Recent News from KAGRA

Masaki Ando

(National Astronomical Observatory of Japan)

Outline



- News in project management
 - Schedule, budget, organization
 - International collaborations

- News in antenna construction
 - Tunnel, Vacuum, Cryo-system
 - Vibration Isolation

KAGRA



KAGRA (かぐら)

<u>Large-scale Cryogenic Gravitational-wave Telescope</u>

2nd generation GW detector in Japan



Large-scale Detector

Baseline length: 3km High-power Interferometer

Cryogenic interferometer

Mirror temperature: 20K

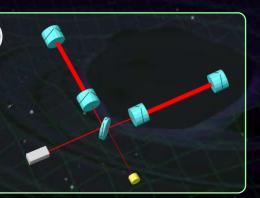
Underground site

Kamioka mine, 1000m underground

KAGRA Schedule

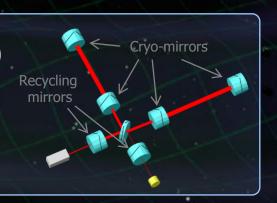


- •iKAGRA (2010.10 2015.12)
 - 3-km FPM interferometer
 - Baseline 3km room temp.
 - Operation of total system with simplified IFO and VIS.





- •**bKAGRA** (2016.1 2018.3)
 - Operation with full config.
 - Final IFO+VIS configuration
 - Cryogenic operation.



Budget Status

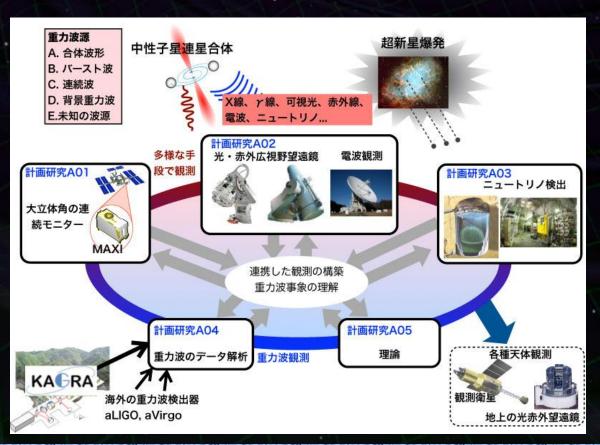


- •FY2012 is the last year of 'Leading-edge Research Infrastructure Program'.
 - → We are asking to extend the period.
- •Budget for bKAGRA has been submitted from MEXT to Ministry of Finance. (4 year budget starting in FY 2013)
- The Grants-in-Aid for Scientific Research
 - 'Scientific Research on Priority Areas' for the multimessenger astronomy. → Approved and started.
 - 'Specially Promoted Research' for bKAGRA construction (postdocs, engineers, travels and various studies.)
 Failed in this year → New proposal submitted (Nov 2012).

Multi-messenger Astronomy



- Based on the approved Grants-in-Aid for Scientific Research:
 'Scientific Research on Priority Areas' → 4.5-yr project.
- •GW theory, GW data analysis, EM transients, and Neutrino. (KAGRA and other GW experiment are not included.)



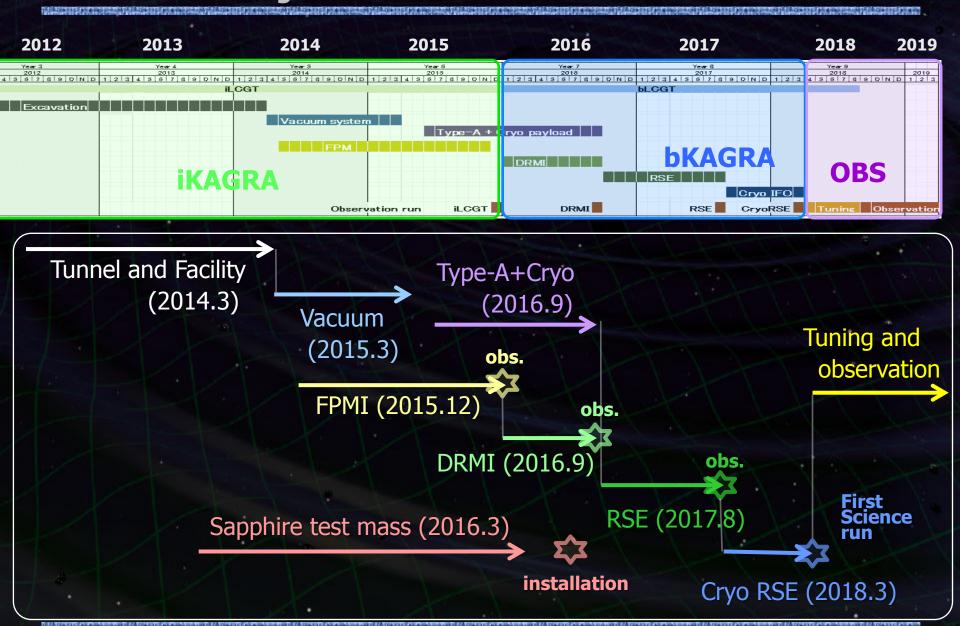
Schedule and Budget



FY2010 FY2011 FY2012 FY2013 FY2014 FY2015 FY2016 FY2017 'Leading-edge Research Infrastructure' 'Specially Promoted Research' program program (~98M\$) for iKAGRA (~5M\$) for detector upgrade **Budget** Budget from MEXT (~20M\$) **Budget from MEXT** for detector upgrade $(\sim 33M\$)$ for excavation **bKAGRA iKAGRA KAGRA** configuration Upgrade KAGRA facility Cryogenic mirrors •3km simple room- Full-power RSE configuration temp. interferometer Preparation of GW detection and **Purpose** infrastructure astronomy

Major Milestones of KAGRA





iKAGRA Scope



- Discussions on iKAGRA scope
 - External Review (April 17- 20, 2012)
 - * Recommend to consider about enhancing the scope.
 - * Better isolation, Silica fiber, Higher power.
 - Program Advisory Board (Aug 17 18, 2012)
 - * Recommend to consider about de-scoping iKAGRA.
 - * Schedule, financial, and man-power constraint.
 - MEXT recommendation: start obs. in 2017.



- We decided not to change iKAGRA scope.
- Observation start is first run with cryogenic interferometer.

Organization Updated



- Data analysis subgroup was divided into two subgroups
 → Data Analysis (DAS) and Data Management (DMG)
- Principal Program Advisory KAGRA Council Investigator Board T.Kajita R&D for International Board of Theory Future Representatives Affiliation for ICAR is System Project Management omitted Executive Office Member Support T.Naka-mura S.Kawamura Executive Office International Representatives N.Kanda (OCU) T Hohiyama M. Andu(NAOd) Y.Saito (KEK/UT) S. Kawamura System Engineering R.Takahashi K.Kuroda Office T.Suzuki (KEK/UT) External N.Mio(LIT,PSC) N.Kimura (KEK/UT) Review 1.Nakatani Chief, S.Kawamura K.Yamamoto Board M. Ohashi Dep. Chief M,Ando(NAOJ) Y.Aso (UT.Phys.) Y.Saito(KEK/UT S.Miyoki O.Miyakawa K.Somiya(TITEC) S.Telada (AIST) T.Akutsu (NAOJ) A.Araya (UT,ERI) Input/ Data Geophysics Vibration Main Inter-Auxiliary Analog Digital Date Tunnel Facility Vacuum Cryogenics Mirror Laser Output Manage-Interfero-Optios Electronics Analysis Isolation ferometer System (TUN) (FOL) (VAC) (CRY) (MIR) (LAS) Optios ment meter (MIF) (AUX) (AEL) (DAS) (DGS) (VIS) DMG (GIF) (100) Y. Saito N.Mio T.Uchiyam T. Suzaki N.Mio Y.Aso S.Telada T. Akutsu G.Miya N.Kanda H.Tagochi A.Arava S.Miyoki (KEK/UT R.Takahashi (UT.PSC O.Miyakawa (UT Phys.) (UT,PSC) (NAOJ) (KEK/UT) (AIST) (OCU) (OU) (UT,ERI) sub-chiefs. E.Hirose H.Hayama -A.Ueda (NAOJ

MoU with LSC and VIRGO



- Draft finished, waiting for sign
 - MoU General part.
 - Attachment A (LSC-KAGRA)
 Technical collaboration
 Information sharing
 - Attachment B (K-L-V)
 Data sharing.
 - Attachment C (VIRGO-KAGRA)
 Technical collaboration
 Information sharing

Memorandum of Understanding between

KAGRA, LIGO and Virgo Scientific Collaborations

A. Purpose of the agreement:

The purpose of this Memorandum of Understanding (MOU) is to establish a collaborative relationship between the signatories who are seeking to discover gravitational waves and pursue the new field of gravitational wave astronomy. The main scientific motivation is that the maximum return from gravitational wave observations is through simultaneous joint measurements by several instruments.

This MOU provides for joint work between the scientific collaborations of KAGRA, LIGO and Virgo. We enter into this agreement in order to lay the groundwork for decades of world-wife collaboration. When sensitive detectors are in operation, we intend to carry out the search for gravitational waves in a spirit of teamwork.

Details and extensions to this MoU will be provided in Attachments agreed by the parties.

•••

EliTES



- •ET-KAGRA collaboration program mainly on underground site and cryogenics.
- First general meeting on Oct. 3rd and 4th at Tokyo.





Construction Status

KAGRA Site



Underground site at Kamioka, Gifu prefecture

Facility of the Institute of Cosmic-Ray Research (ICRR), Univ. of Tokyo.

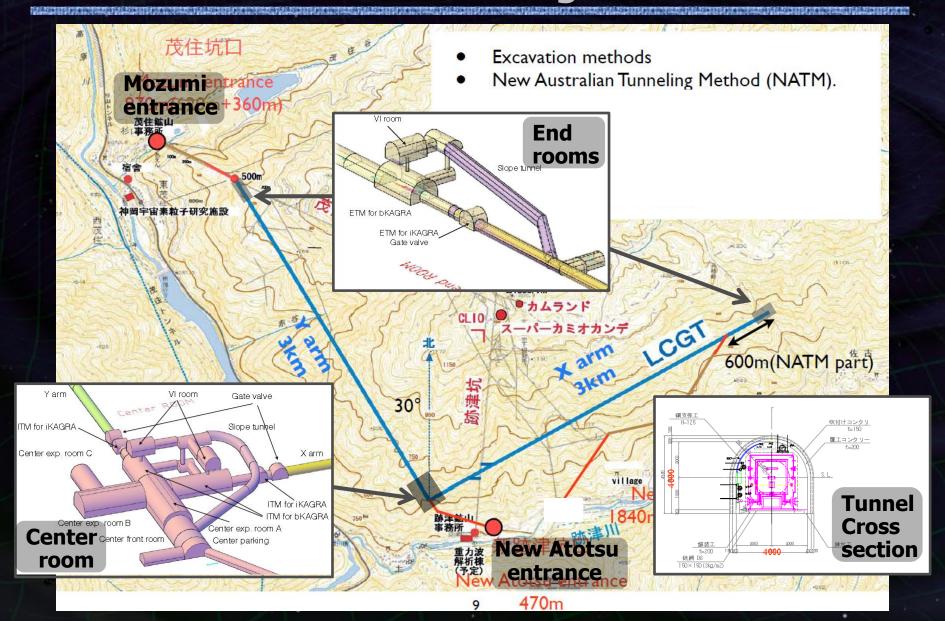


Neutrino
Super Kamiokande, Kamland
Dark matter
XMASS
Gravitational wave
CLIO, KAGRA
Geophysics
Strain meter

- 220km away from Tokyo
- 1000m underground from the top of the mountain. (Near Super Kamiokande)
- •360m altitude
- Hard rock of Hida gneiss(5 [km/sec] sound speed)

Tunnel Design





KAGRA Site



• Construction companies were decided, final design were almost fixed, construction was started.



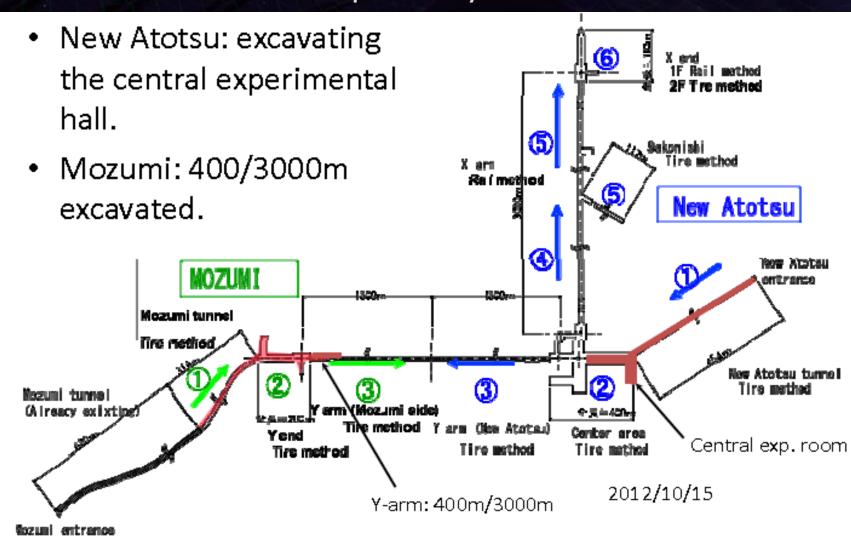


Field development for the 'New Atotsu' entrance, and for office building of the companies. (Photo by T.Uchiyama, April 2012)

Excavation Status



The tunnel will be completed by March 2014

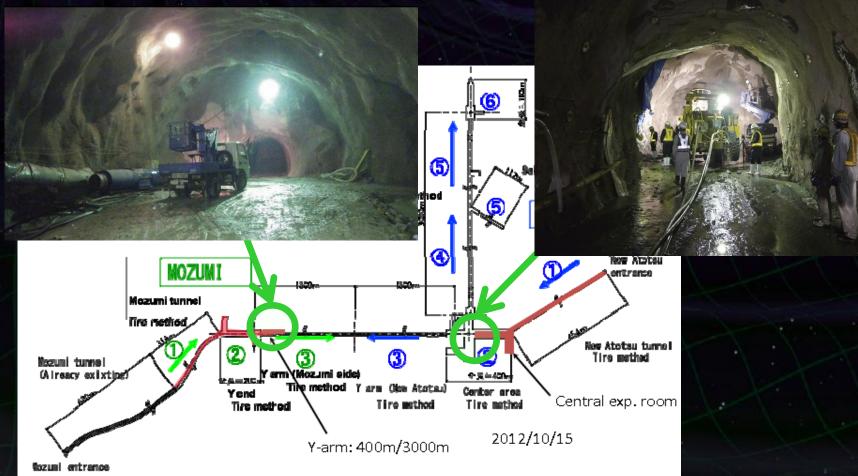


Excavation Status



- Mozumi entrance
 - Y-end room almost completed
 - Y-arm tunnel 400m/3000m

- New Atotsu entrance
 - Center Hall excavation



Surface Facility at Kamioka



Rent and remodel a public building (140m²) for free.

→ On-site office and laboratory for GW group.







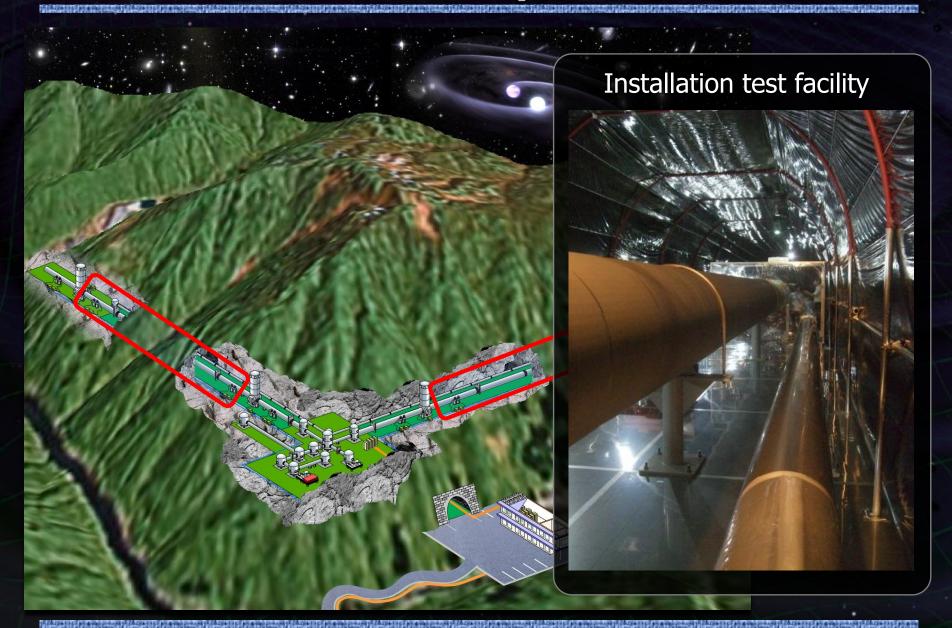
Aug. 29, 2012
Announcement for local people

→ Open as office in Nov.



Vacuum System





KAGRA Vacuum duct



- •12m, Φ800mm ducts for 3km x 2 arms.
 - → 70% of 478 ducts have been delivered.



Press to form a duct



Baking at MIRAPRO Co. Noda/MESCO, Kamioka



Bellows for each duct



Test at MIRAPRO Co. Noda

Presentation
By Y.Saito (KEK)

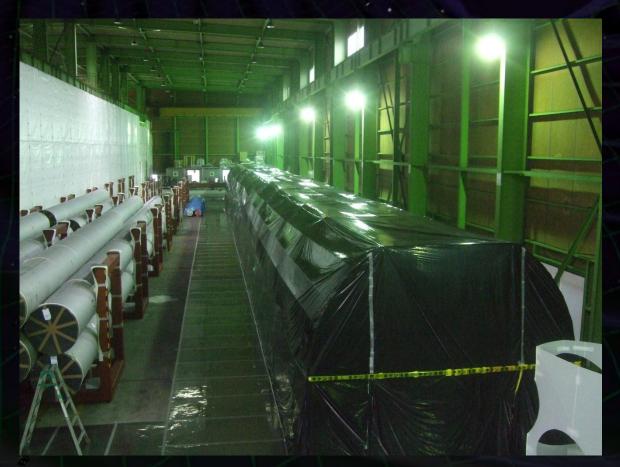


Transportation to Kamioka

Installation Test Facility



KAGRA tunnel simulator for installation test
(MIRAPRO, Noda factory)



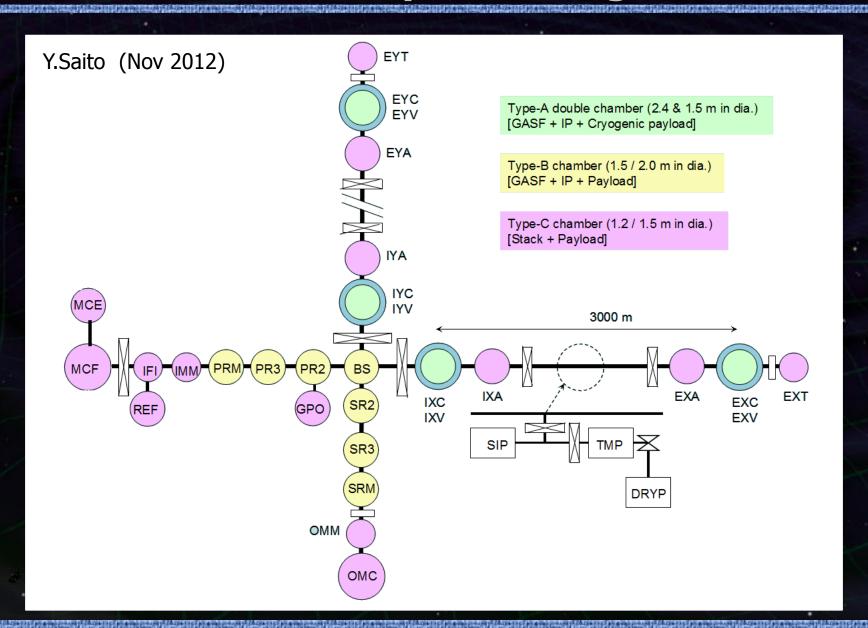






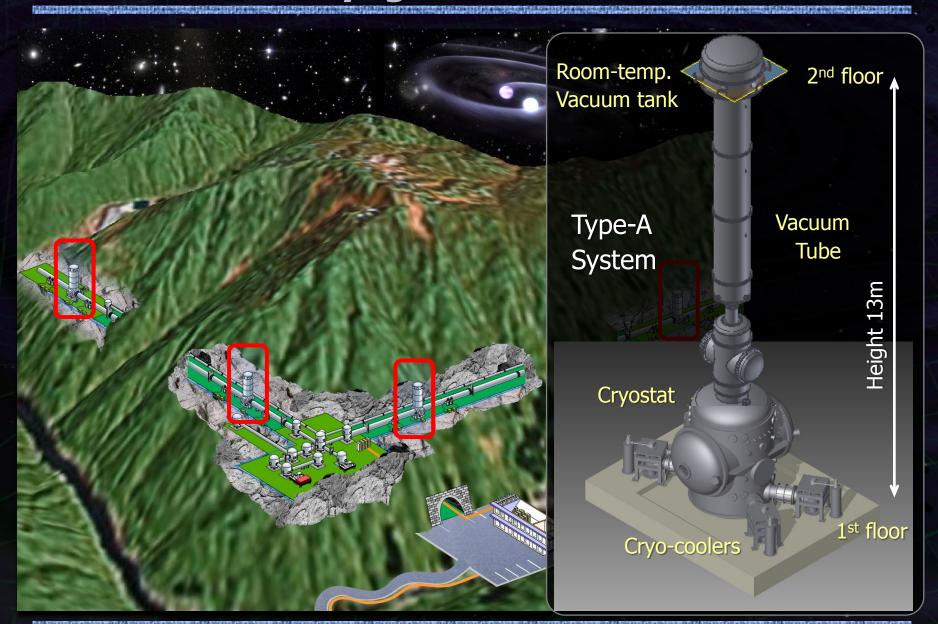
Vacuum System Design





Cryogenic Isolator



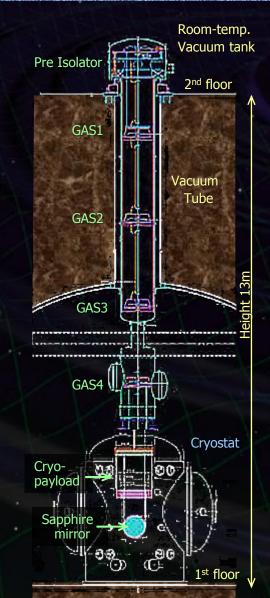


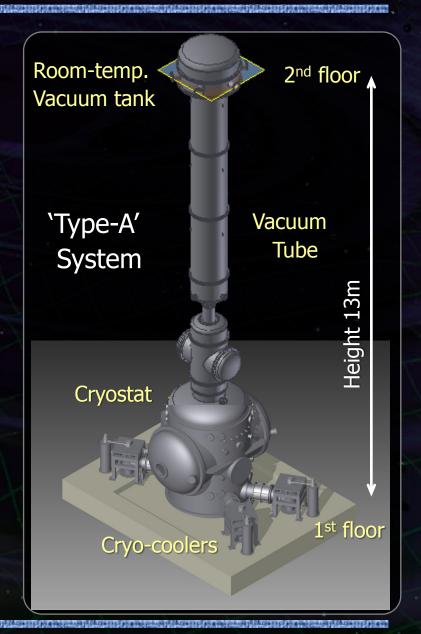
Cryogenic Mirror Isolator



Room-temp. Filter chain

Cryogenic Payload

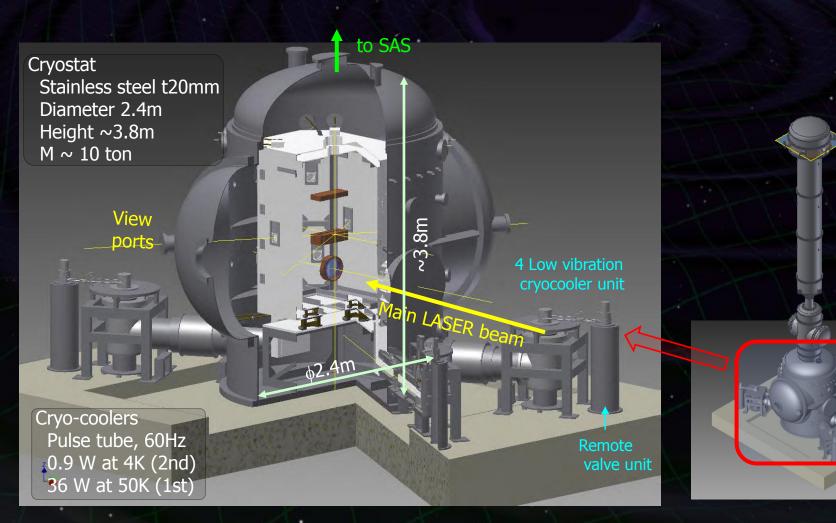




Cryostat Design



•4 Cryostats and 9 cryo-coolers in FY2012.



Cryostat Construction



Cryostat #1 in preparation for installation of radiation shield.

Cryostat #2 in leak test.

3rd and 4th cryostats under construction



Radiation shield



Toshiba Keihin Factory (Oct 31, 2012)

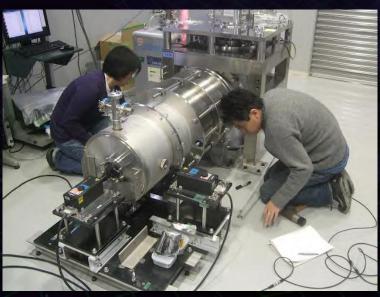
Cryo-cooler Construction



Cryo-cooler units at ICRR (Kashiwa)



Vibration measurement



Storage at ICRR (Akeno)

T.Suzuki at External Review (April 2012)





Cooling Test Plan



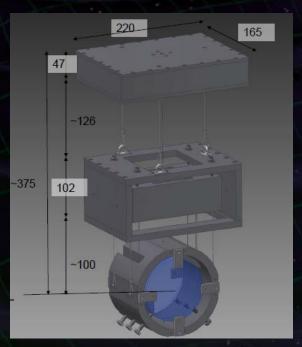
- Cryostat #1: 12 Nov. : Dummy sphere cooling
- Cryostat #2: 22 Nov. : DLC coated sphere
- •Cryostat #3: 7 Jan. : Vibration measurement

(Roma accelerometer, Interferometer)

•Cryostat #4: 21 Jan. -: 1/2-scale dummy payload.







Sapphire Mirror Isolator

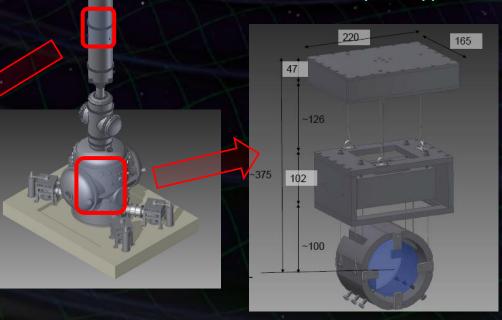






'Type-A' system

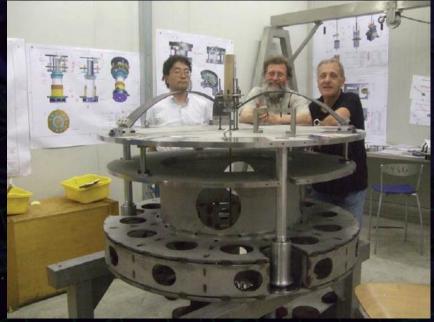
Cryogenic payload ½-scale prototype



Isolator Prototypes











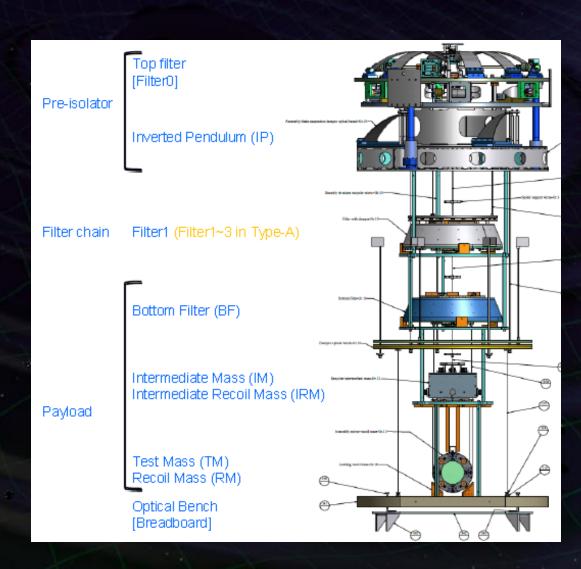
Vibration Isolation Design



'Type-B' SAS

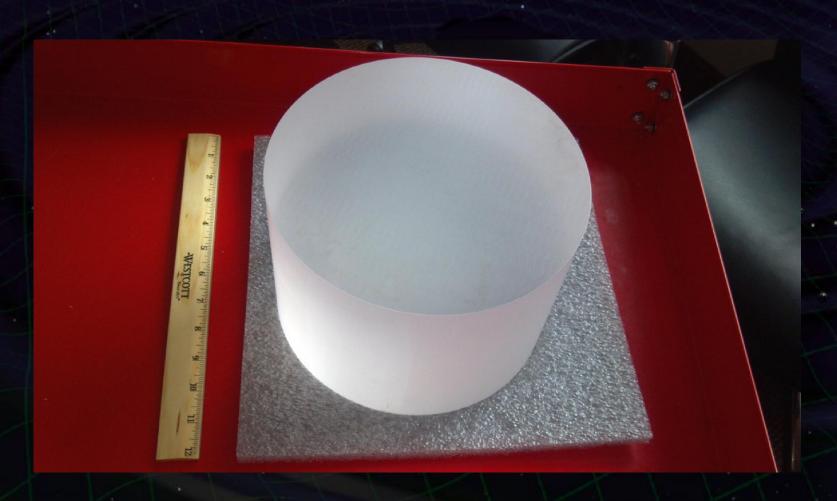
- Multi-stage, low-freq. isolator for silica room-temp. optics (BS, PRC, and SEC).
- •Full system test at TAMA300 facility from next April.

R. Takahashi at External Review (April 2012)



Sapphire Mirror



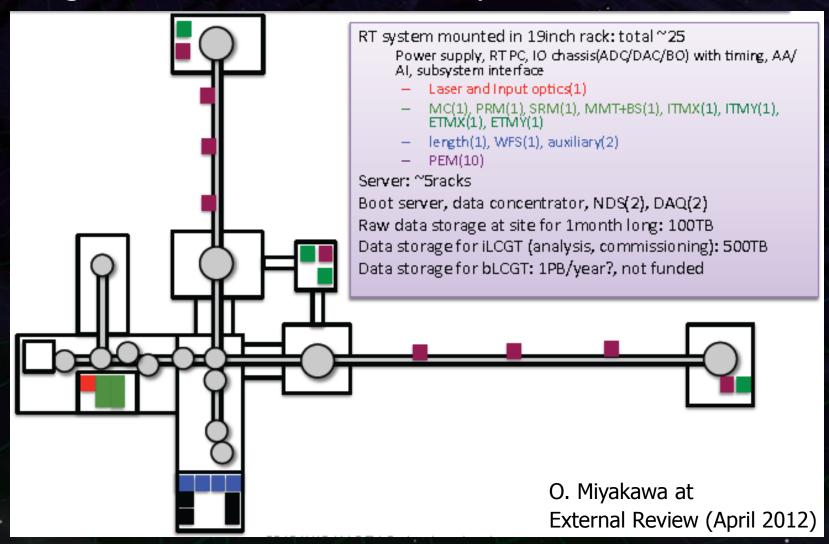


2 Sapphire substrates were delivered (Φ220mm, t 150mm, c-axis)

Digital Control System



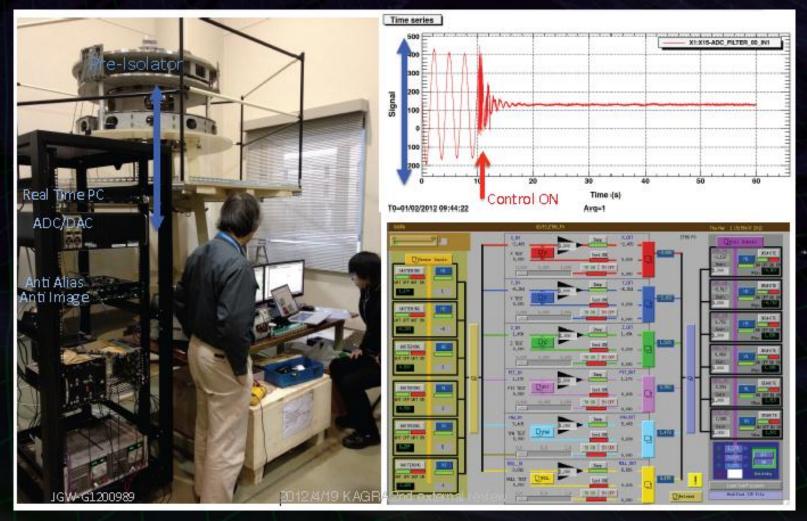
Digital control based on LIGO system.



Digital Control Test



• Damping of SAS pre-isolator (at ICRR Kashiwa).



O. Miyakawa at External Review (April 2012)



Summary

PAC33, The 33th Program Advisory Committee Meeting (November 8th 2012, Hanford, USA)

Summary



- A lot of progress in KAGRA!
 - Tunnel construction in progress.
 - Vacuum, Cryo system, etc. are becoming real.
 - Sorry for skipping important subsystems: FCL, MIF, IOO, AOS, MIR, LAS, DAS, AEL, GIF.
 - However, the budget, manpower and schedule are still very tight.
- We appreciate supports from LIGO.
 - Digital system and components
 - DCC access and many information
 - Human exchange
 - Reviews



End

PAC33, The 33th Program Advisory Committee Meeting (November 8th 2012, Hanford, USA)