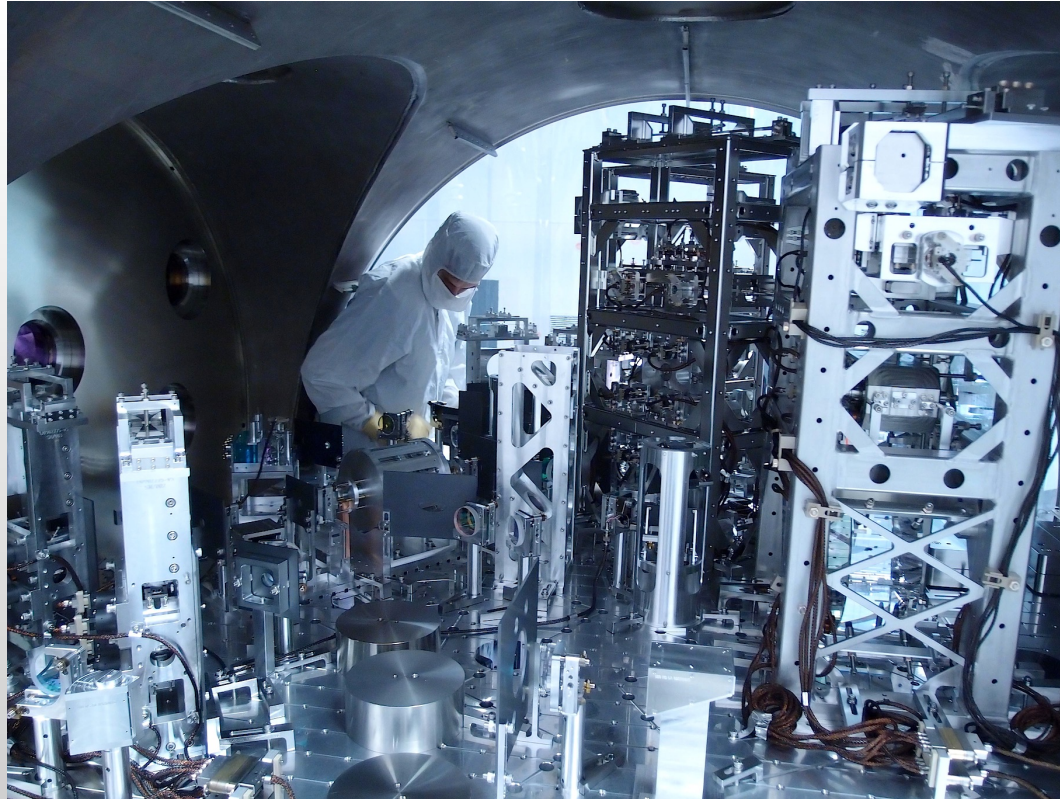


News from LIGO - October 2013



Koji Arai

LIGO Project, California Institute of Technology

Domestic KAGRA Collaboration Meeting (Oct. 9th, 2013)

JGW-G1301893

LIGO Livingston

Dual-recycled Michelson Interferometer (DRMI)

- All interferometer components installed at vertex.
- DRMI locked, PRMI locked with DC Readout (OMC)

Plan:

- End test masses now being installed
- Installation completion in Feb 2014, 2hrs lock by Oct 2014

LIGO Hanford

Half-interferometer (HIFO-Y) completed

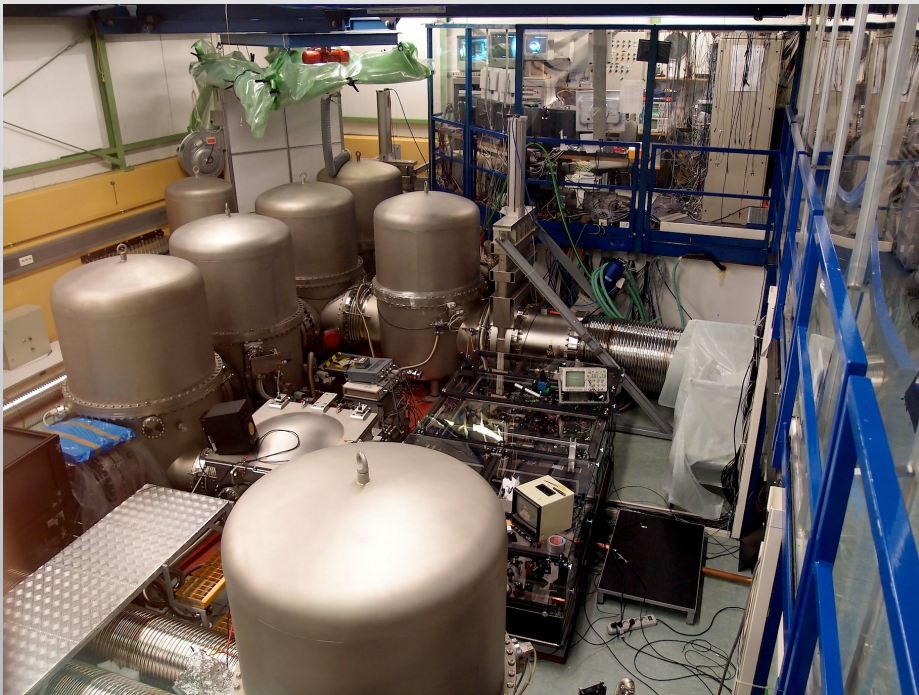
- First arm cavity test with the main IR beam
- Demonstration of Arm Length Stabilization (green locking)

Plan:

- To be followed by HIFO-X, HIFO-XY
- DRMI in May 2014
- Installation completion in July 2014, 2hrs lock by Dec 2014

GEO-HF

- Astrowatch, >1year data with 2~2.5dB squeezing.
- Long-term operation with the squeezed light (~3yrs at GEO)
- Steady improvement with patience
 - Power increase / Thermal Compensation / Scattered light
 - OMC control / Squeezing control



Advanced LIGO Output Mode Cleaner (OMC)

- An optical cavity to remove unwanted optical fields from the interferometer output beam.
- The first OMC was built and tested at Caltech, and installed at LLO
- The design, fabrication, and test results

OMC ISC team

Rich Abbott¹, Koji Arai¹, Sam Barnum³, Peter Fritschel³, William Korth¹, Jeffrey Lewis¹, Charles Osthelder¹, Sam Waldman³

OMC SUS team

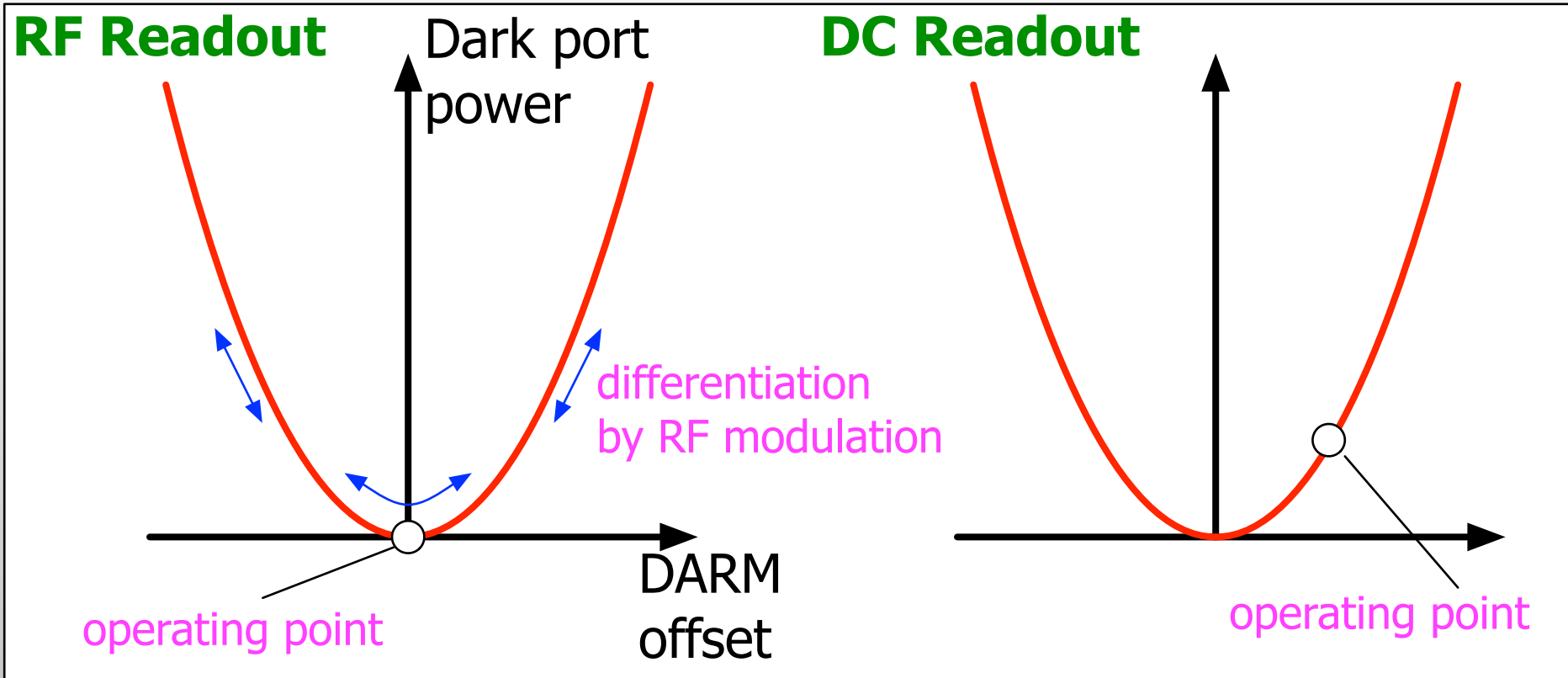
Stuart Aston², Jeffrey Bartlett⁴, Derek Bridges², Jeffrey Kissel³, Norna Robertson¹

LIGO Laboratory: California Institute of Technology¹, LIGO Livingston Observatory², Massachusetts Institute of Technology³, LIGO Hanford Observatory⁴

Mission of the OMC

DC Readout

aLIGO employs a DC readout scheme for sensing of GW signals



DC Readout is good:

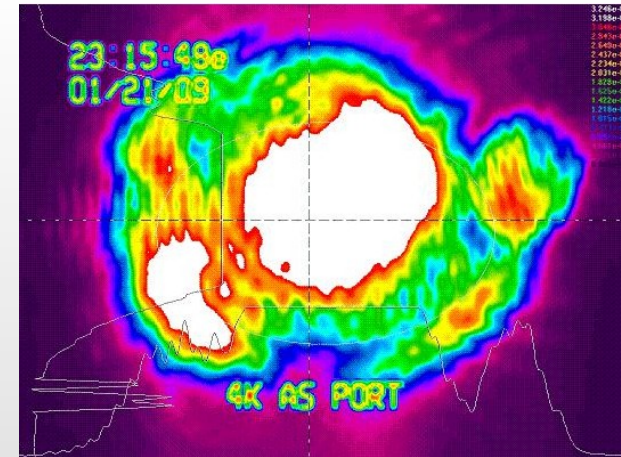
- removes nonstationary shot noise
- mitigates technical noises associated with the RF sidebands

Mission of the OMC

Enemies of the DC Readout

- Carrier HOMs (higher-order modes)
- RF modulation sidebands (any spacial modes)

do not contribute to the signal
and increase the shot noise



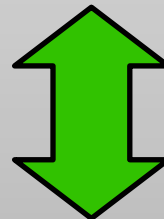
eLIGO AS port beam

Output mode cleaner

the idea is to use a short ($\sim 1\text{m}$) optical cavity for the filtering

Filtering performance

Trade-off



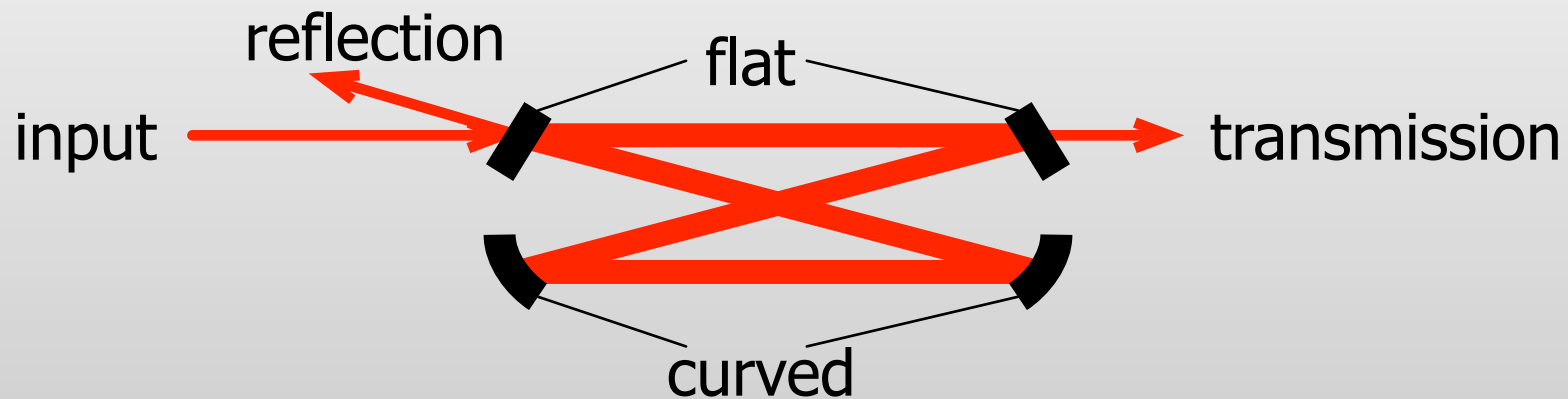
higher finesse

lower finesse

Signal transmittance

Basically eLIGO OMC design was followed

- Bowtie 4-mirror ring cavity
even mirrors => simpler HOM structure
ring cavity => less back scattering



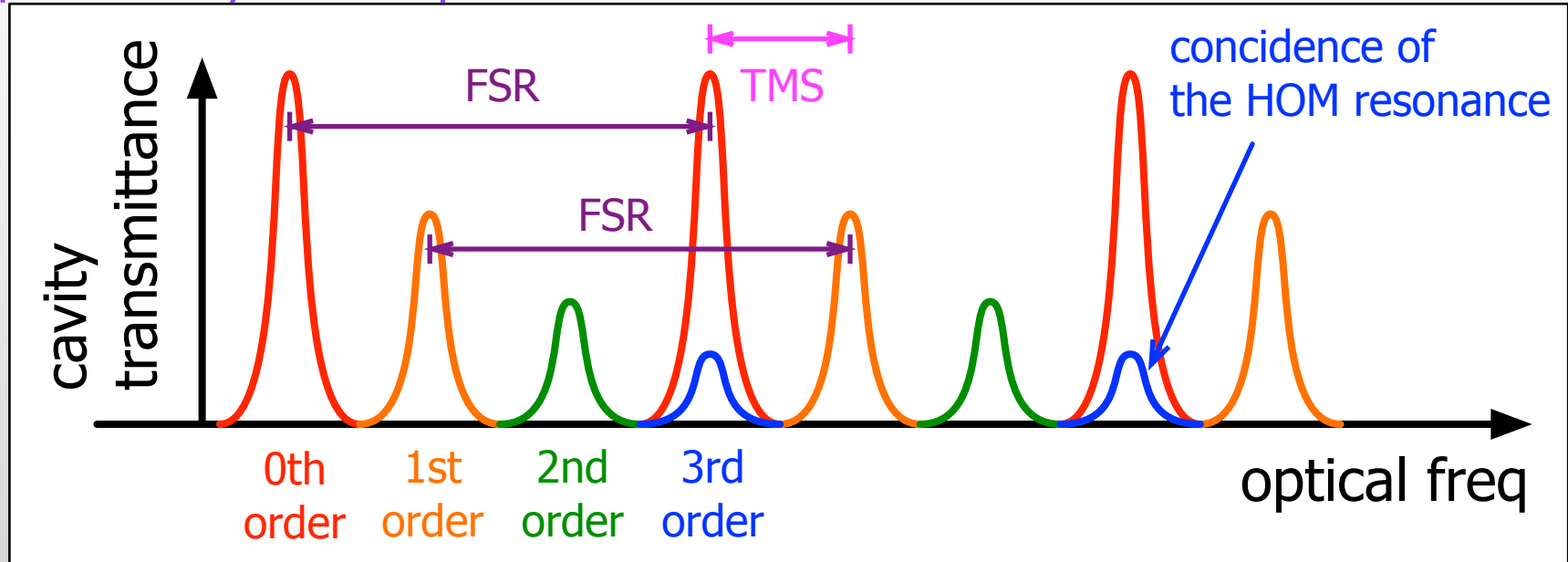
- Finesse: ~ 400 (for $\sim 98\%$ transmission)
Roundtrip length $\sim 1\text{m}$ (the breadboard size)
Curved mirror radius $\sim 2.5\text{m}$

Filtering Performance

Koji Arai
KAGRA meeting, Oct. 9, 2013
JGW-G1301893 P8

Important parameter: Transverse Mode Spacing (TMS)

An optical cavity has a repetitive resonant structure



If TMS/FSR is a rational number (m/n), n -th order HOMs get transmitted

MOREOVER: The vertical and horizontal modes have different TMSs due to astigmatism of the curved mirrors (i.e. non-zero incident angle)
=>The higher the mode number, the wider the resonance is.

TMS/FSR is dependent on the cavity geometry

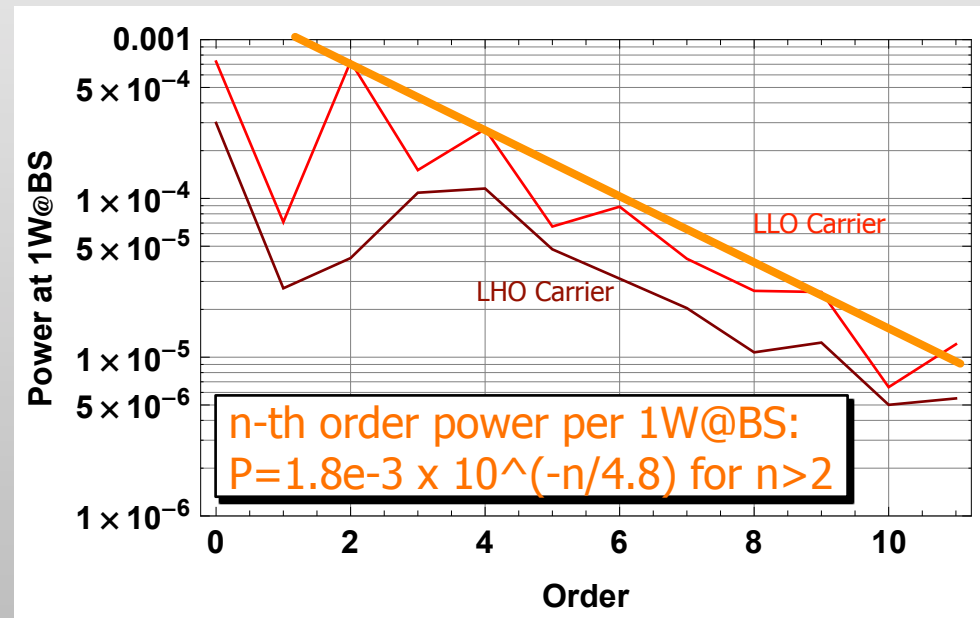
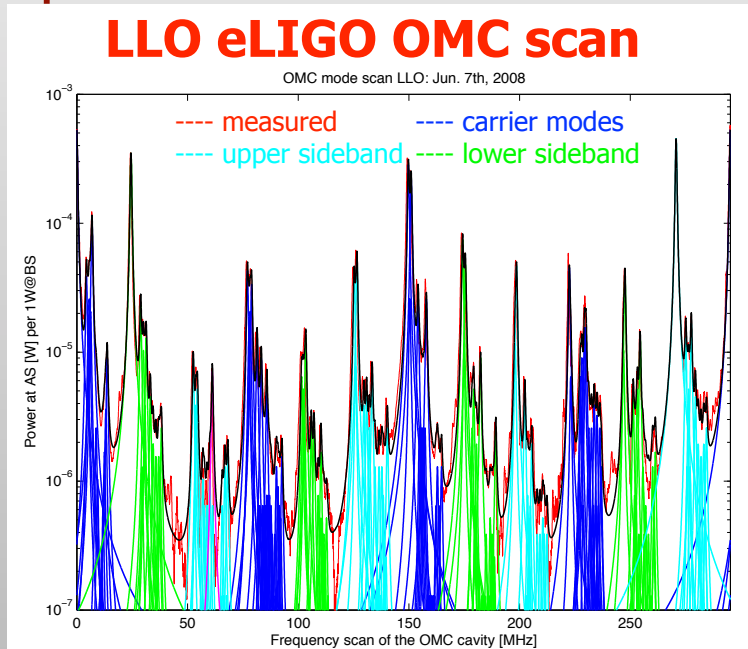
=>Careful adjustment of TMS/FSR is the key to avoid HOMs

Estimation of the filtering performance

Total transmitted power

$$= \sum (\text{power in each mode}) \times (\text{transmission of each mode})$$

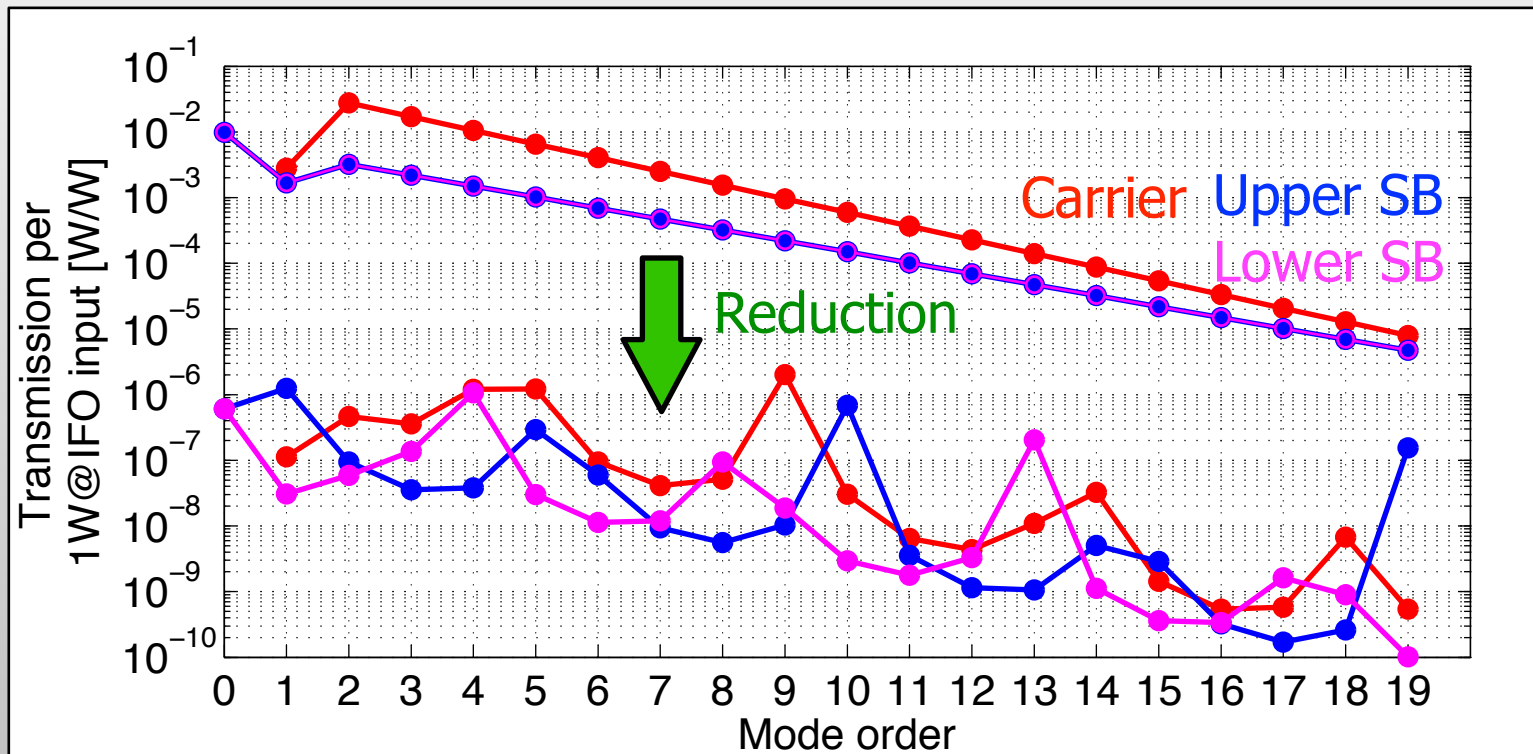
- Modeling of the interferometer output beam (details in G1201111)
power laws based on the eLIGO performance of the IFO optics



This wouldn't be a prediction, but have some usefulness, anyway

Estimated filtering performance

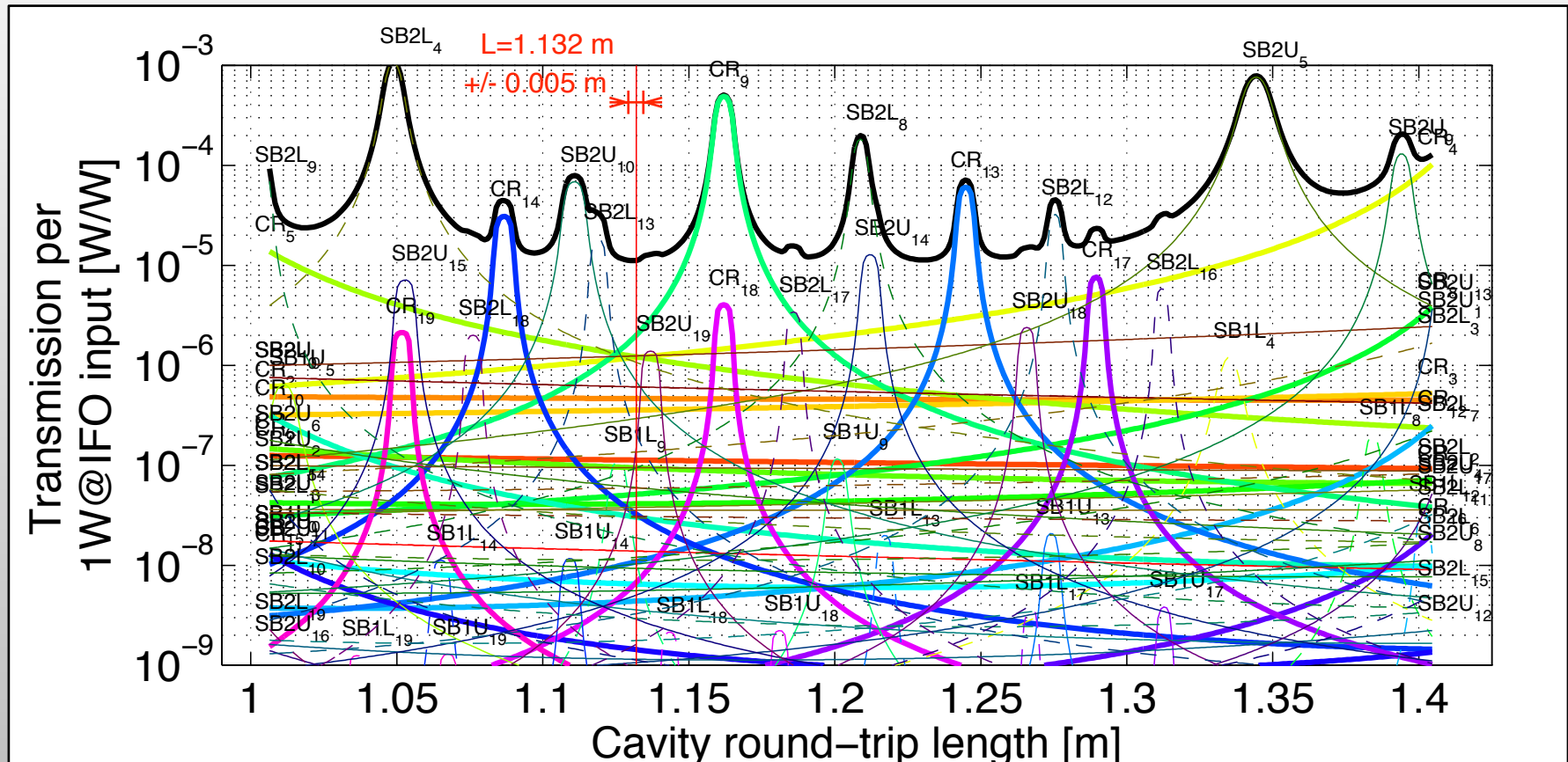
- Expected junk light power at the dark port (100W input = 4kW @BS)
~12W leakage => filtered down to 1mW. Well within the PD capability.
This could become better thanks to mode healing and better optics in aLIGO



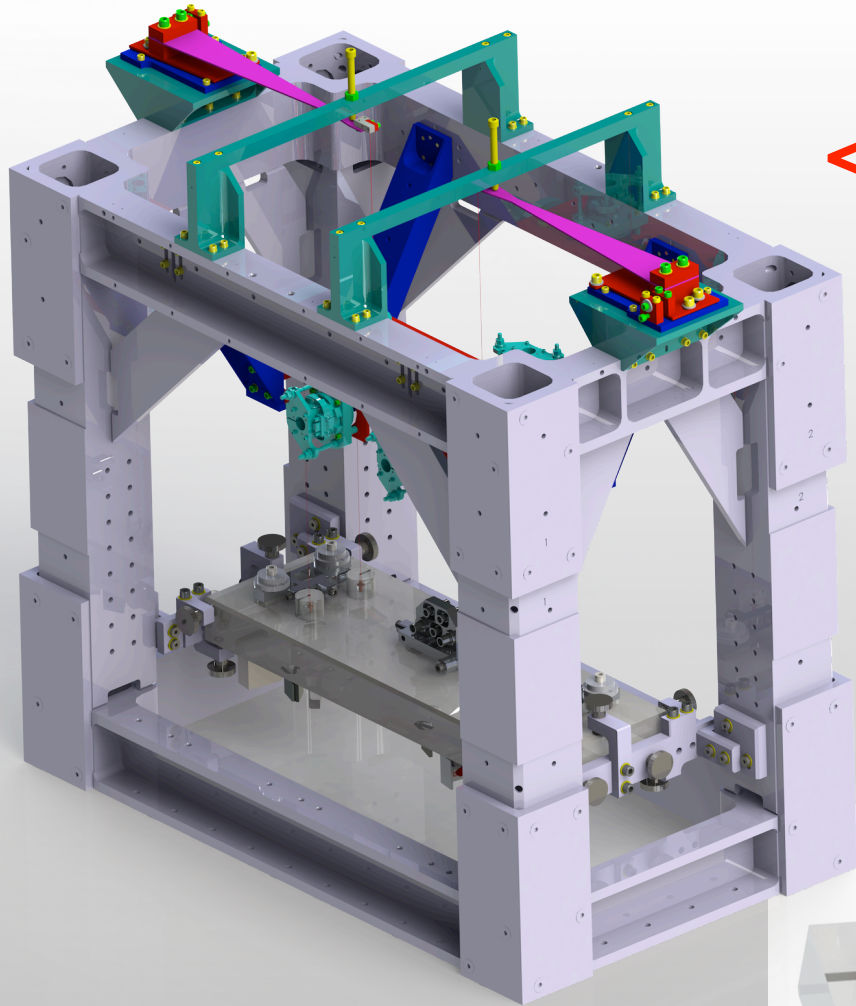
9MHz sidebands omitted in the plot due to small contribution

Parameter tolerance

- Cavity length tolerance: $L=1.132 \pm 0.005$ [m]
- Mirror RoC tolerance: $R=2.575 \pm 0.015$ [m]

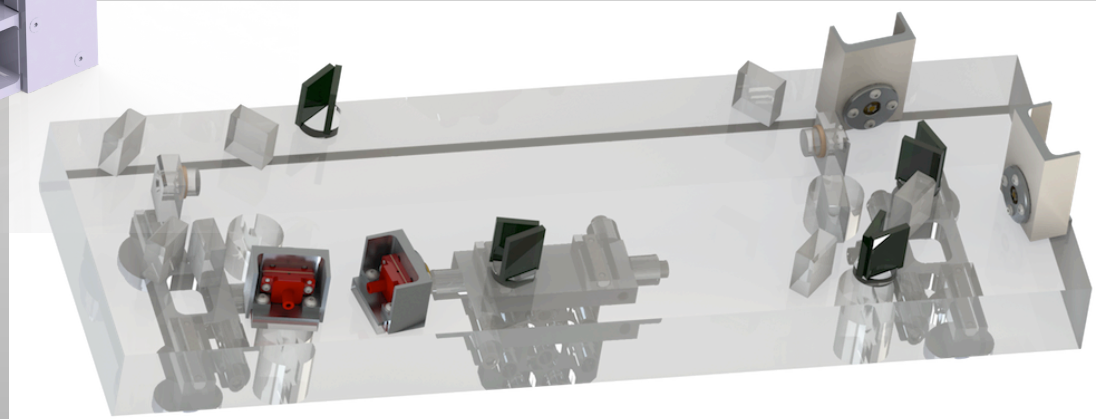


CR_n - carrier n-th mode, SB(1,2)(U,L)_n - sideband n-th mode,
SB1 - 9MHzSB, SB2 - 45MHz SB, U - upper SB, L - lower SB



<- OMC Suspension

v- OMC Breadboard



Comparison between eLIGO/aLIGO OMCs

eLIGO

aLIGO

Cavity design:

semi-monolithic

same

(for flexibility of the cavity parameters)

Breadboard material:

ULE

fused silica (dT of the order of 0.1K -> just 0.06um)

Mirrors

Glued on the back side of
the fused silica prisms

Glued on the front side of
the fused silica prisms (better access)

Actuators

1 PZT + 1 heater

Two PZTs (for redundancy)

Suspension

double pendulum
(actively damped)

same

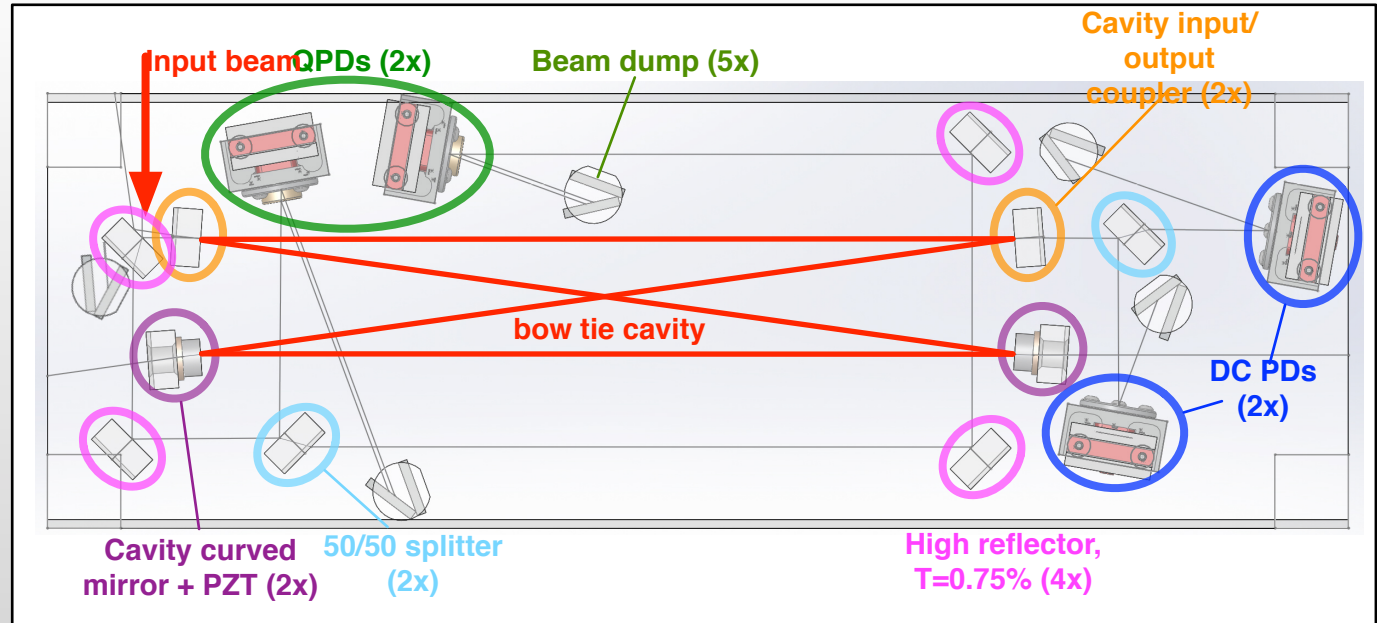
Wire clamps

screwed on the breadboard

glass wire bracket on the top side

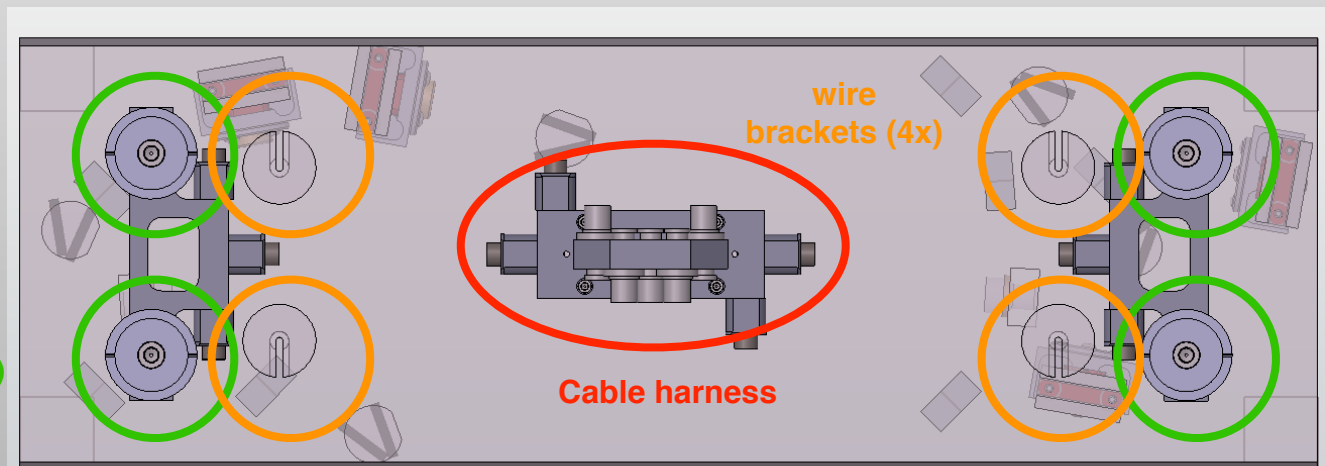
How the OMC breadboard looks like

Bottom side
(cavity side)



Top side
(cable side)

balance weight (4x)

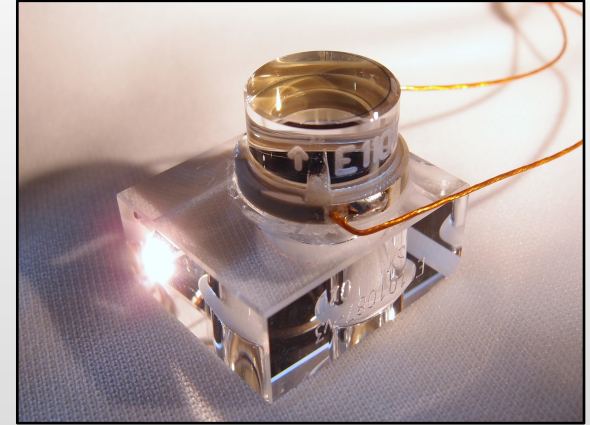
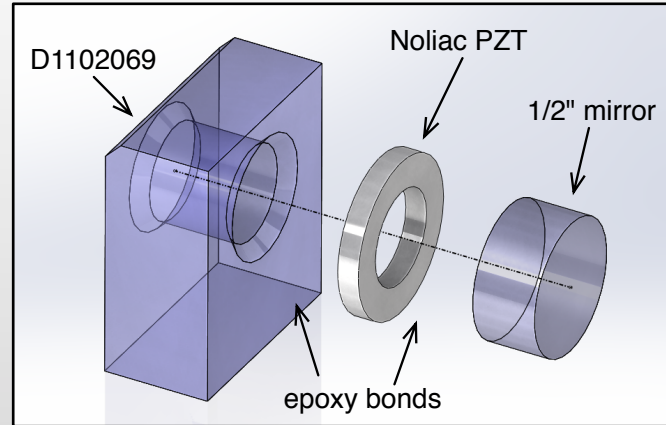


Procedure

- PZT & curved mirror sub-assembly
- Bottom (cavity) side gluing
- Top (suspension) side gluing
- Vacuum bake
- Optical test
- Cabling
- Shipping
- Installation

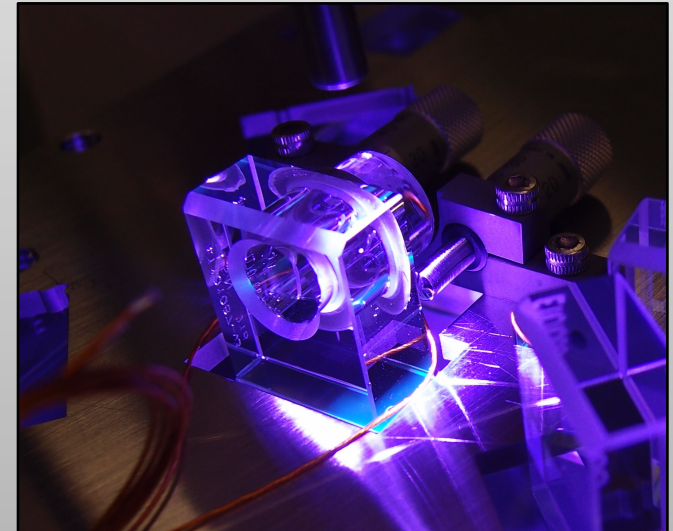
PZT-curved mirror subassembly

- PZT + 1/2" curved mirror



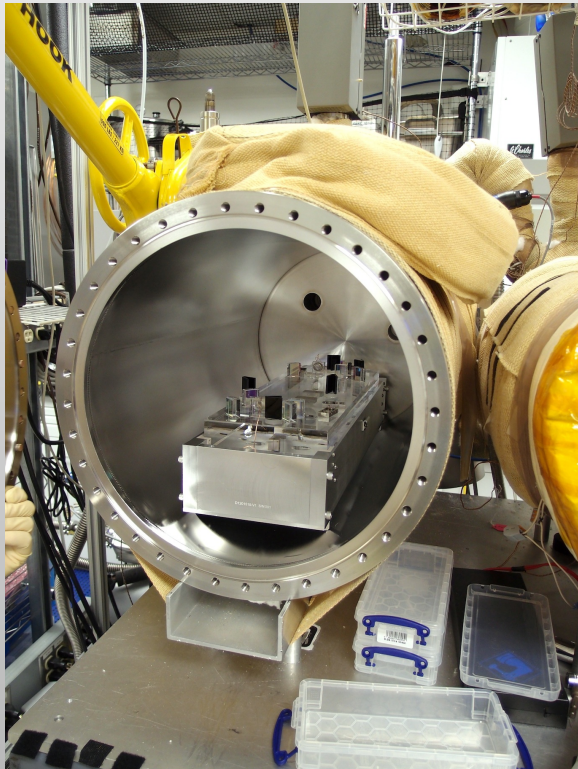
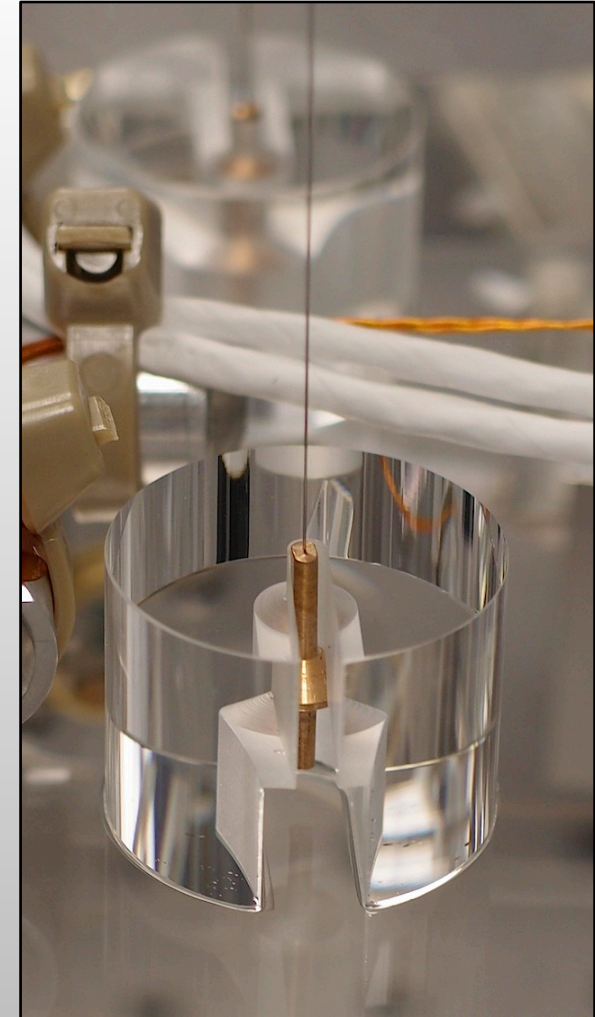
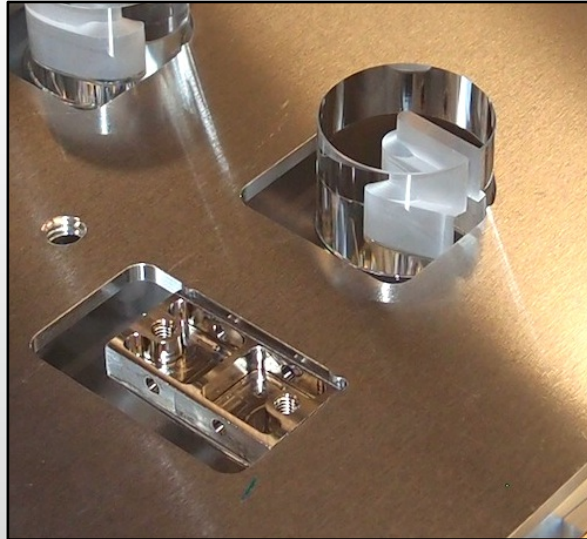
Cavity side gluing

- Template
- UV epoxy
- Bonding while the cavity is monitored



Top (cable) side gluing

- Suspension interface
- Mounting blocks

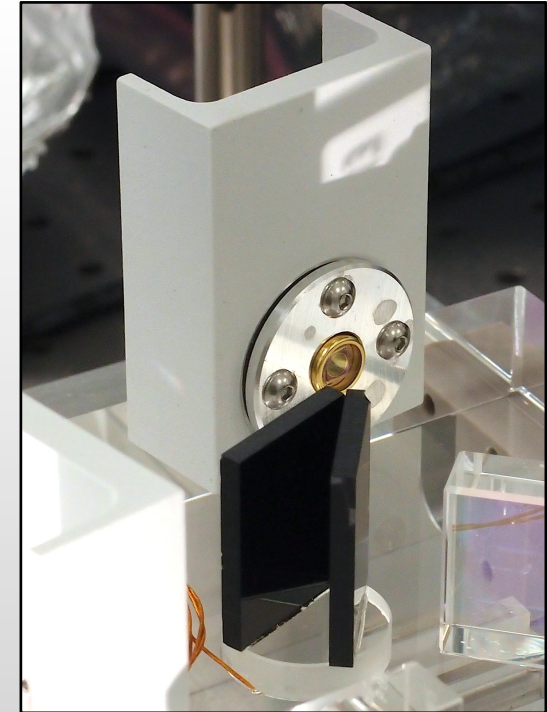
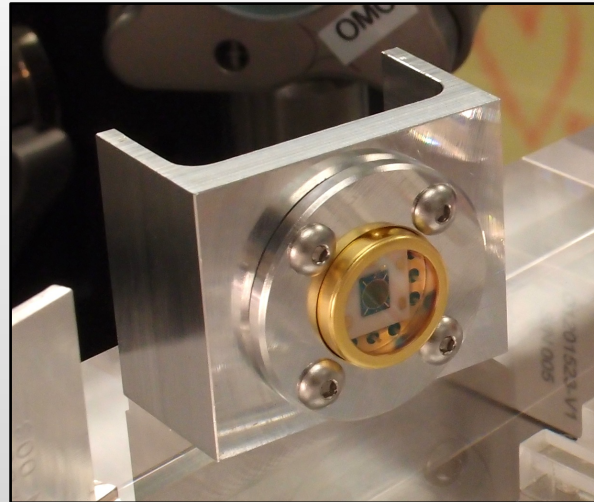


Vacuum Bake

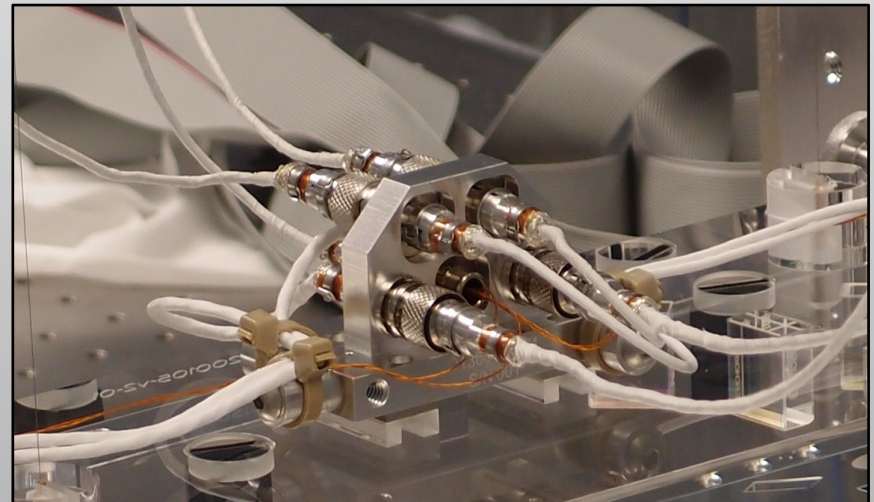
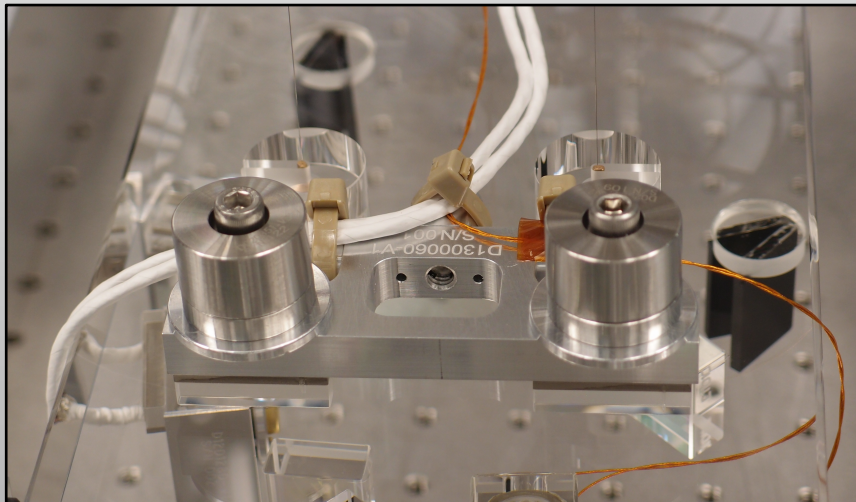
- reduce outgass
- completion of epoxy cure

Attaching peripheral components:

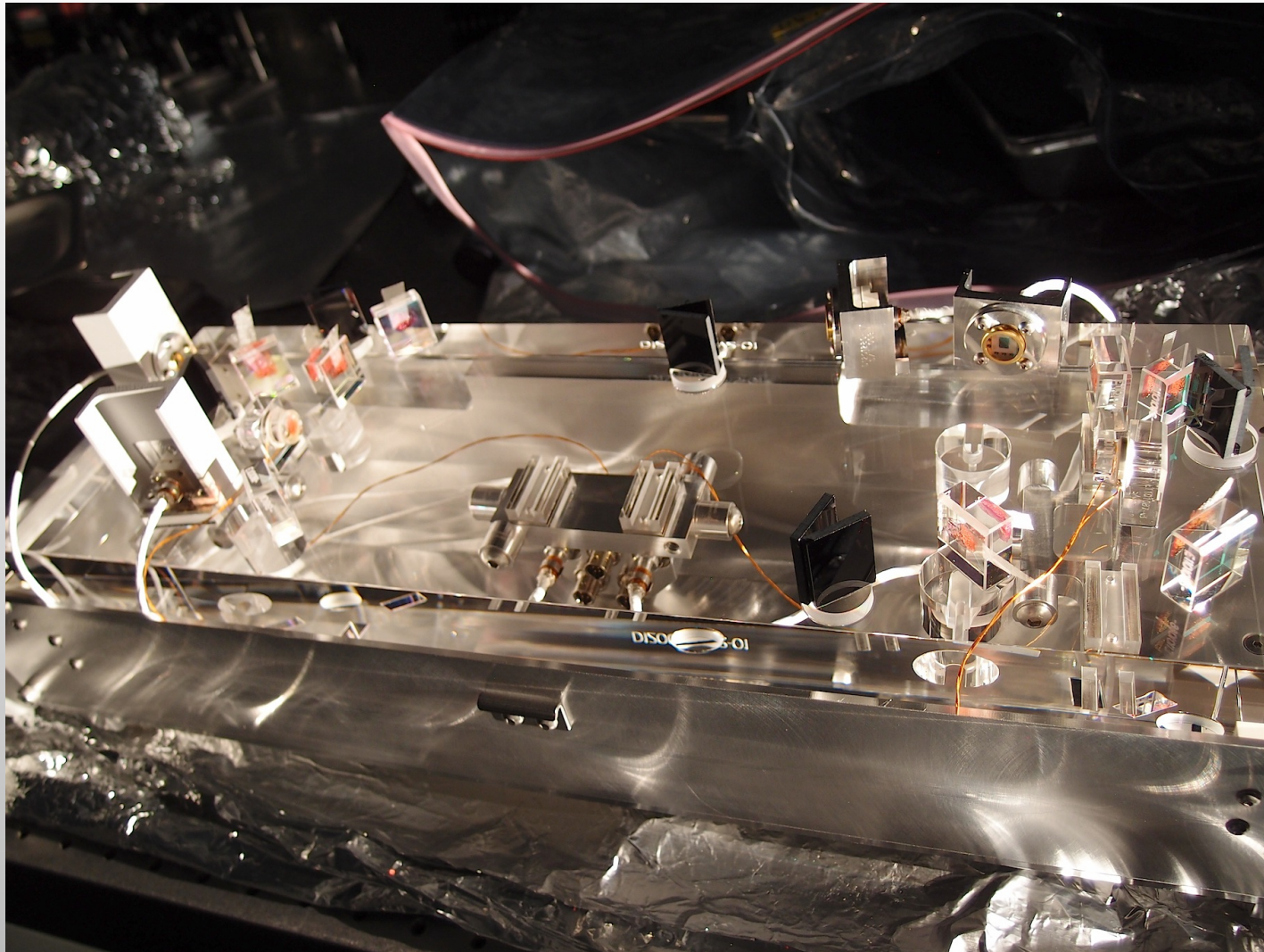
- DC photodiodes / QPDs



- Cables



OMC ready for the shipment



Power Budget

Estimated from the input power, transmitted power, visibility, and cavity finesse

Specification

Cavity transmission for TEM ₀₀ :	97.8 %	98.4 %
Curved mirror transmission:	42 ppm	50 ppm
Loss per bounce:	22.3 ppm	10 ppm
Loss per roundtrip:	173 ppm	140 ppm
PD Q.E.	92%	
Total thruput of TEM ₀₀	90%	(PD Q.E. = 92%)

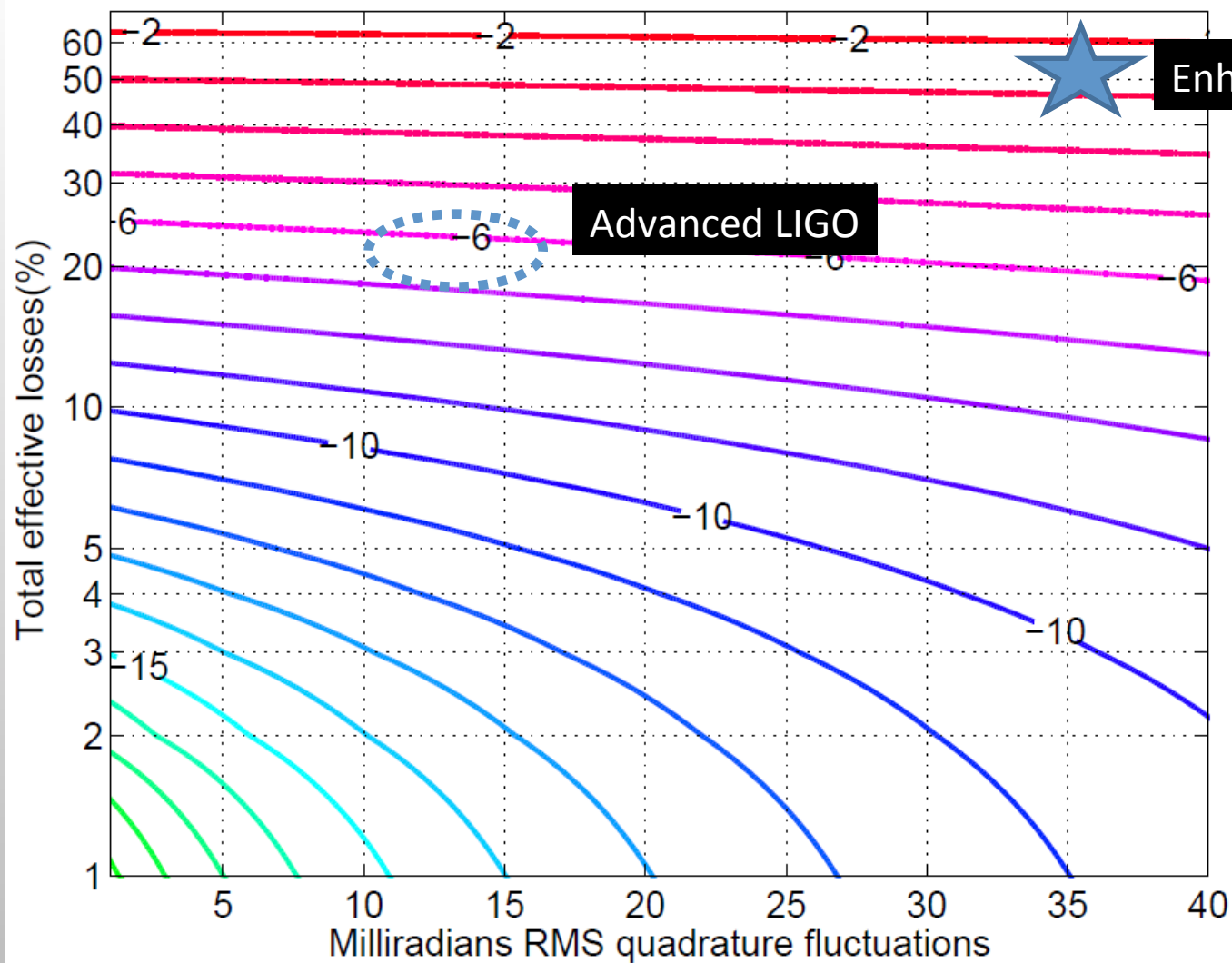
About 20% total loss allowed for 6dB squeezing.

A half of the budget already eaten up by the OMC. **(not nice)**

These PDs were previously (eLIGO) reported to have Q.E.>95%

Need further investigation (or replacement)

Optical testing



Cavity round-trip length

Cavity length: 1.131421 $\pm 3 \times 10^{-6}$ m

(Spec.: 1.132 ± 0.005 m)

Finesse: 403.79 ± 0.07

(Spec. 390)

Transverse Mode Spacing (TMS)

Pitch TMS/FSR: 0.218822 $\pm 1 \times 10^{-6}$

(Spec. 0.2188)

Yaw TMS/FSR: 0.219218 $\pm 1 \times 10^{-6}$

(Spec. 0.2194)

N. Uehara and K. Ueda, "Accurate measurement of the radius of curvature of a concave mirror and the power dependence in a high-finesse Fabry-Perot interferometer", Appl. Opt. 34, pp. 5611-5619 (1995).

TMS/FSR was successfully adjusted such that the HOMs were avoided until 32nd order.

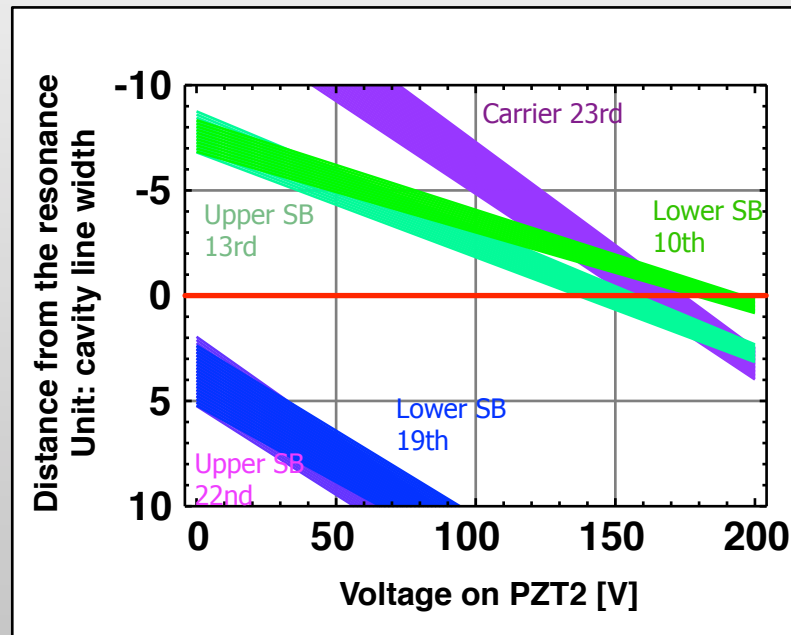
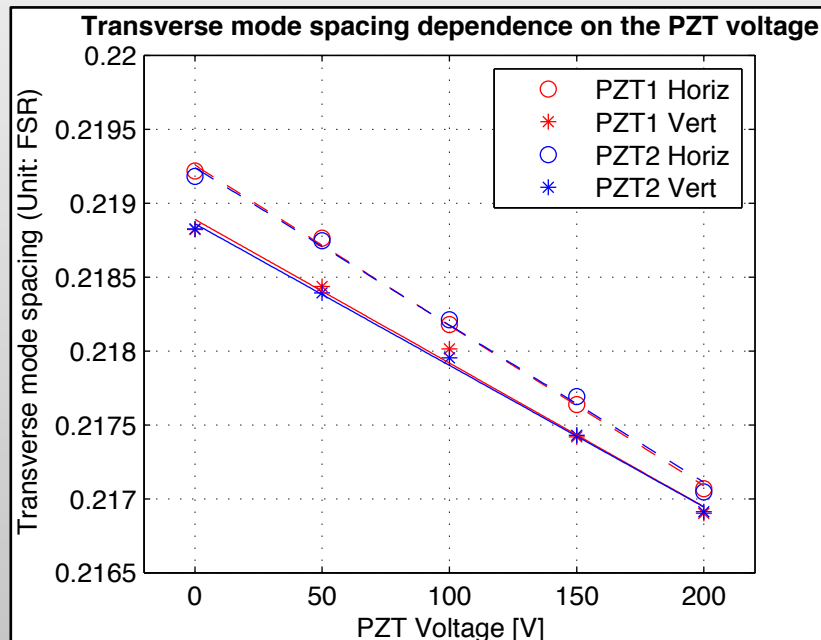
...BUT

TMS depend on the PZT voltage

- 3D deformation of the ring PZT?

Pitch TMS/FSR: $0.2189 - 9.7 \times 10^{-6} V_{\text{PZT1}} - 9.6 \times 10^{-6} V_{\text{PZT2}}$

Yaw TMS/FSR: $0.2192 - 10.8 \times 10^{-6} V_{\text{PZT1}} - 10.6 \times 10^{-6} V_{\text{PZT2}}$



The HOMs comes into the resonance at PZT voltage of $\sim 150\text{V}$.

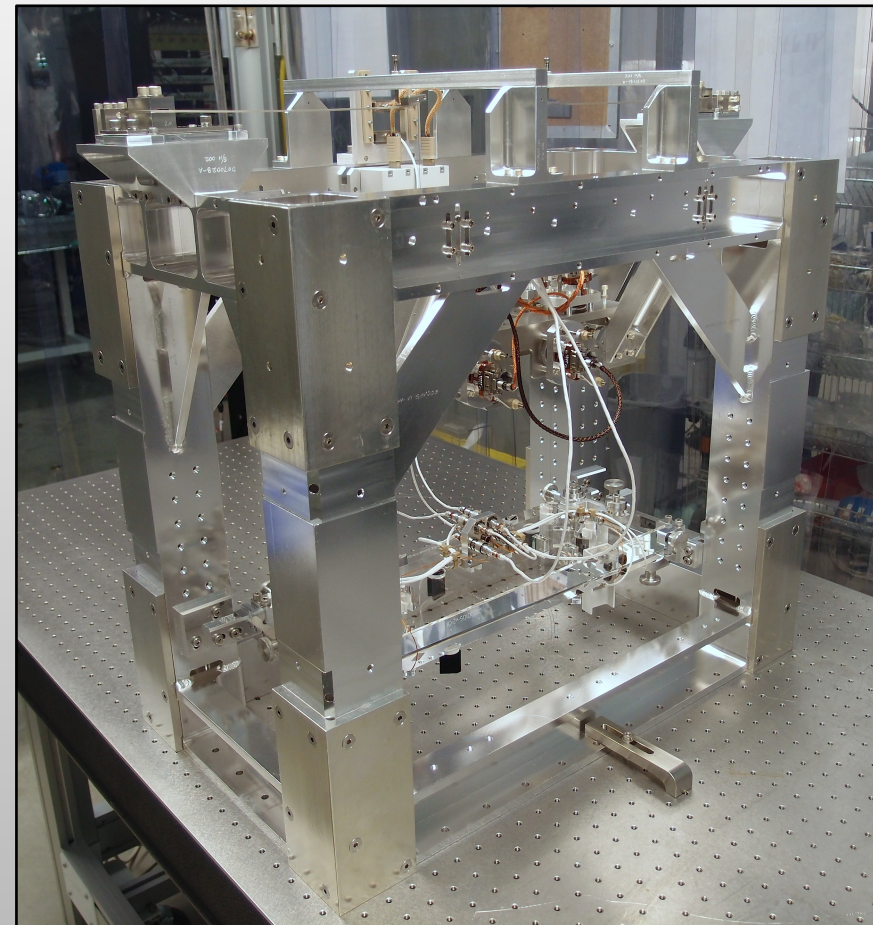
This actually does not happen as we limit the PZT voltage to 100V because of some other reason

Integration with the OMC suspension (OMCS)

- OMCS prepared with a metal bench
- Swapped the metal bench with the glass bench at LVEA
- Cabling / Weight balance
- Suspension tests in LVEA
transfer functions
damping control

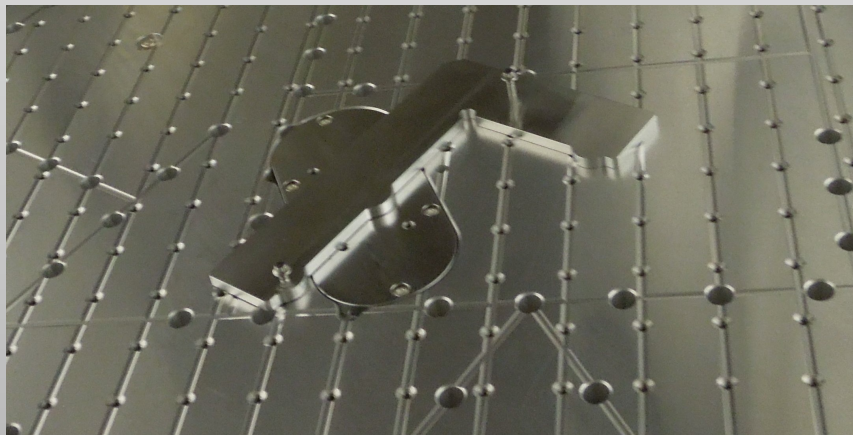
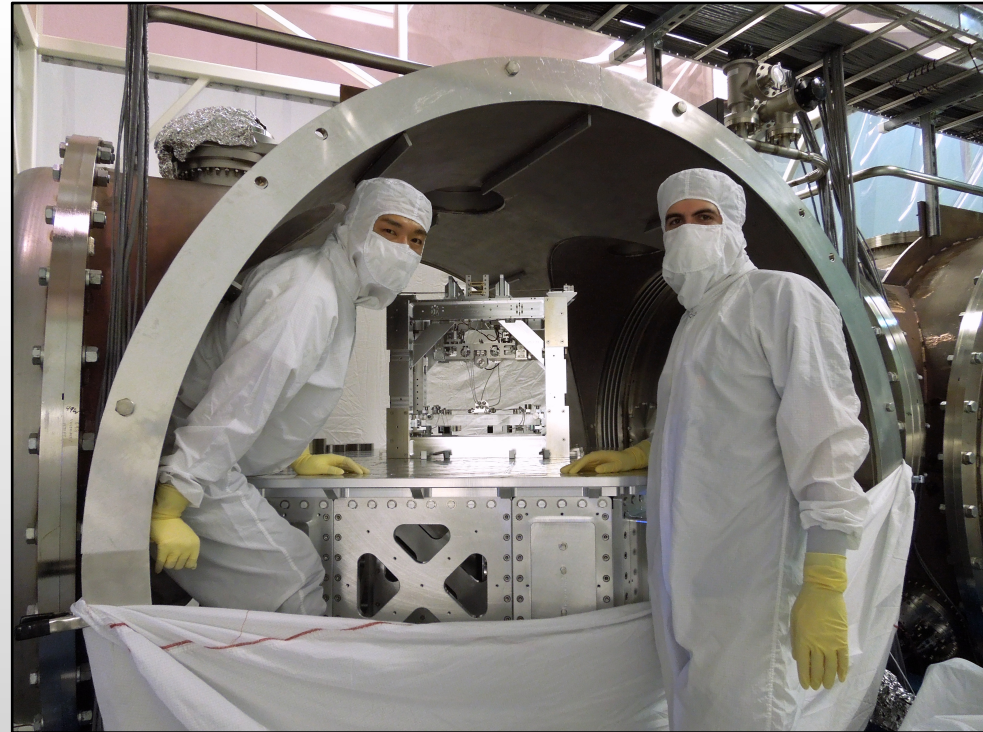
Placement on HAM6

- Loading on the ISI
a compact lift truck
to raise the OMCS
- "Cookie cutter"
- Suspension tests in HAM6



Installation @LLO

Koji Arai
KAGRA meeting, Oct. 9, 2013
JGW-G1301893 P25

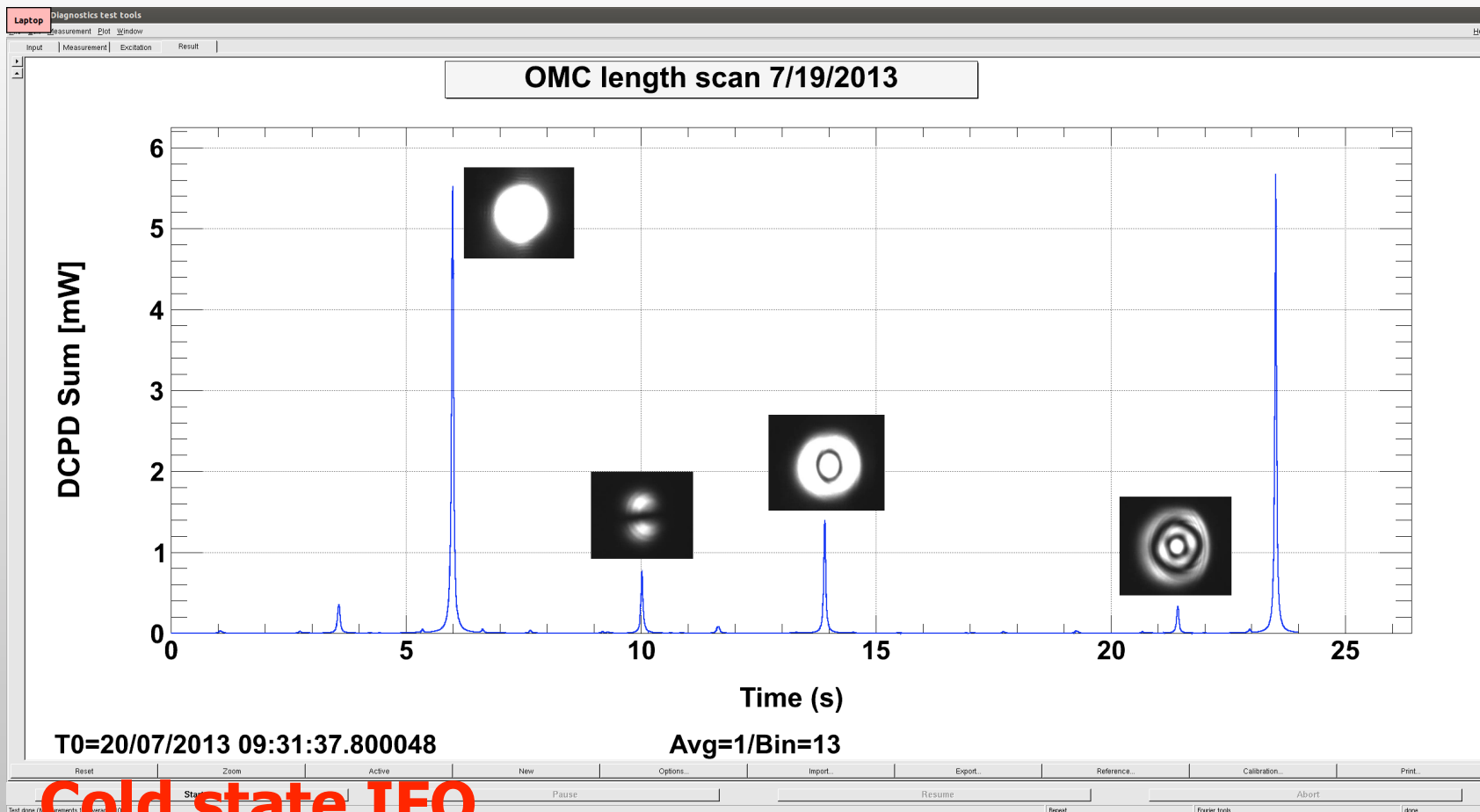


An installed scientist

Lift truck

Cookie cutter

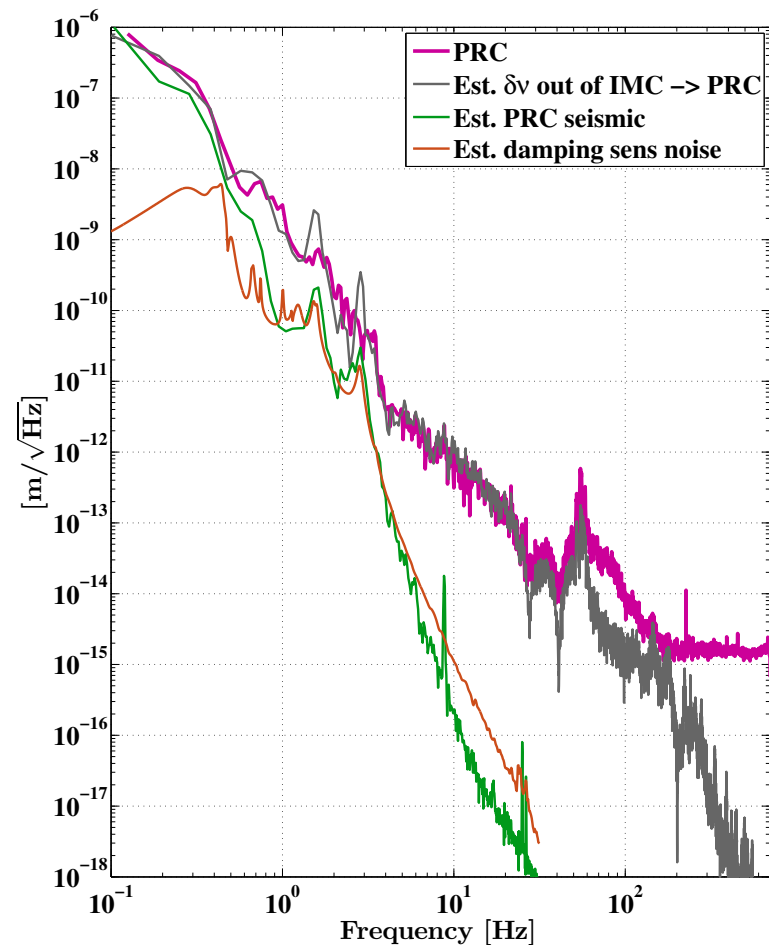
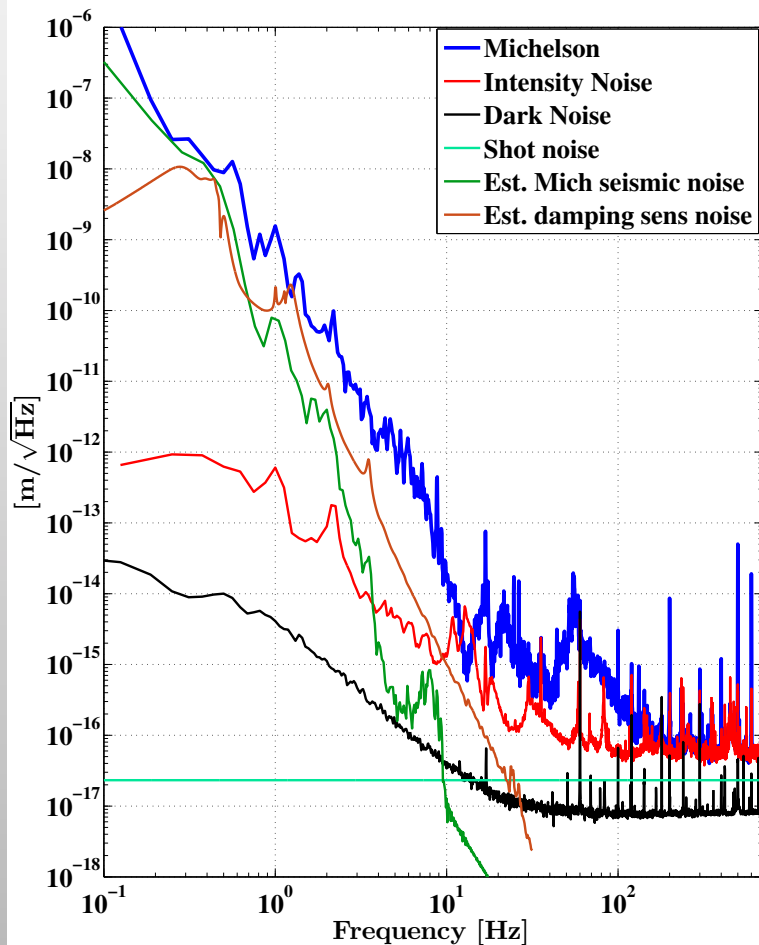
MODE MISMATCH?



- Cold state IFO
- Difference of the arms emphasized by the incorrect PR2/3 distance?

- By locking the PRMI with an offset we can use DC readout to improve the sensing noise
 5×10^{-17} m/rtHz above 100Hz, Intensity noise limited

Power Recycled Michelson locked on Carrier with DC Readout (~ 1 nm offset)



OMC Team

Koji Arai
KAGRA meeting, Oct. 9, 2013
JGW-G1301893 P28



**Jeff Lewis
(Engineer)**



Koji Arai (Scientist)



**Zach Korth
(PhD student)**

OMC Team

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Fabrication procedure
Mechanical design
Engineering issues



Jeff

Installation issues
Preparation@LLO



Koji

Optical design
Feedback control
Scientific requirements
Optical tests



Zach

Encounter with engineers

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- CAD

- Project control

- Sense of mechanics

easy-to-use / precisions / cost

e.g. transport fixture

- Specifications

PZT

Bonding

EP30 glass sphere bond-liner

UV cure

Bonding tests

Torquing

Thermal expansion coefficients

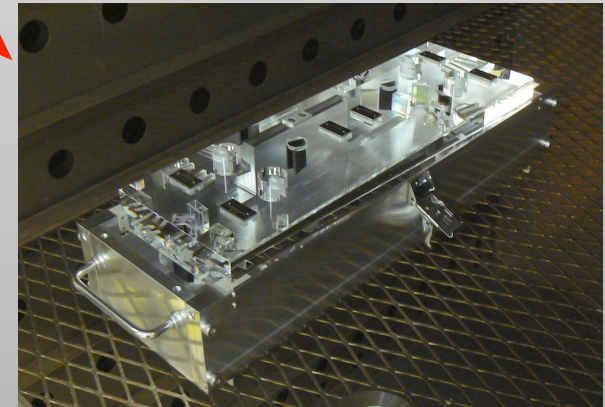
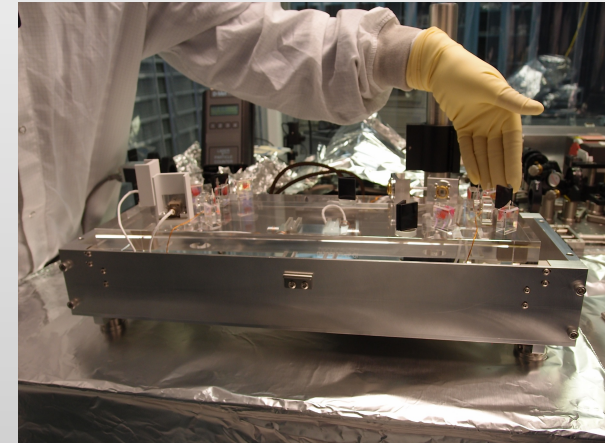
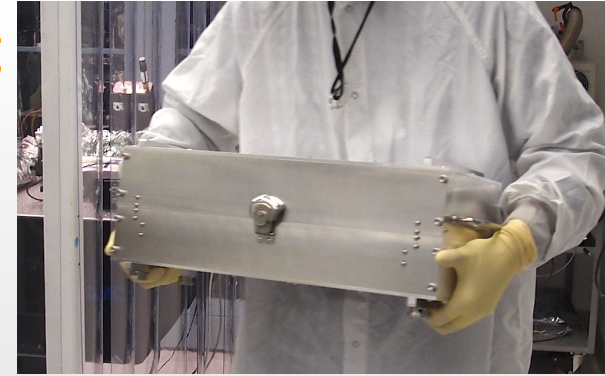
Thermal radiation

Contamination control

transport

testing

baking



Engineers are complimentary to scientists!