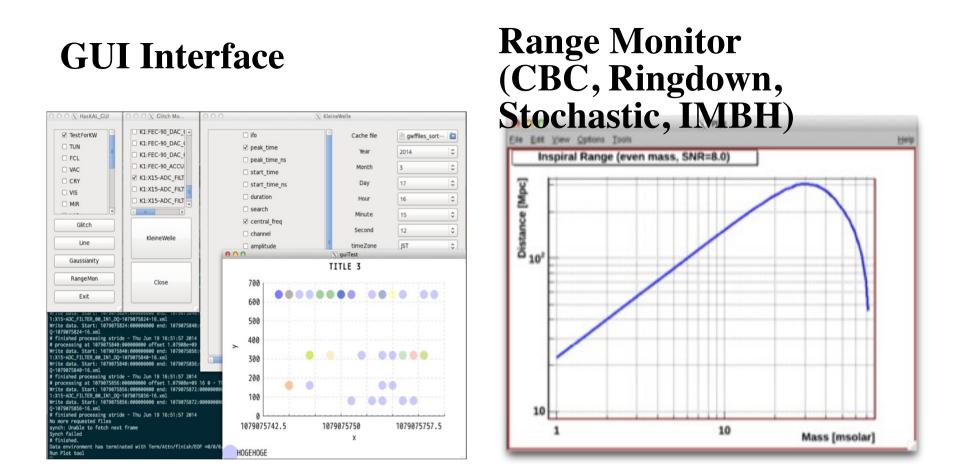
TAMA300のデータから採集した 2万6000の非定常雑音カタログを クラスタリング





HasKAL Detector Characterization Analysis Tools

upload MBLT items		
asano0622 authored 9 days ag	0	latest commit 9fa358144c
DetectorUtils	working around injection	19 days ag
ExternalUtils	Mine.hs updated	14 days ag
FrameUtils	small change	11 days ag
GUI_Utils	changed GUI_Utils for plot tool update	11 days ag
LineUtils/LineRemoval	upload MBLT items	9 days ag
Misc	move haskalOpt to Environment module	2 months ag
MonitorUtils	change plot tool of RayleighMon from Chart to HROOT	14 days ag
PlotUtils	modified plot tool	11 days ag
SearchUtils	added SearchUtils	22 days ag
SignalProcessingUtils	minor update	13 days ag
SimulationUtils	add injection function which uses bang method for memory saving	18 days ag
SpectrumUtils	minor change of DetectorSensitivity	11 days ag
StatisticsUtils	change module name	13 days ag
TimeUtils	change function fromGPS to deformatGPS	19 days ag
WaveUtils	add dropWaveData, takeWaveData	13 days ag
DetectorUtils.hs	added module of modules	2 months ag
TimeUtils.hs	added, module-setting module	or characterization
WaveUtils.hs	ttps://github.com/gw-analysis/detect	OI-CITATACLETIZATION 19 days ag



Developers: Asano, Hayama, Itoh, Mano, Ono, Ueno, Yamamoto, Yokozawa, Yuzurihara,

And so on ...

Plan

- In 2014、GUI tools、Web-based summary tools will be version1
- Mar. in 2015 Participation of PEM monitor operations by GIF
- \sim 2015 Dec, improvement, modification
- IKAGRA operation

