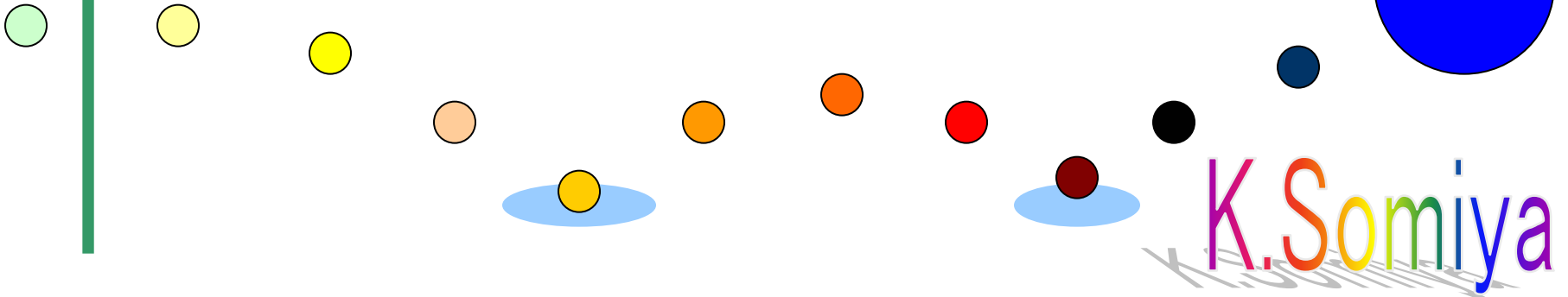


Requirements of LCGT parameters

LCGT F2F meeting @ ICRR
Feb. 2011

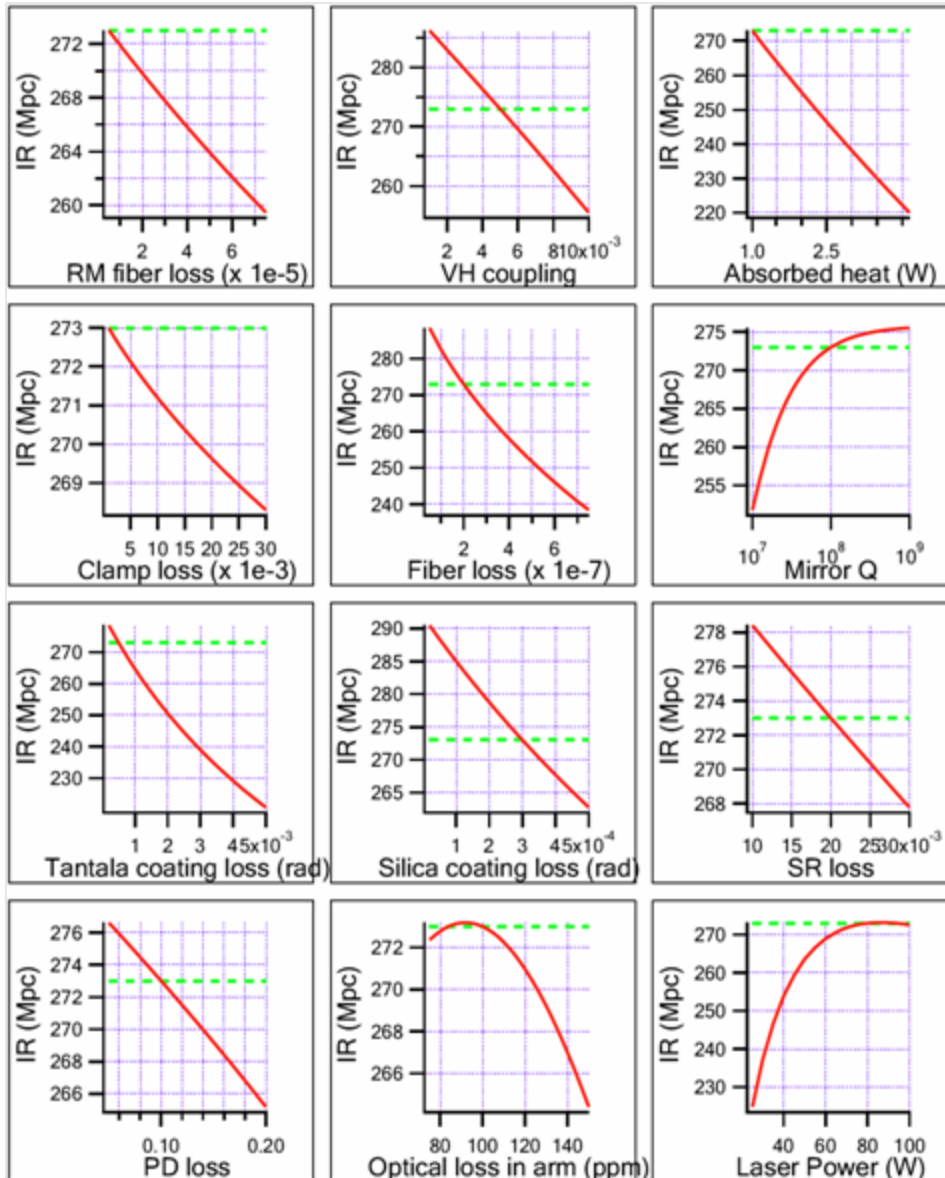
Waseda Inst for Advanced Study
Kentaro Somiya



Contents

- Fundamental noise parameters
 - LCGT design sensitivity is determined by QN, TN, seismic
 - Deterioration of the parameters degrades the sensitivity
- Other noise sources
 - Residual gas, laser noise, control-loop noise, etc.
 - Each noise should be 1/10 or less of the design sensitivity
- Interface parameters
 - Listing up all the parameters
 - Recognizing common parameters to subsystems
- Discussions

Requirements for QN/TN parameters



- IR reduction by degrade of some major parameters
- IFO setup is fixed
- Most of them changes the IR almost linearly
- Let's see the values that cause 1-2% IR reduction for each parameter

Optical loss and laser power

IR reduction	0%	1%	2%
Carrier power	81W	64W	57W
Arm loss	100ppm	124ppm	138ppm
SRC loss	2%	2.5%	3.0%
PD loss	10%	13.6%	17.1%

- IFO setup is fixed (detuning and DC readout phase)
- Arm loss causes (i) PRG reduction (ii) signal reduction
- SRC loss and PD loss cause signal reduction
- IR reduction due to the mode mismatching has been recently calculated by Nishida-san (TBA)

Mirror Qs and temperature

IR reduction	0%	1%	2%
Temperature (Absorbed heat)	20K (1W)	20.7K (1.16W)	21.4K (1.31W)
Silica coating loss	3e-4	3.5e-4	4.0e-4
Tantala coating loss	5e-4	6.6e-4	8.2e-4
Substrate Q	1e8	5.0e7	3.3e7

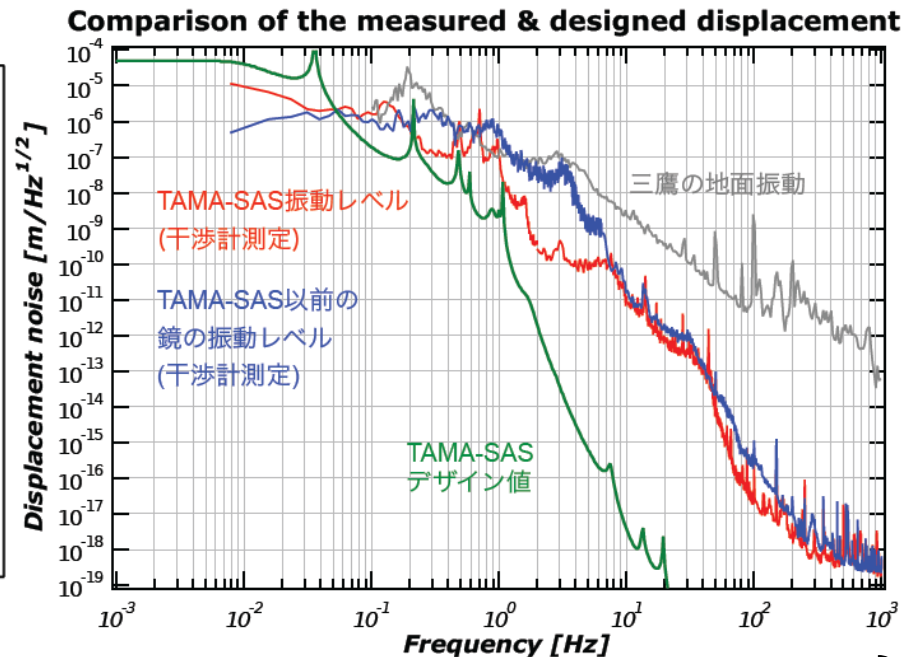
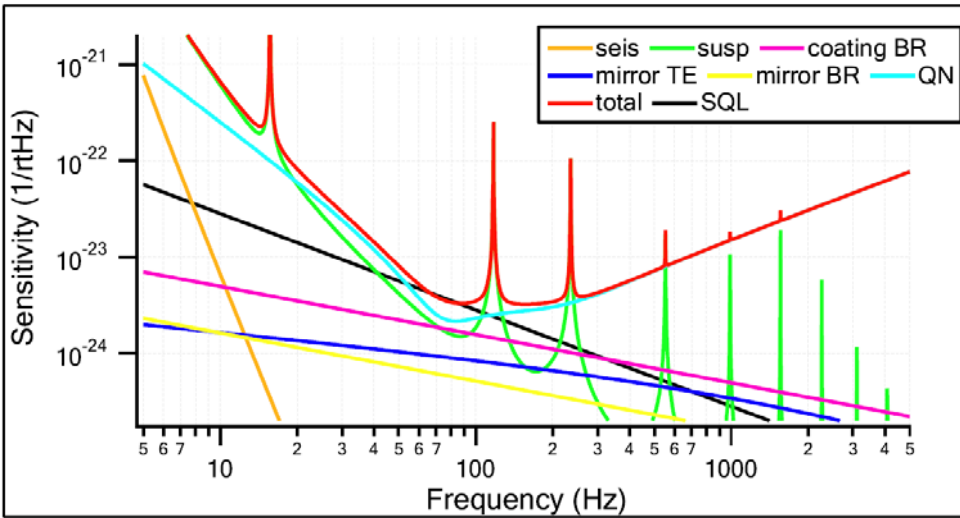
- Increasing heat absorption raises the temperature and increases thermal noise
- Measured coating losses are 5e-4/8e-4 but the values can be better with a proper annealing

Fiber losses and VH coupling

IR reduction	0%	1%	2%
Sapphire fiber (TM)	2e-7	2.32e-7	2.67e-7
Tungsten fiber (IM)	1e-4	N/A	
BeCu fiber (RM)	5e-6	1.8e-5	3.1e-5
Clamp loss	1e-3	1.6e-2	3.6e-2
VH coupling	1:200	1:172	1:152

- TM, RM, and VHC are severe
- IM TN is diluted with a mini-GAS btw IM and PF
- Mini-GAS resonance is set 0.4Hz (-> 3Hz?)
- The robe of vertical resonance may increase

Requirements for seismic isolation



left: LCGT design (seismic noise in orange)
 right: TAMA seismic noise with SAS (design in green, result in blue)

- 5-order degradation makes only 30% IR reduction
- In fact, it is easy for such a thing to happen
- Takahashi-san says the worst-case noise level can be calculated

Other noise sources

Excess noise should be 1/10 or less of the design sensitivity

- **Laser noise**

Transfer function has been calculated analytically, but it's probably better to use Optickle code for more detailed calculation (>> ISC task)

- **Loop noise**

Optickle code can calculate loop noise (>> ISC task). Number of issues remain.

- **Residual gas noise**

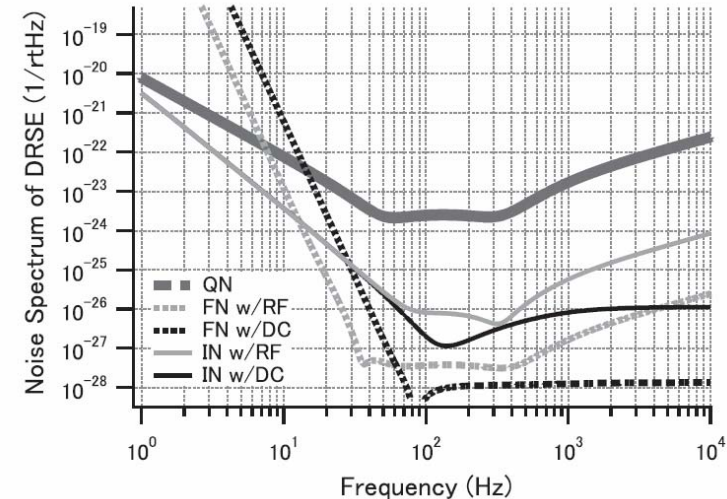
Requirement for the vacuum level is given:

$P < 5e-7$ Pa (with positive g-factor)

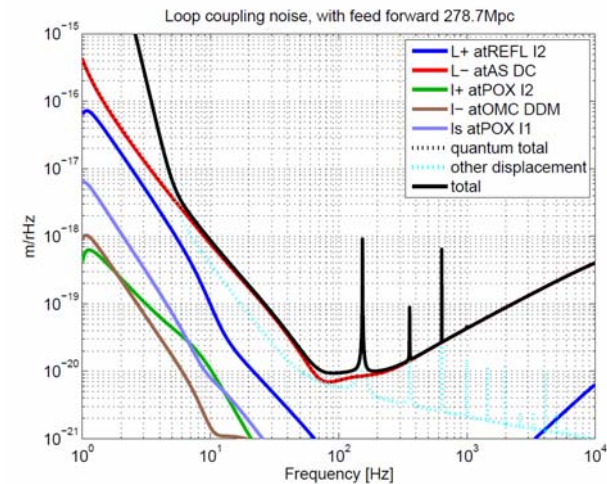
$P < 3e-7$ Pa (with negative g-factor)

- **Analog circuit noise**

- **Digital-system noise**



Laser noise TF for old aLIGO



Loop noise of LCGT

Cleanliness

- Facility support group is required to find a leader!
- Good mirrors can be easily contaminated on the way to its place
- LIGO has a lot of reports on this issue
- We will have only one spare mirror per arm
- Different requirements for different places:
 - CLASS 10,000 in the tunnel (?)
 - CLASS 1,000 in the lab
 - CLASS 100 in the assembly area (clean room)
 - No particles $>1\mu\text{m}$ in the vacuum chamber

(Those numbers are very preliminary)

Requirement report

LCGTのスコープと各サブシステムの要求値

LCGT Optical Configuration Group

January 23, 2011

1 はじめに

ここでは主に LCGT について考える。

2 Duty cycle 目標

年間を通じて 90 % の duty cycle を目標とする。

3 感度目標

中性子連星からの重力波を年に 2-3 回以上の頻度で観測できる感度の実現を目標とする。Duty cycle が 90 % で年 3 回の観測を実現するために必要な inspiral range (IR) は 221 Mpc である。この数字に安全係数 10 % を掛けた 244 Mpc を目指し、地面振動、熱雑音、量子雑音という 3 つの原理的な雑音を下げよう、干渉計の基本構造を決定する。これら以外の雑音は、それぞれが原則としてデザイン感度の 10 分の 1 以下になるようにする。

4 デザイン感度

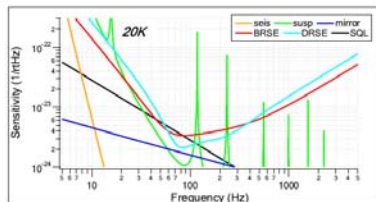


Figure 1: LCGT の感度。

- Requirements are to be set with (i) excess noise level (ii) feasibility taken into account
- Configuration group is preparing a report with Prof. Kawamura and Prof. Ando.
- We will ask subsystems for some helps and opinions
- The report will show requirements for each subsystem

Interface parameters

Value	TU	FA	VA	VI	MI	CR	IF	DG	EL	IO	LA	DA	GE
Digital													
ISC													
data analysis													
Optics													
Optics													
Optics													
Optics													
Cryogenics													
Electronics													
data analysis													
Laser noise													
seismic noise													
seismic noise													
ISC													
Laser noise													
Electronics													
Electronics													
ISC													
Test mass													
Digital													
Digital													
Digital													
Laser noise													
Digital													
Optics													
site information													
ISC													
ISC													
Test mass													
Cryogenics													
ISC													
Quantum noise													
ISC													
ISC													
ISC													
suspension TN													

- We've listed up so far 600+ parameters

- Related subsystems are marked

- List will be distributed to subsystems

⋮

* Values are not filled yet

End